

# Before asymptopia: High-dimensional inference in small sample size regimes in particle physics

Aishik Ghosh

STAMPS Workshop

14 May 2026



**Georgia Institute  
of Technology**



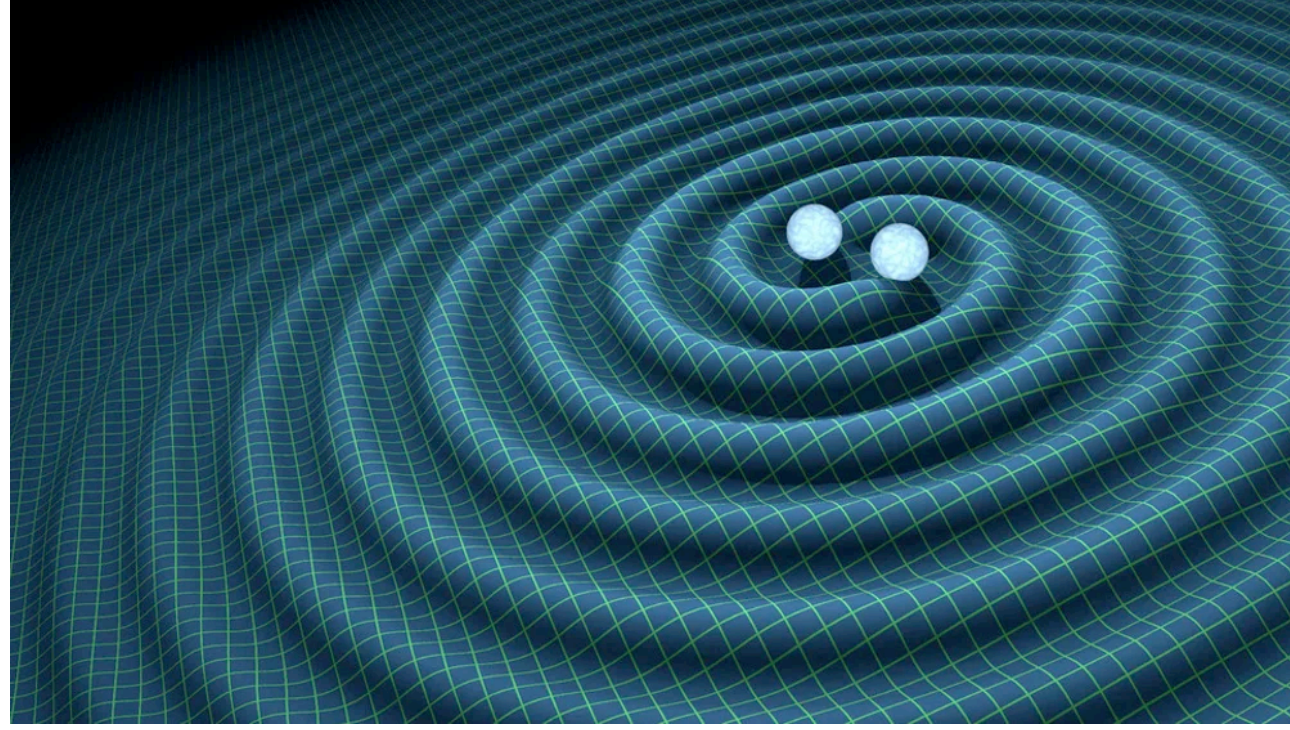
 [Prof-Aishik-Ghosh](#)

 [aishikghosh.bsky.social](#)

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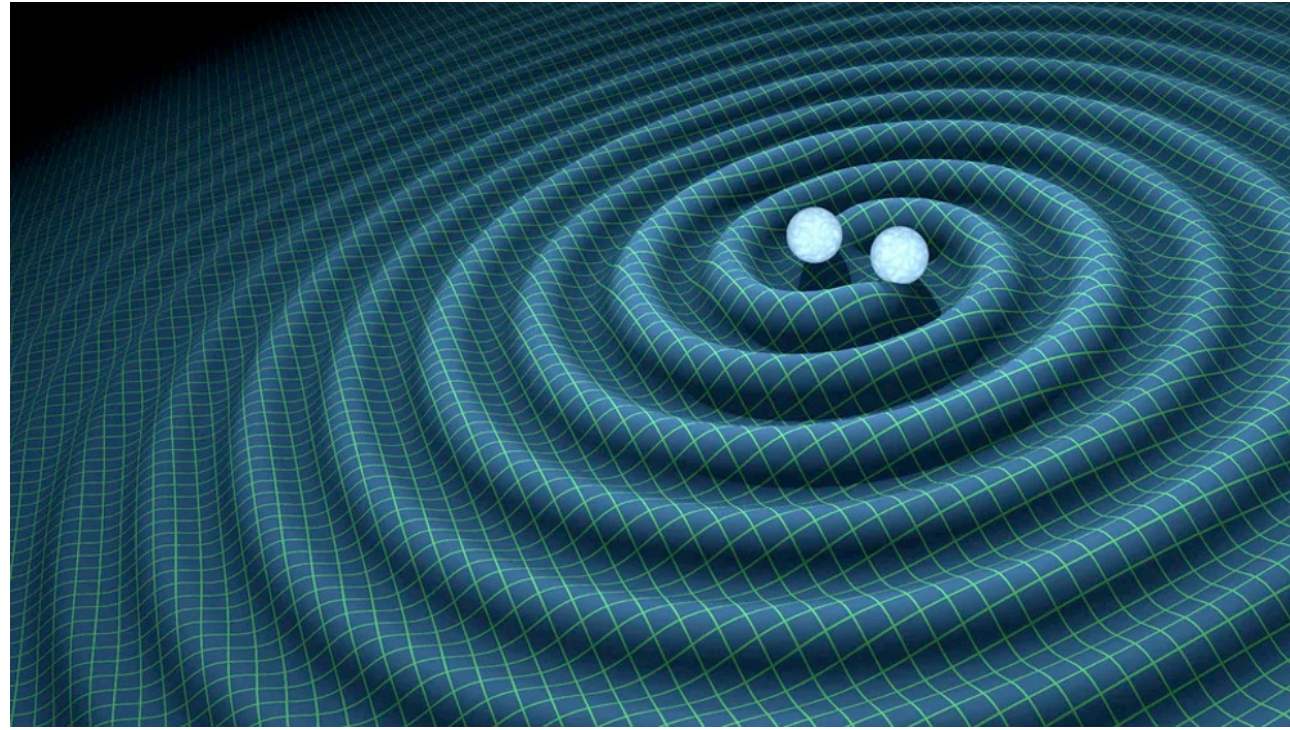
# Gravitational Wave discovery (2016)

Astrophysics:

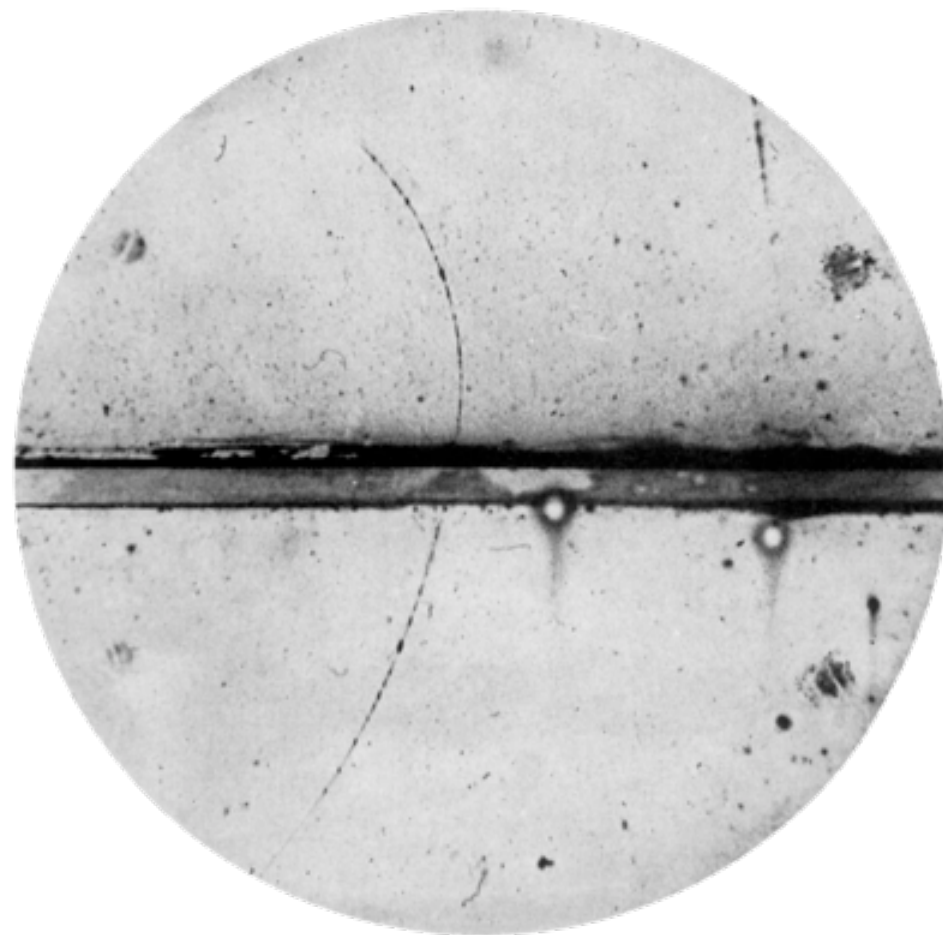


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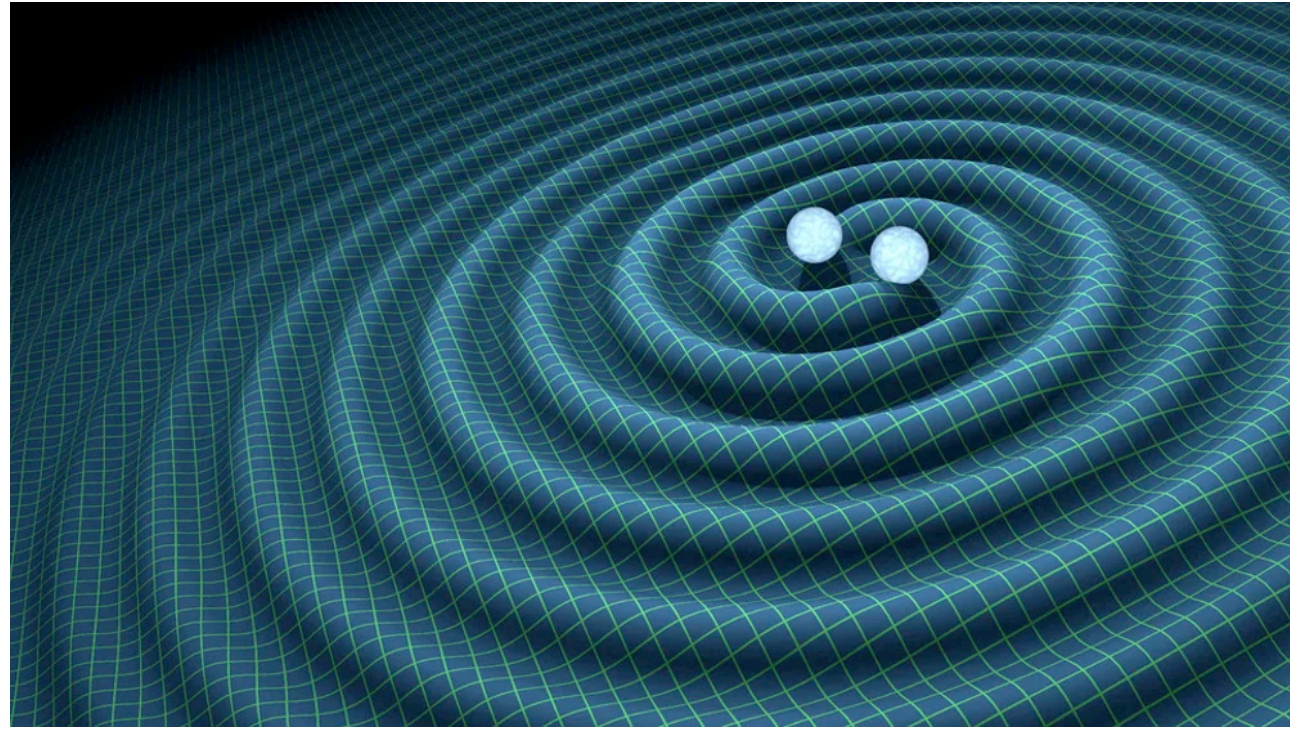
Particle physics:



Positron discovery (1930s)

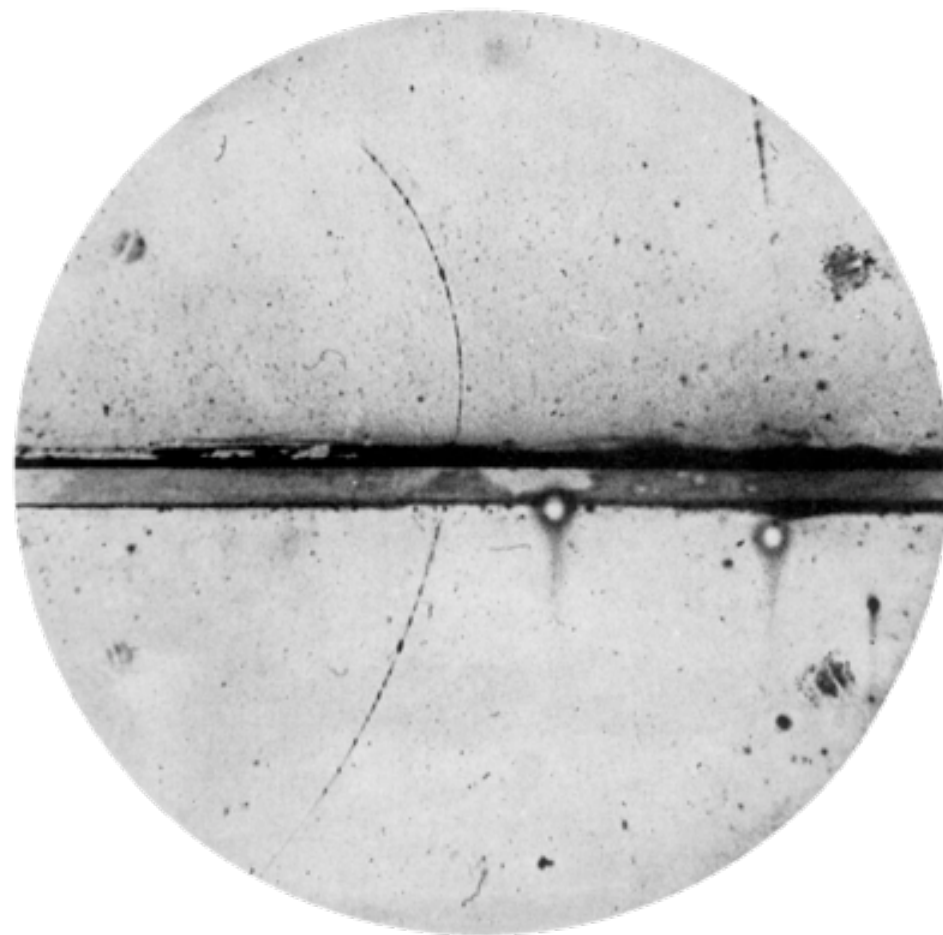
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Astrophysics:



Single event

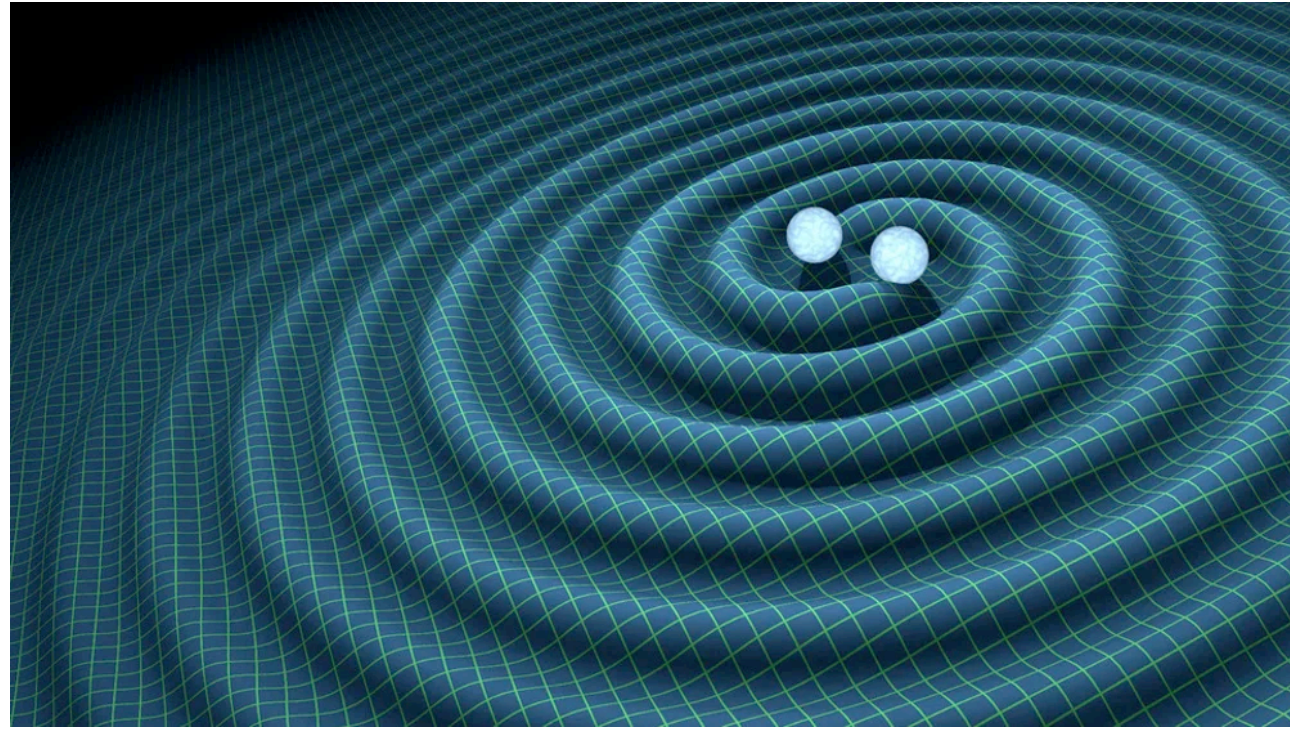
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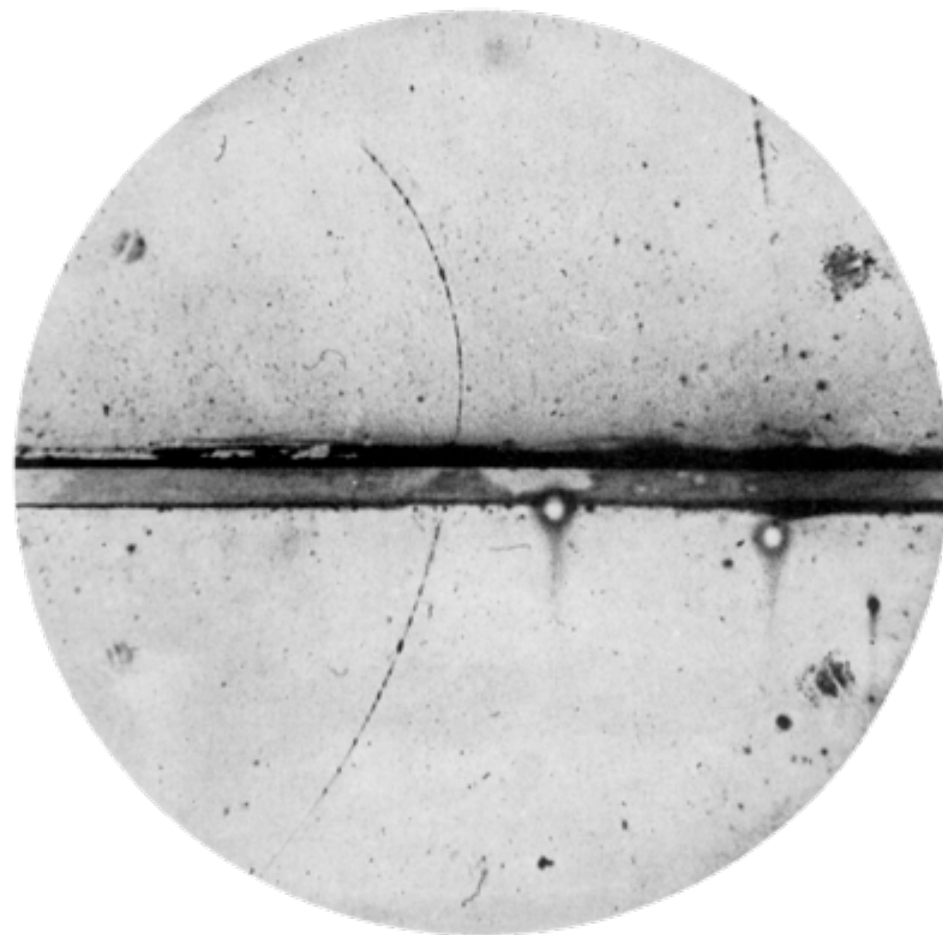
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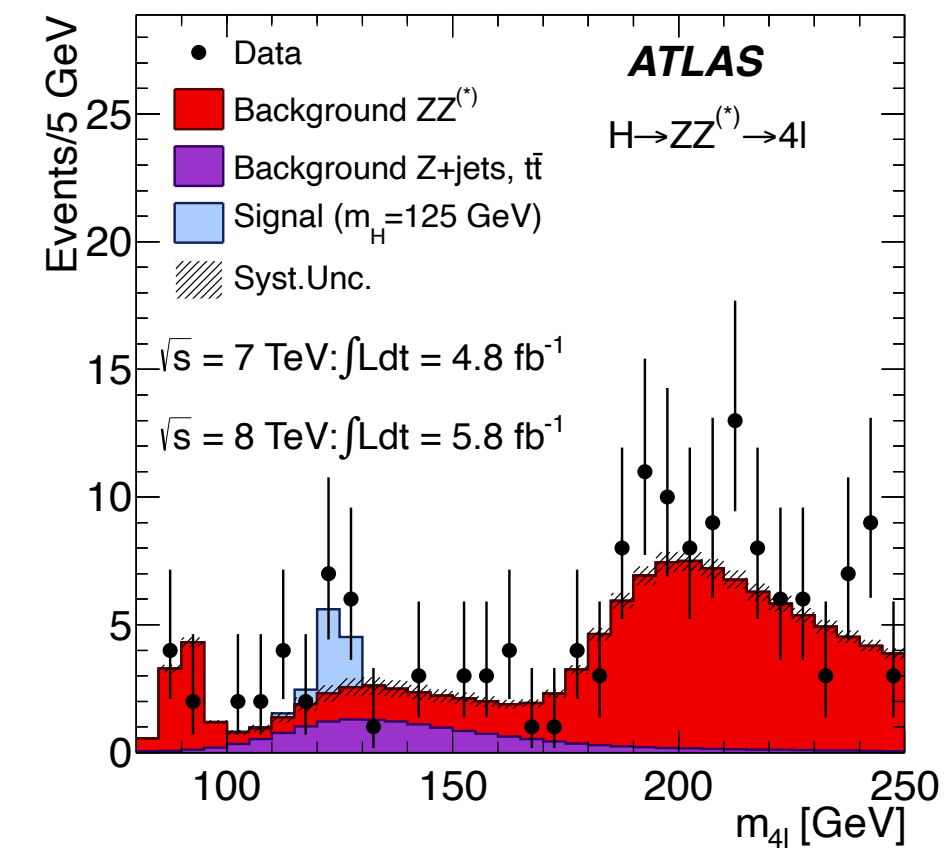
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Collection of events

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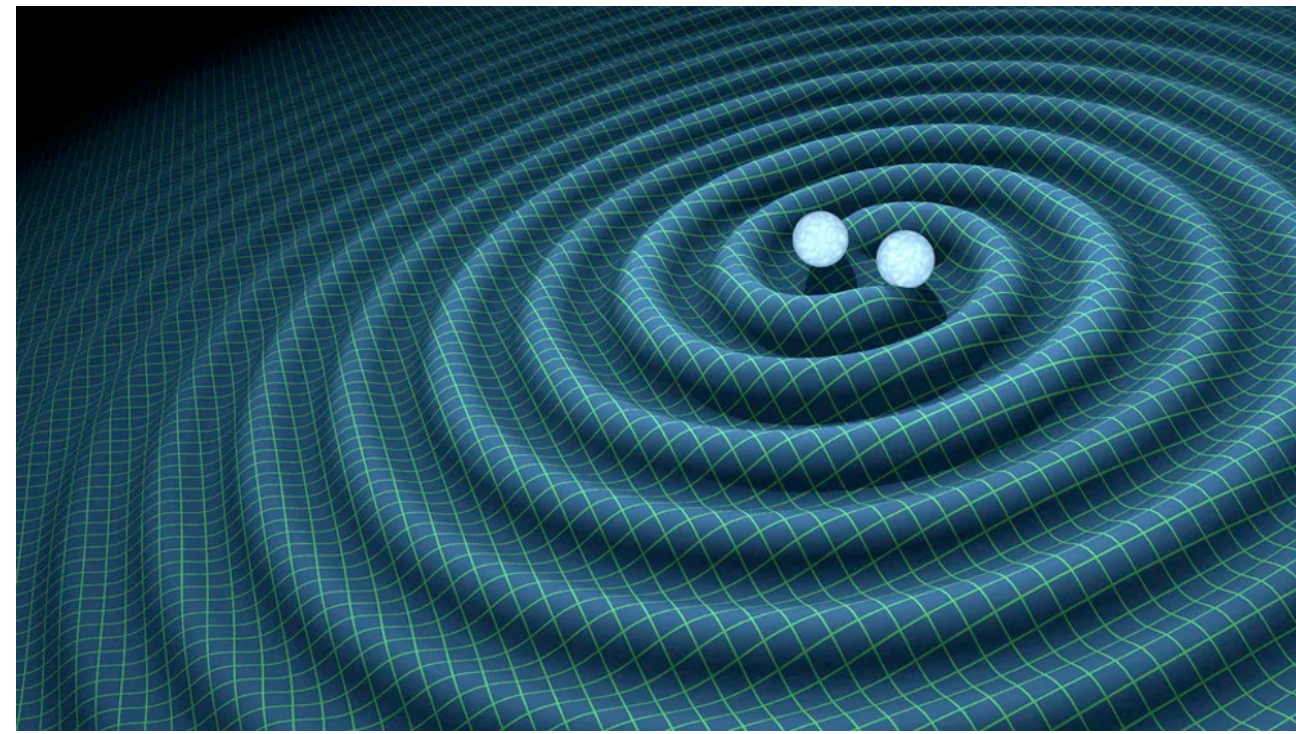


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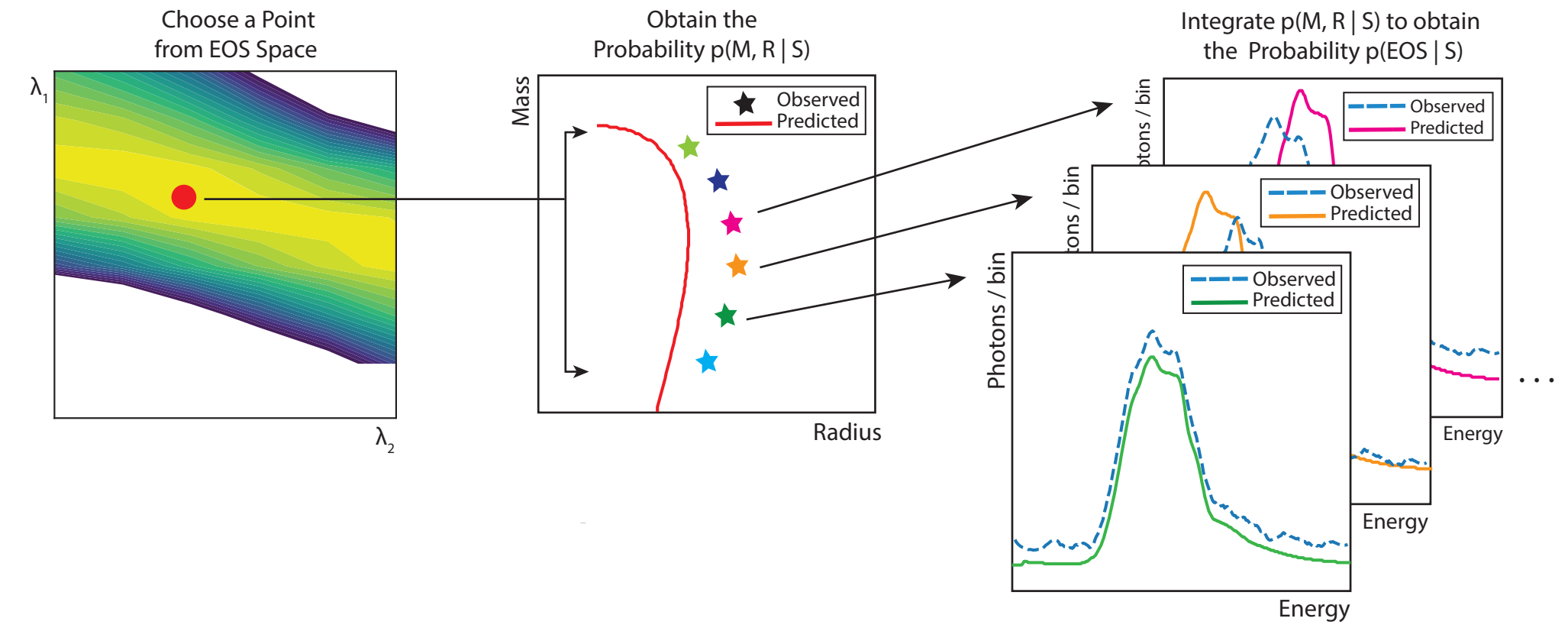
Higgs boson discovery (2012)

# Gravitational Wave discovery (2016)



Astrophysics:

# Probing the interior of neutron stars

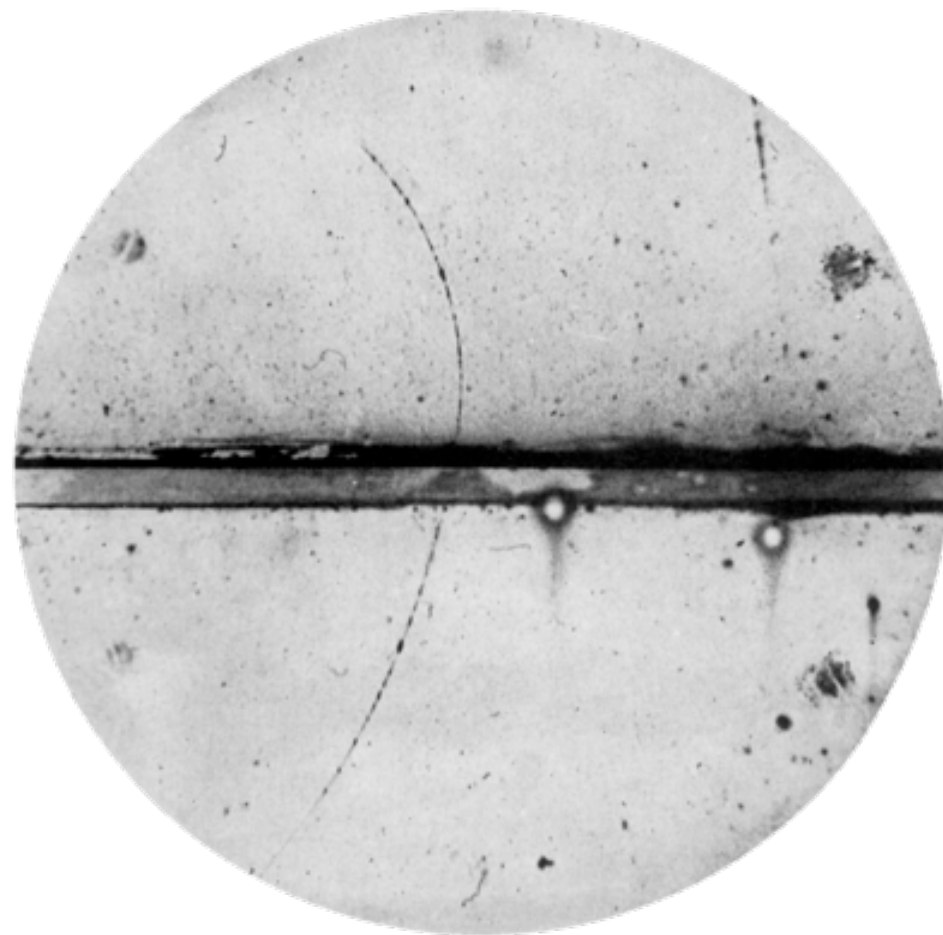


[JCAP12\(2023\)022](#): Farrell et al. (incl. Ghosh)

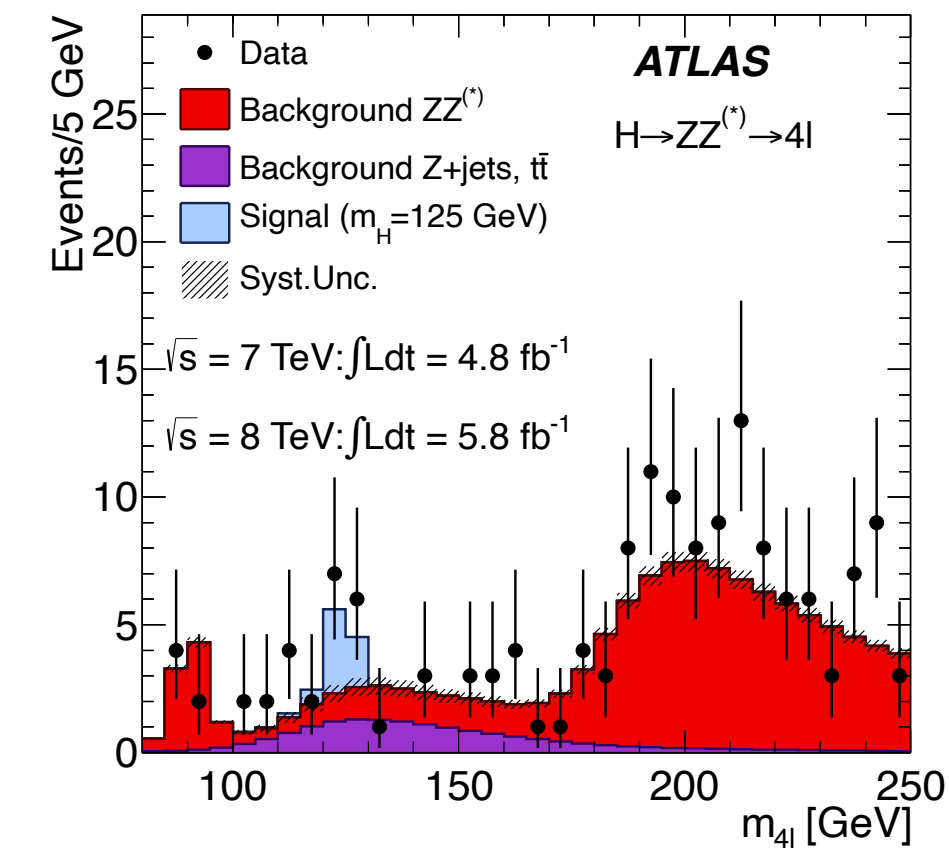
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## ML for Detector Simulation

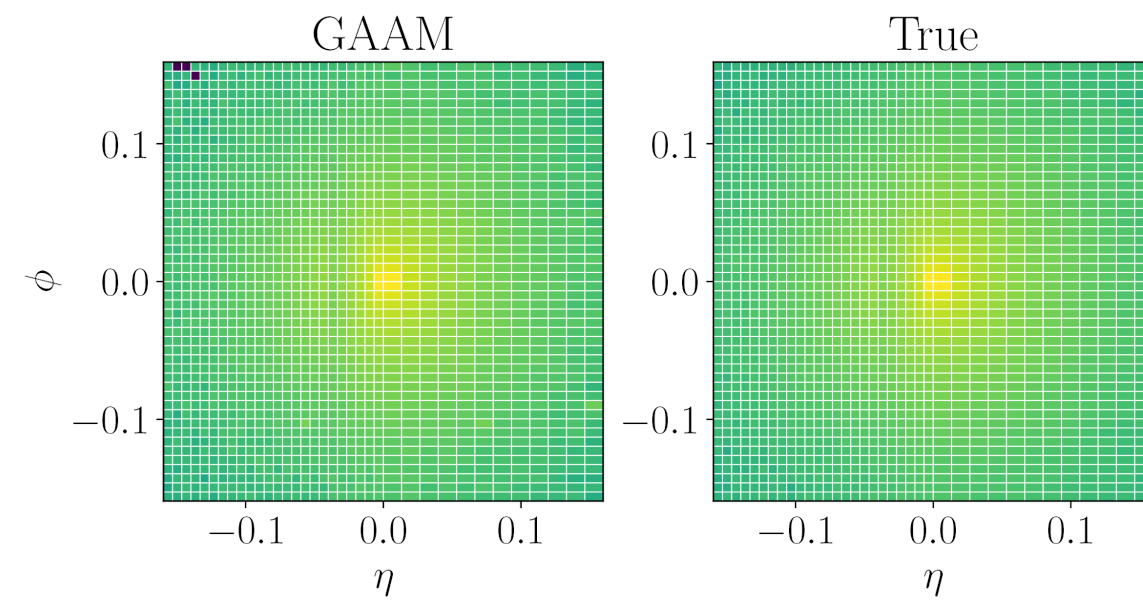


Image: [arXiv:2305.11531](https://arxiv.org/abs/2305.11531)

## ML for Reconstruction

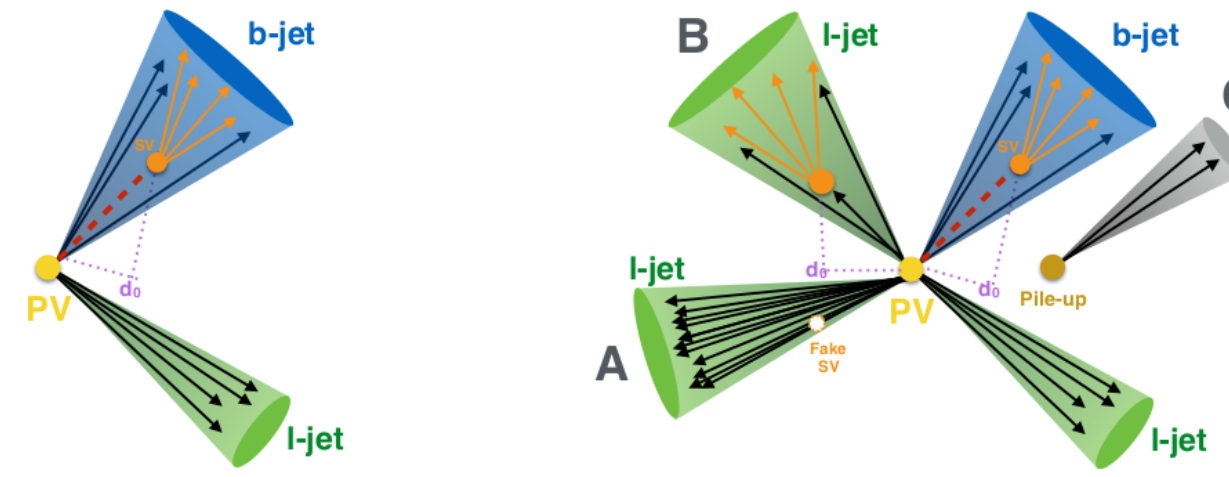


Image: [USiegen group](https://www.usiegen.de/)

## ML for Signal vs Background

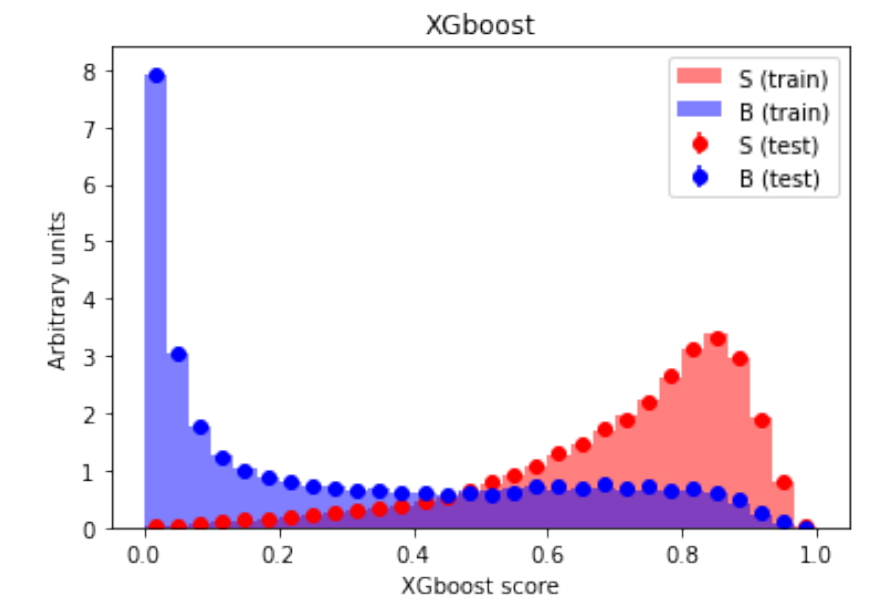


Image: [ML4FP School](https://ml4fp.github.io/)

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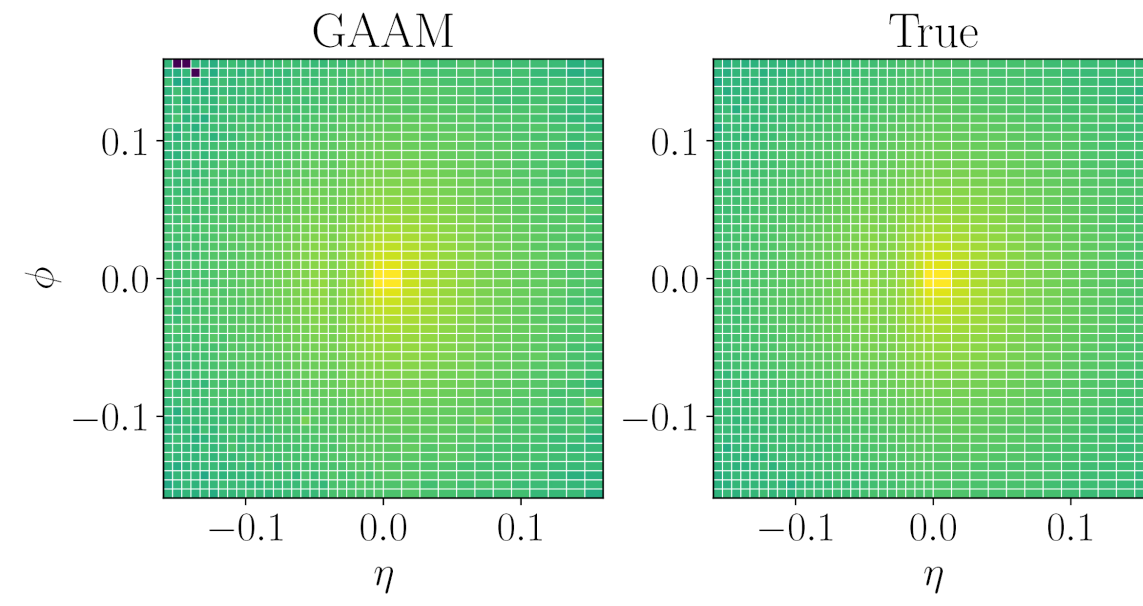


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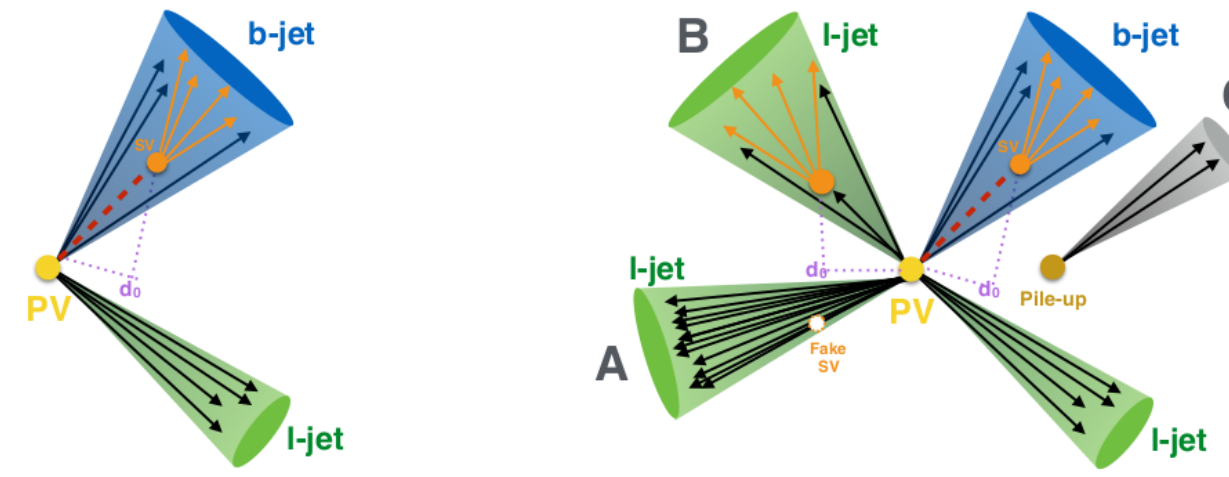


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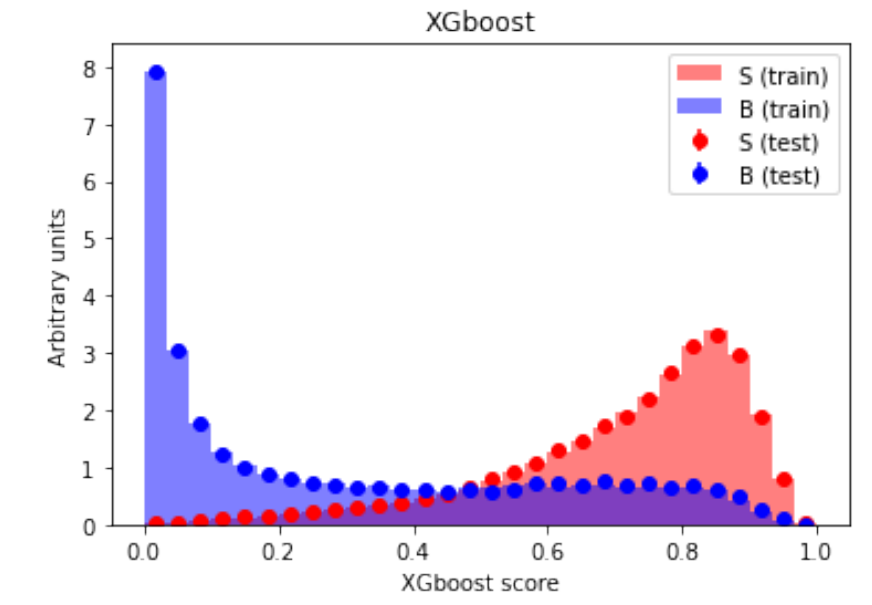


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The final statistical inference step has remained the same

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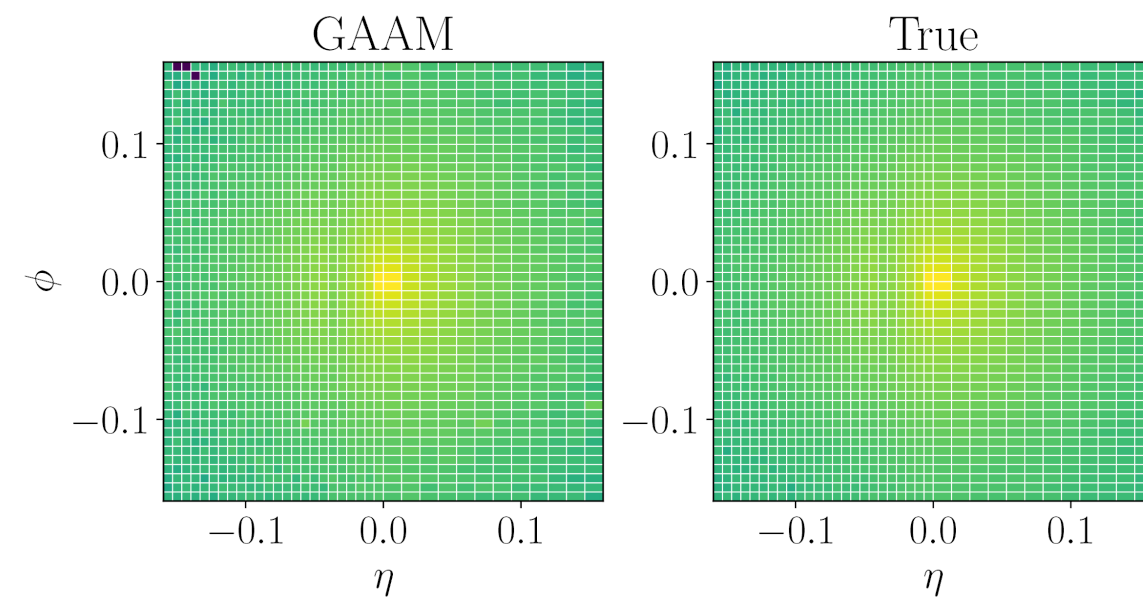


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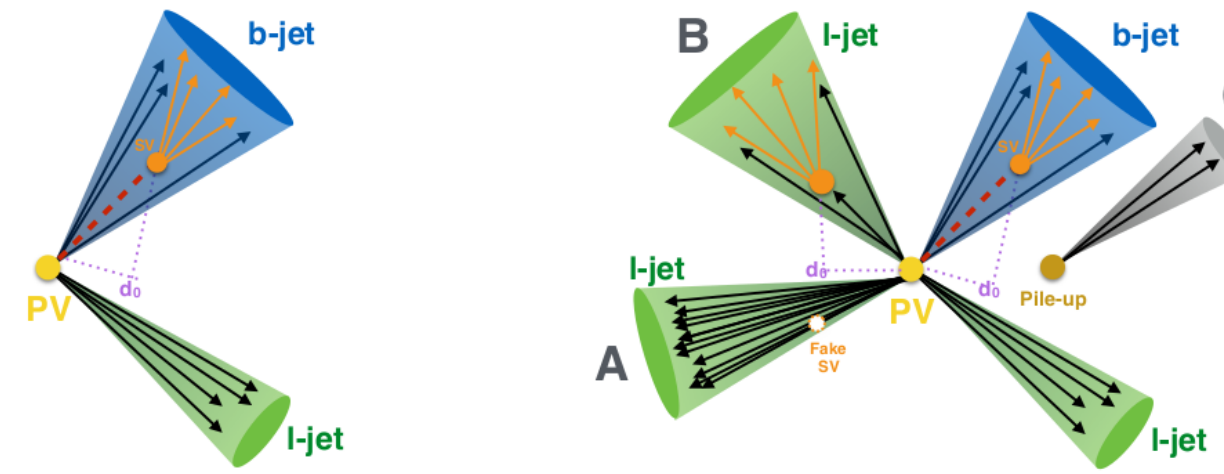


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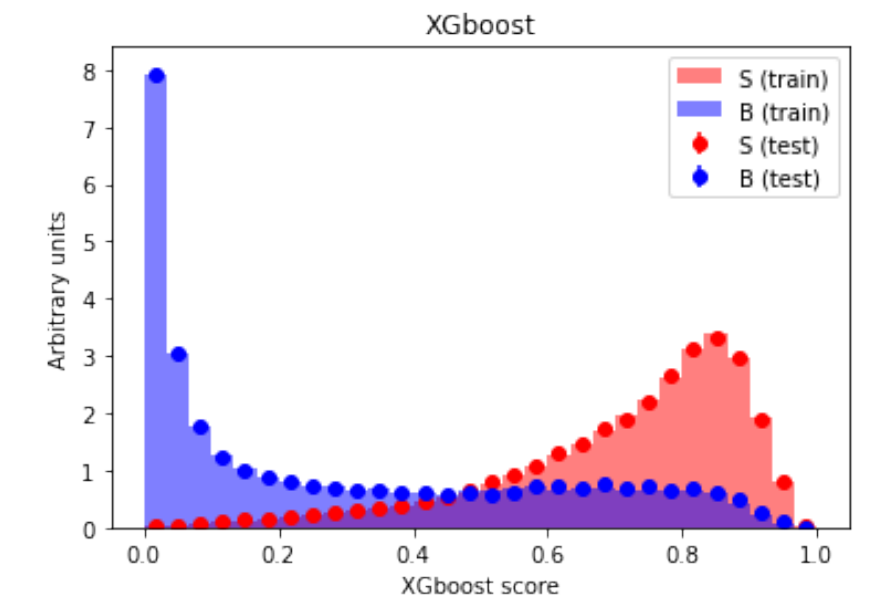
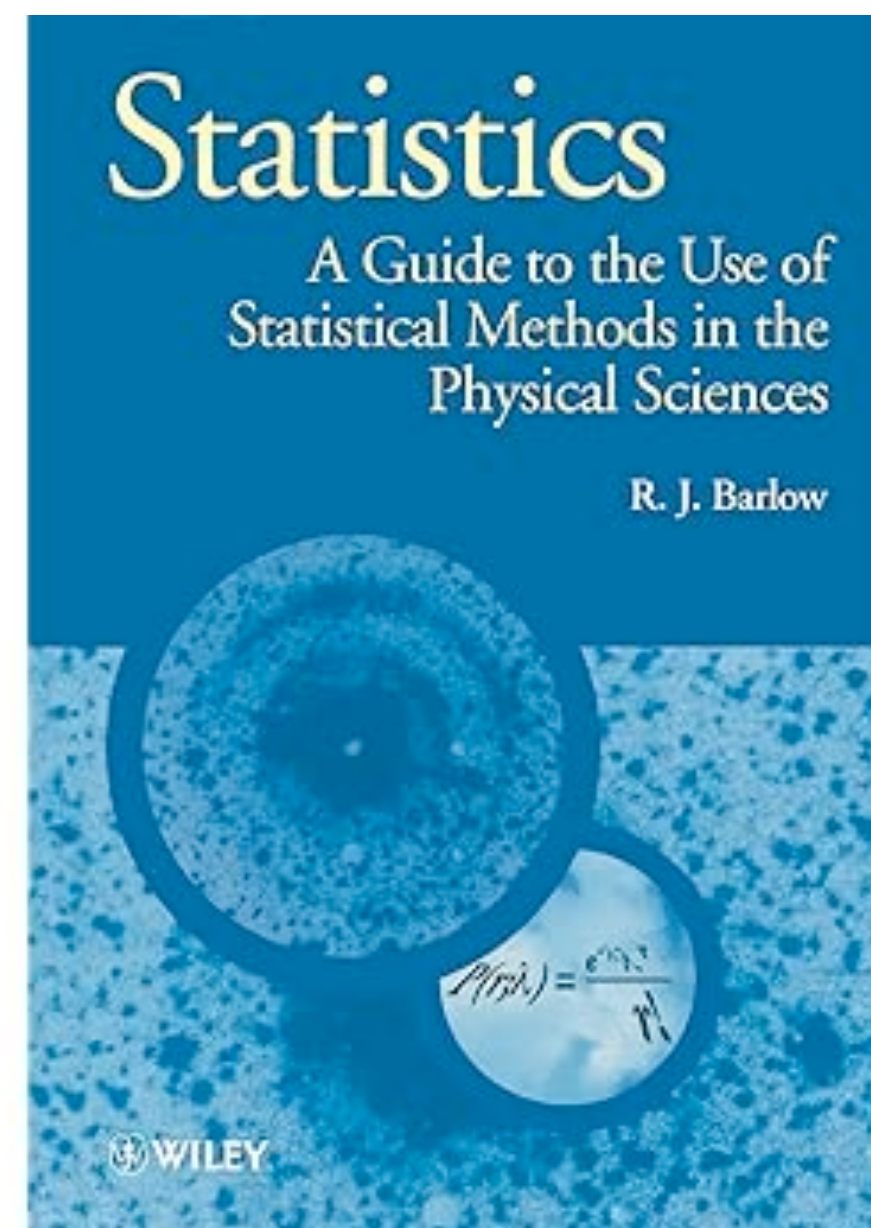
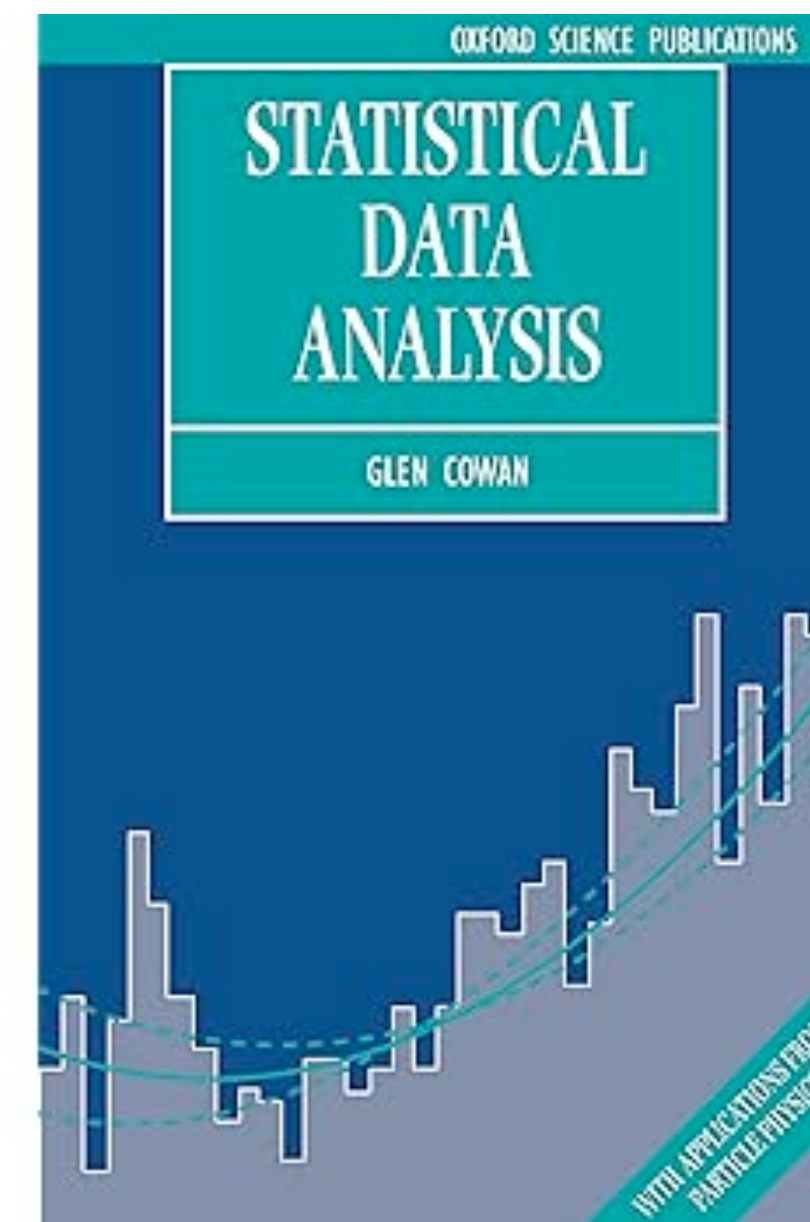


Image: [ML4FP School](#)

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Publication date : November 1, 1993



Publication date : December 4, 1997

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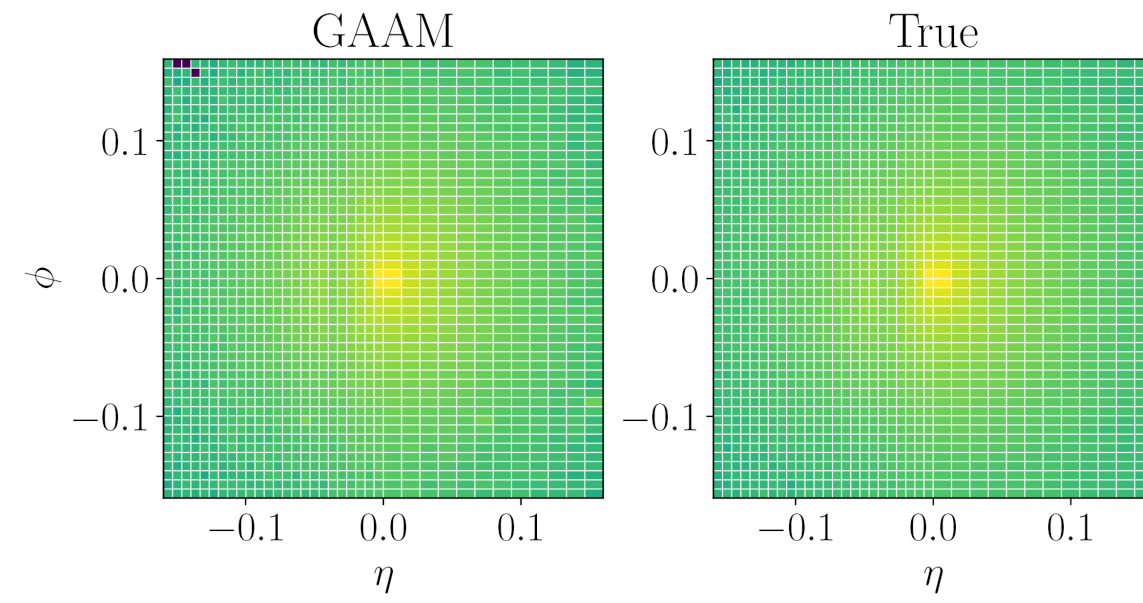


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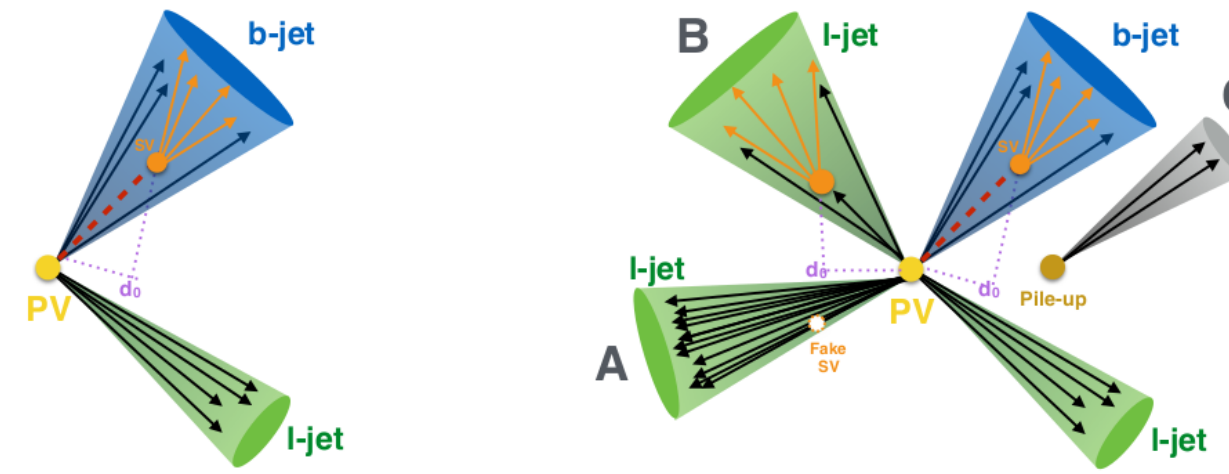


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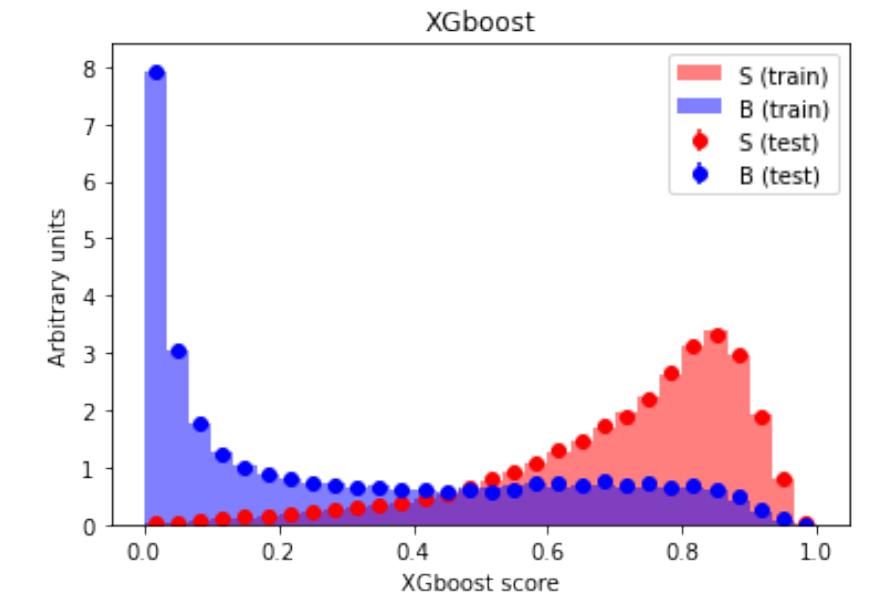
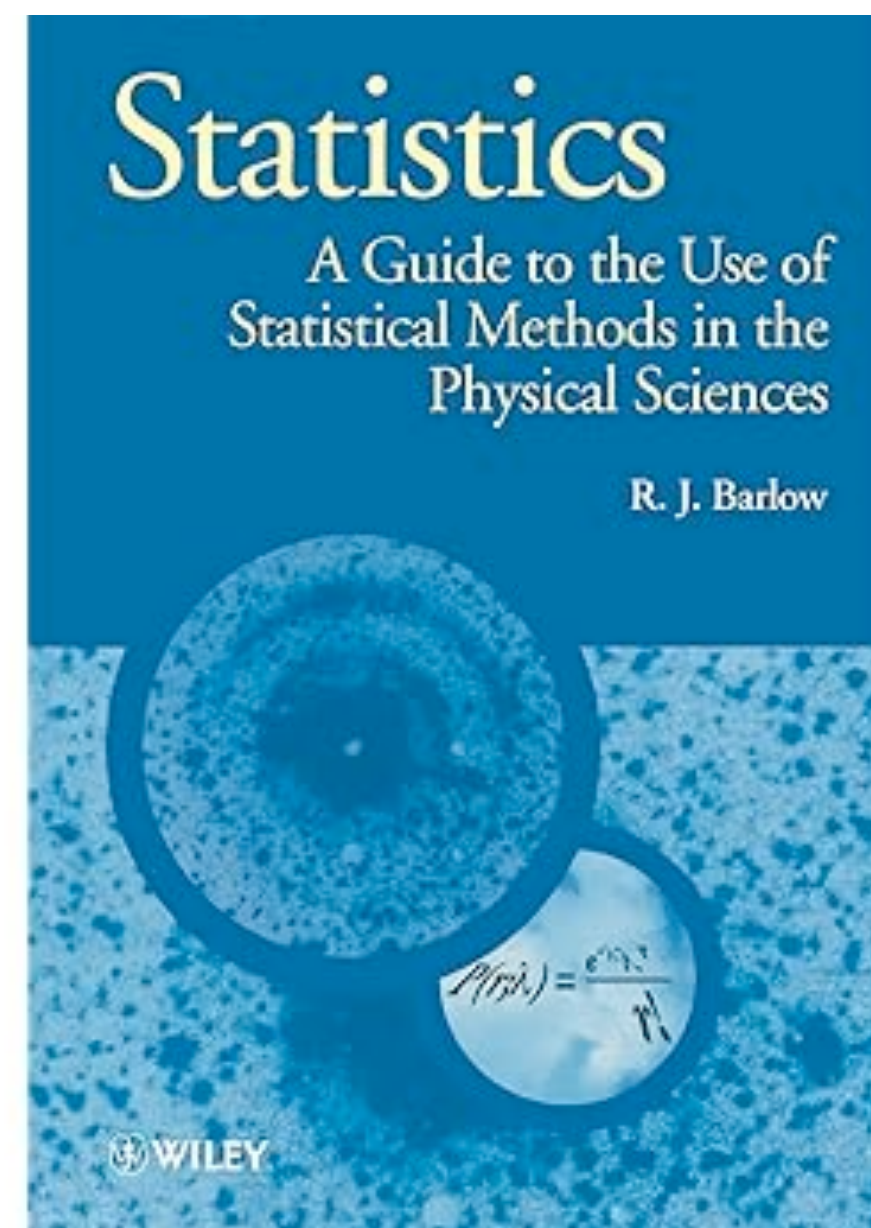


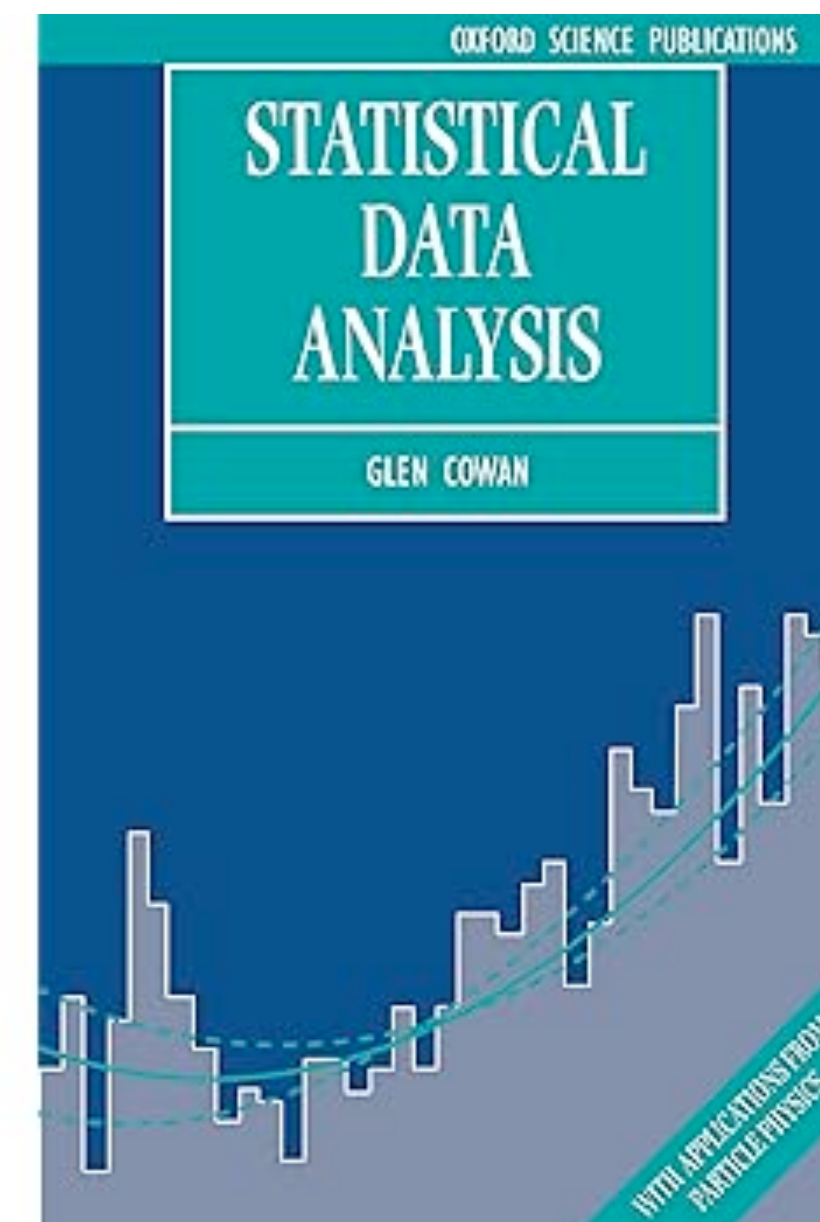
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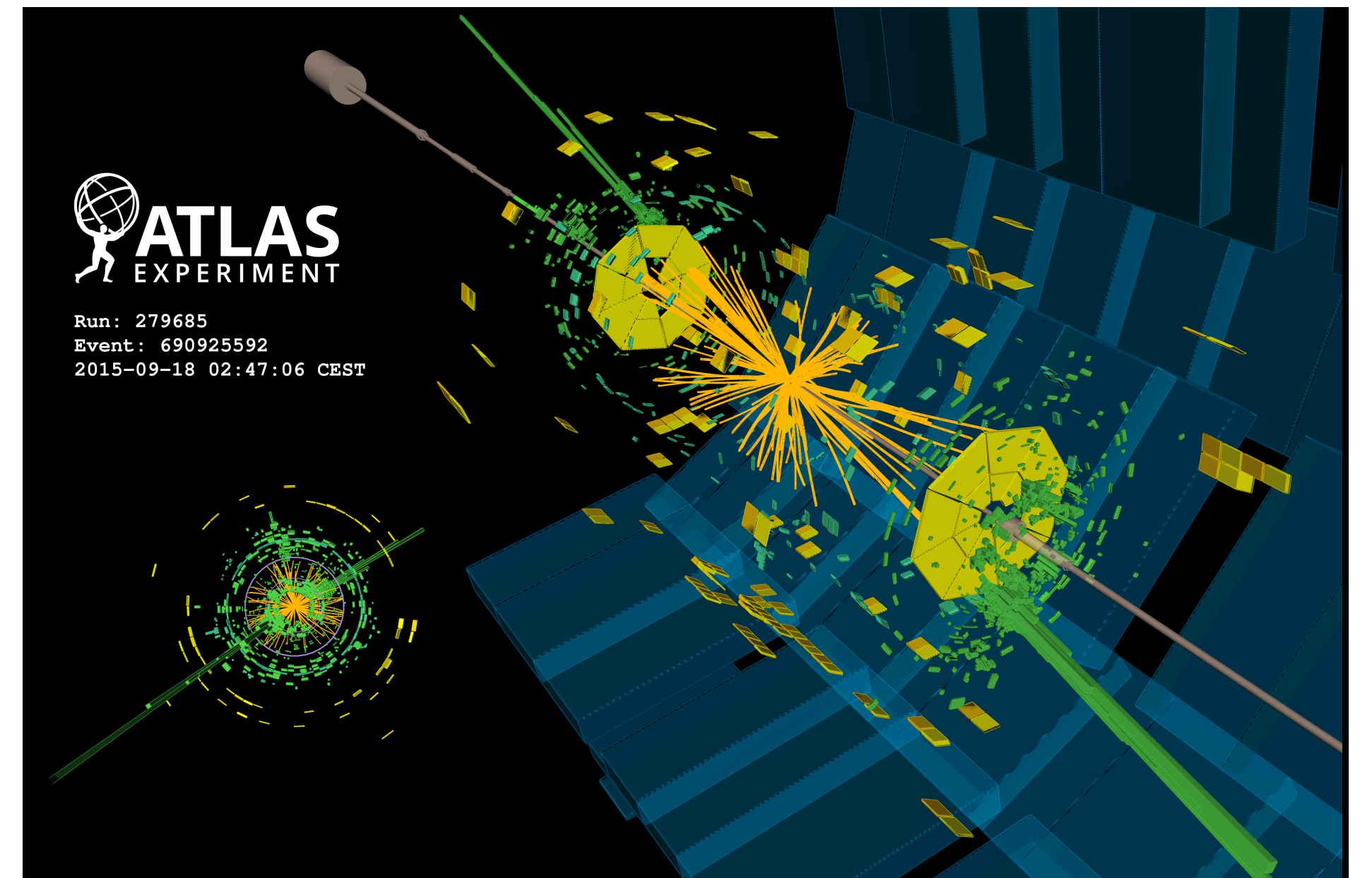
Time to modernise?



**Publication date : December 4, 1997**

# Traditional Approach: Design one sensitive observable

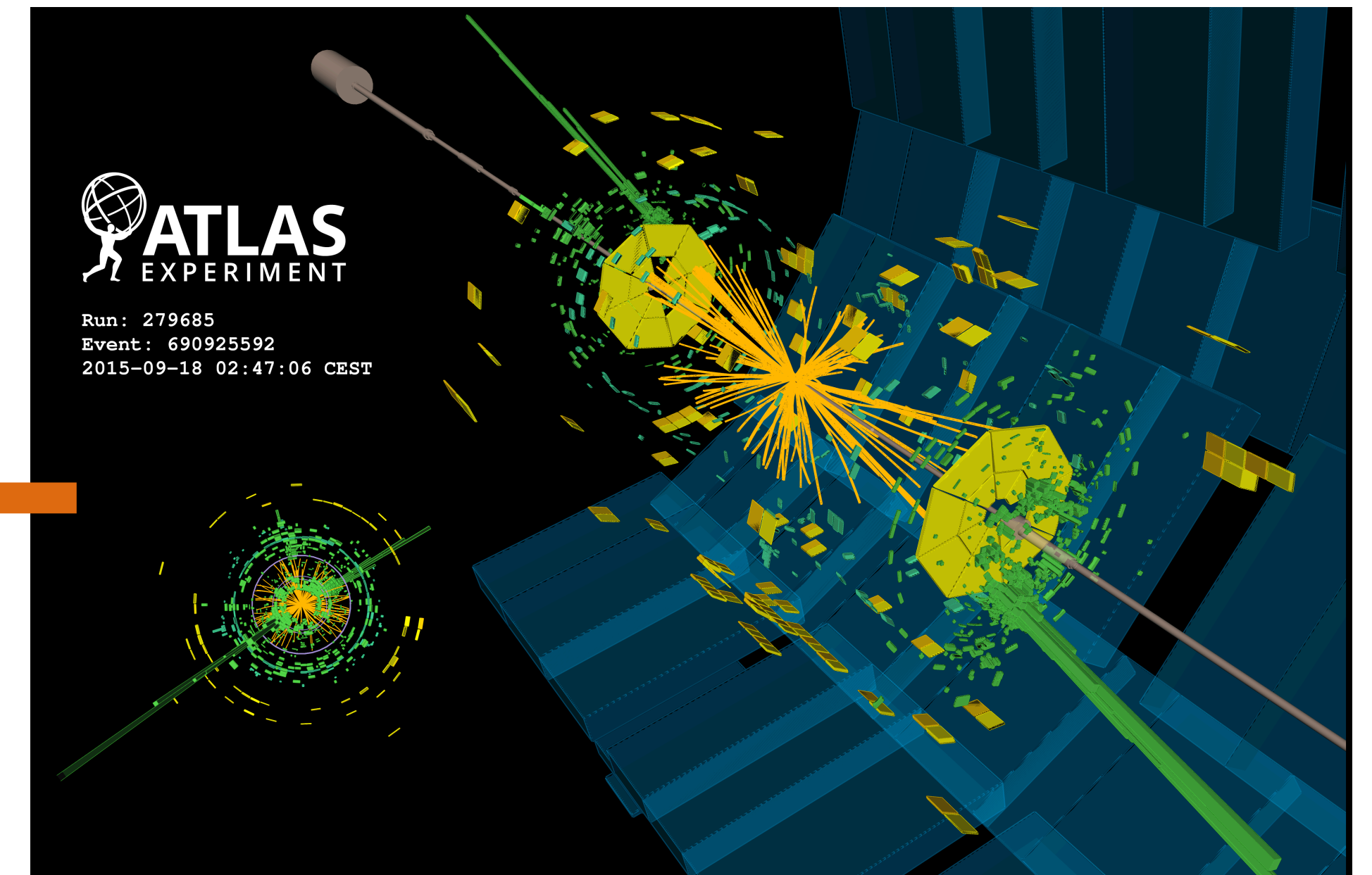
- Detector has  $O(100 \text{ million})$  sensors
- Reconstruction pipeline, event selection
- Design one summary variable
  - Compression:  $O(100 \text{ million}) \rightarrow 1$
- Build a 1-D histogram
- Calculate likelihoods with histograms



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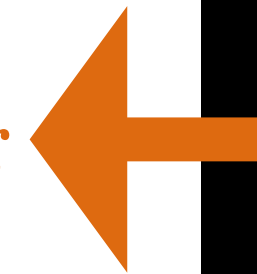
1 number



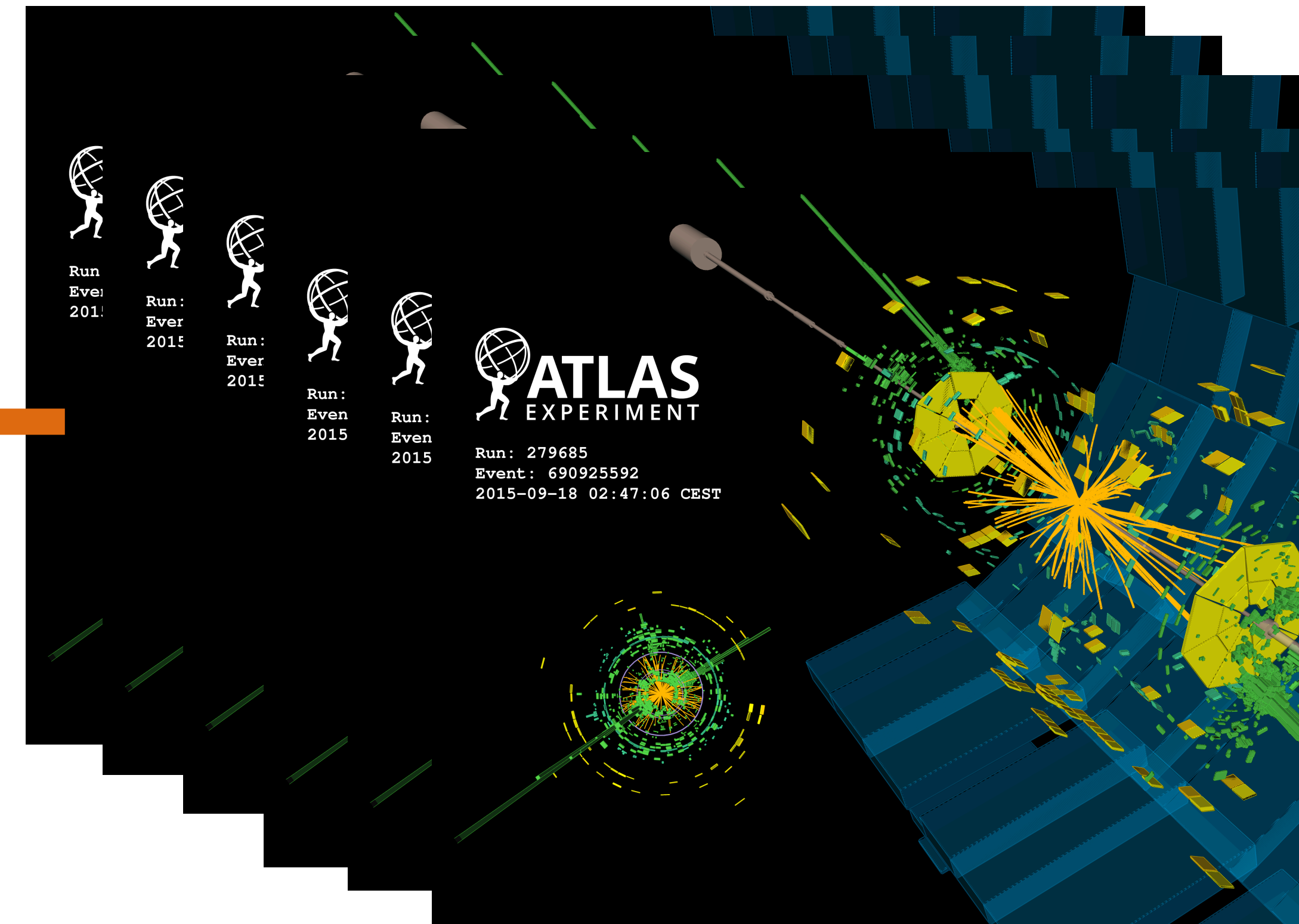
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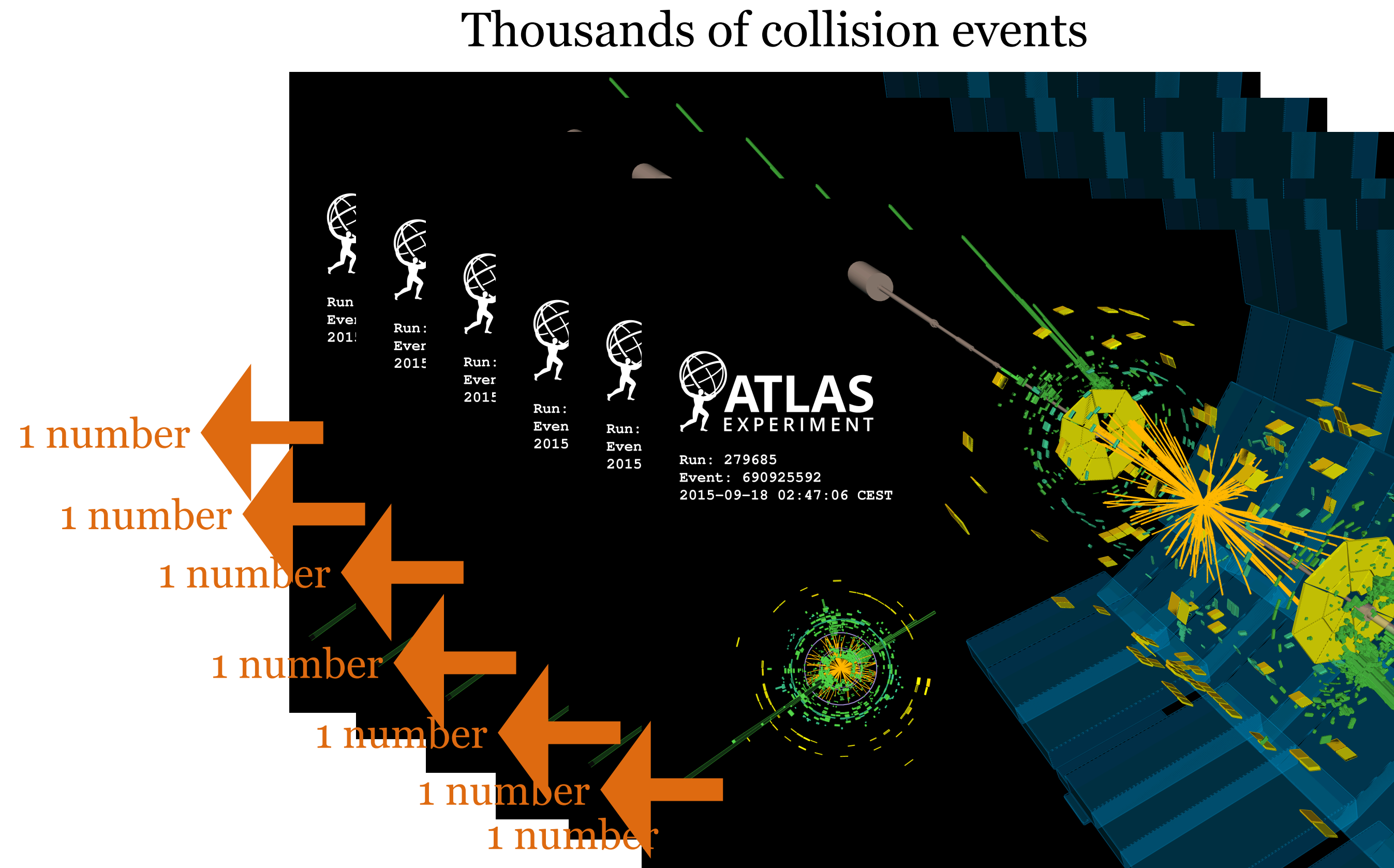


Thousands of collision events



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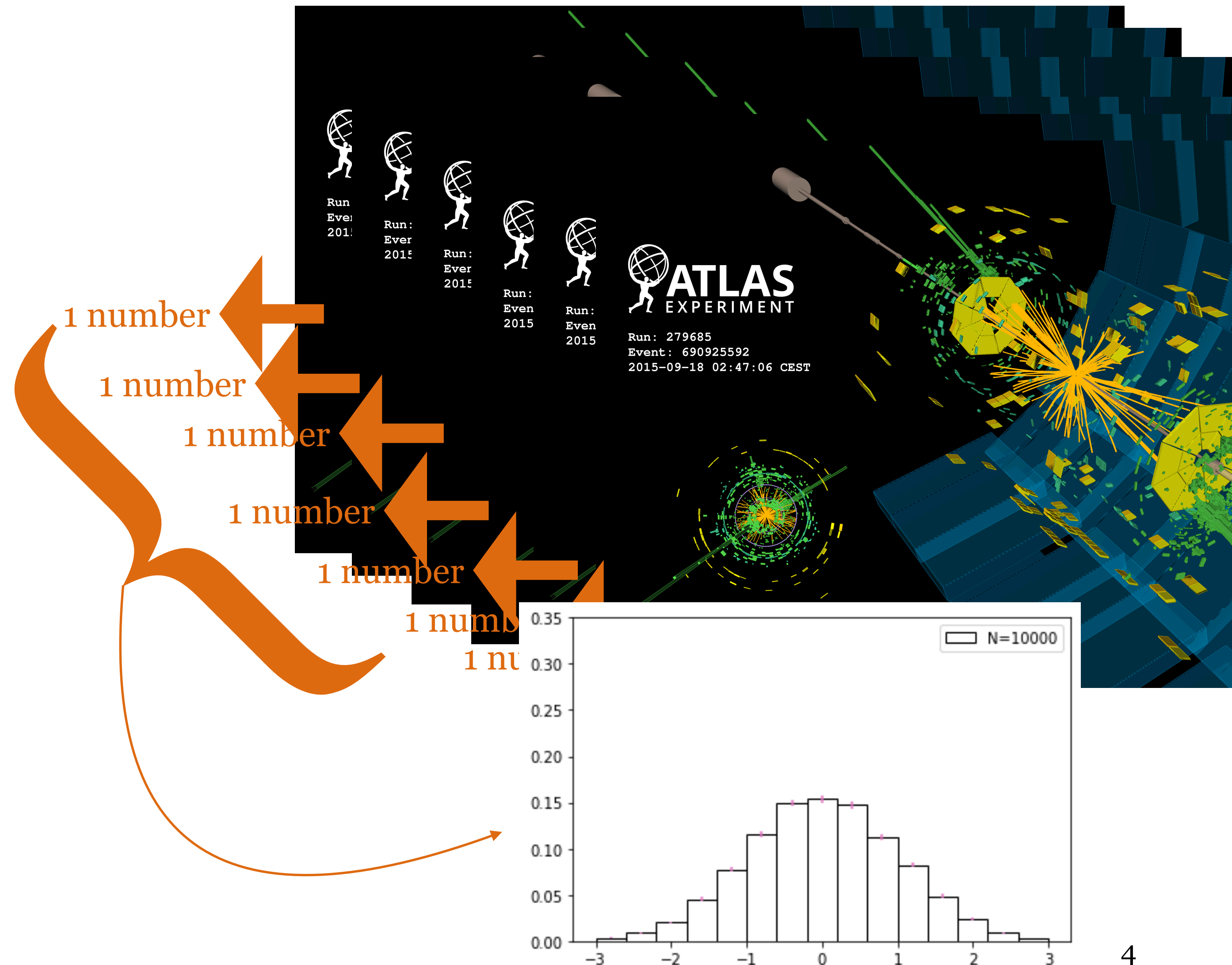
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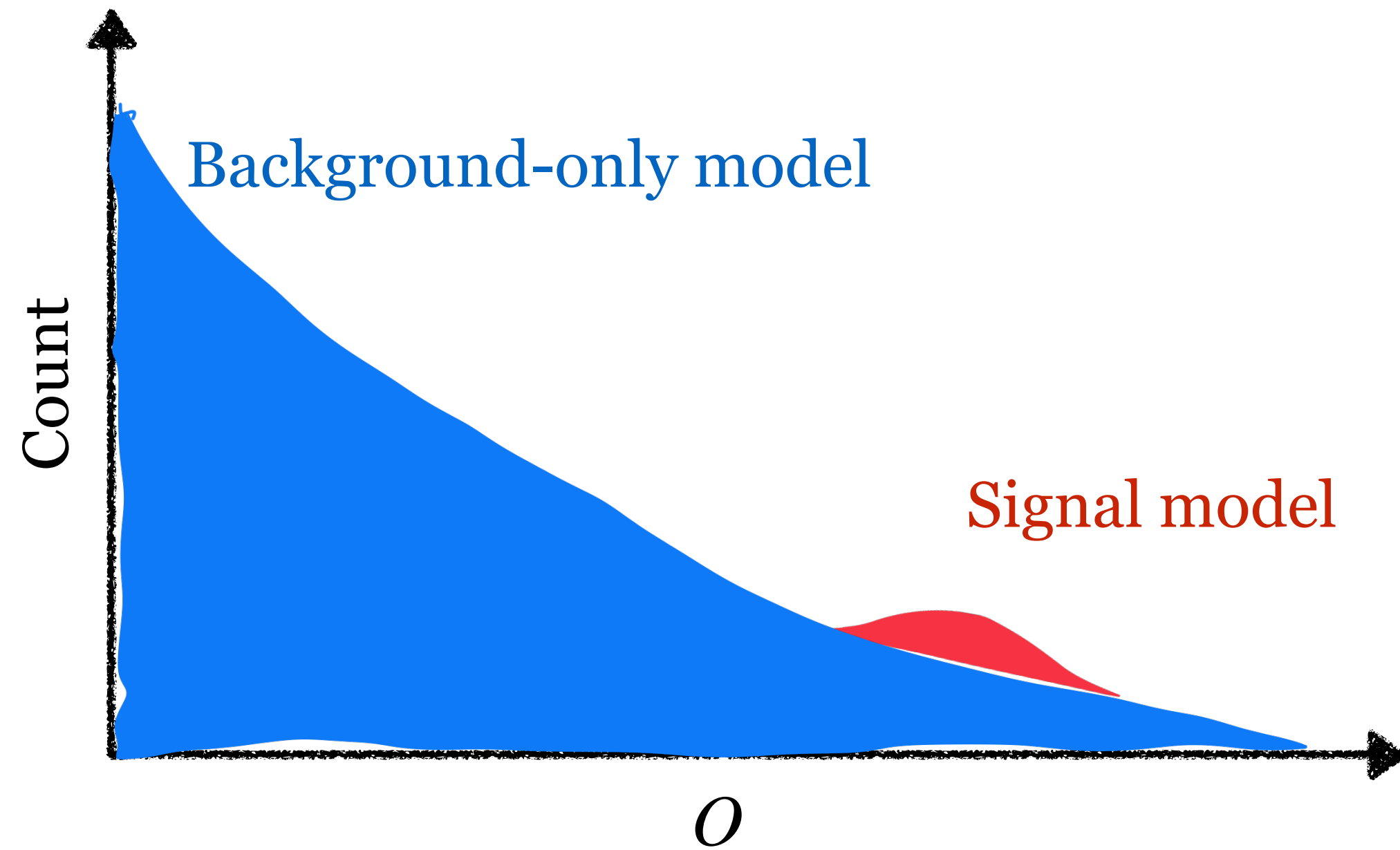
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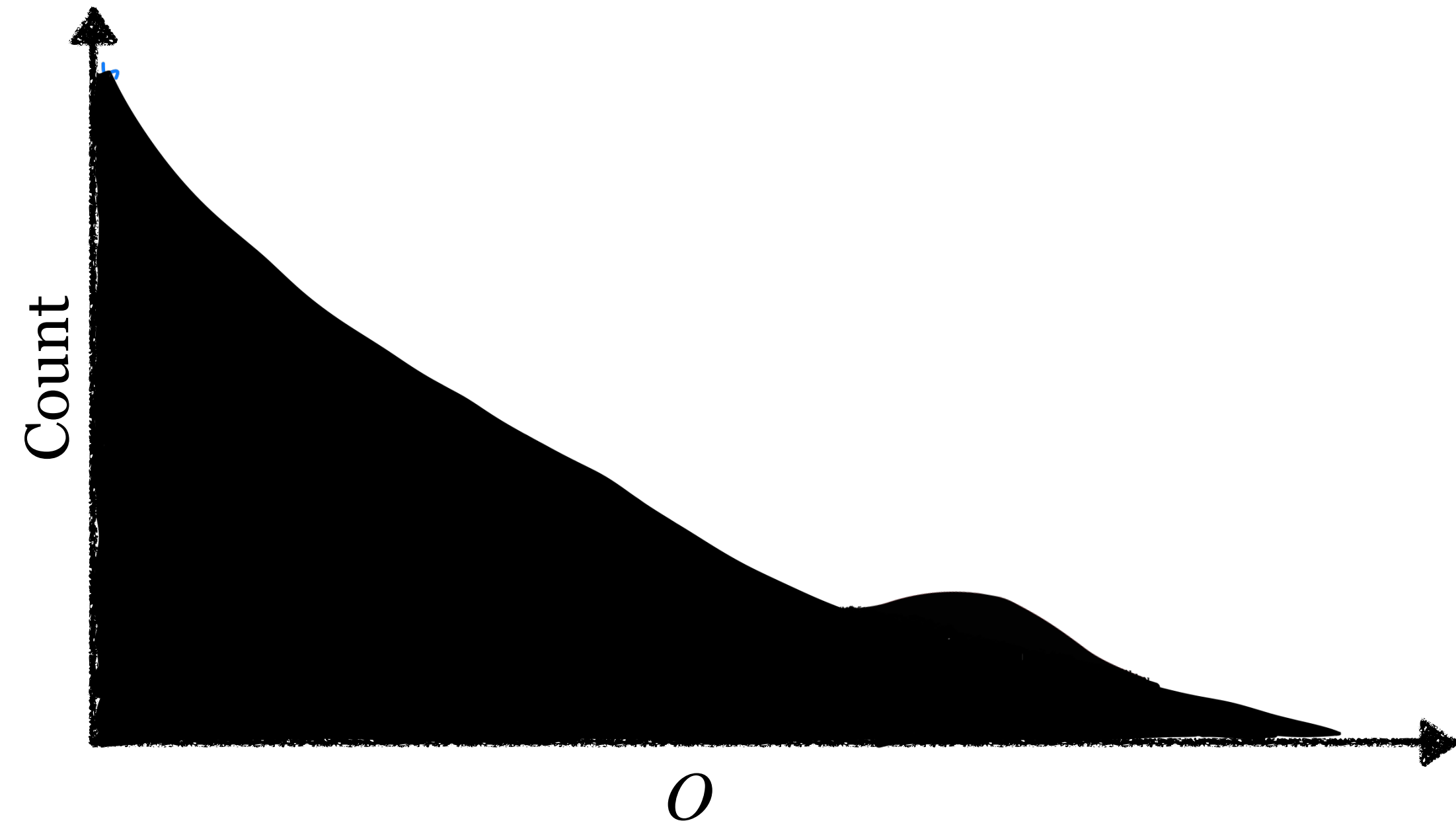


# Probability density estimation with 1-D histogram

Theory Predictions

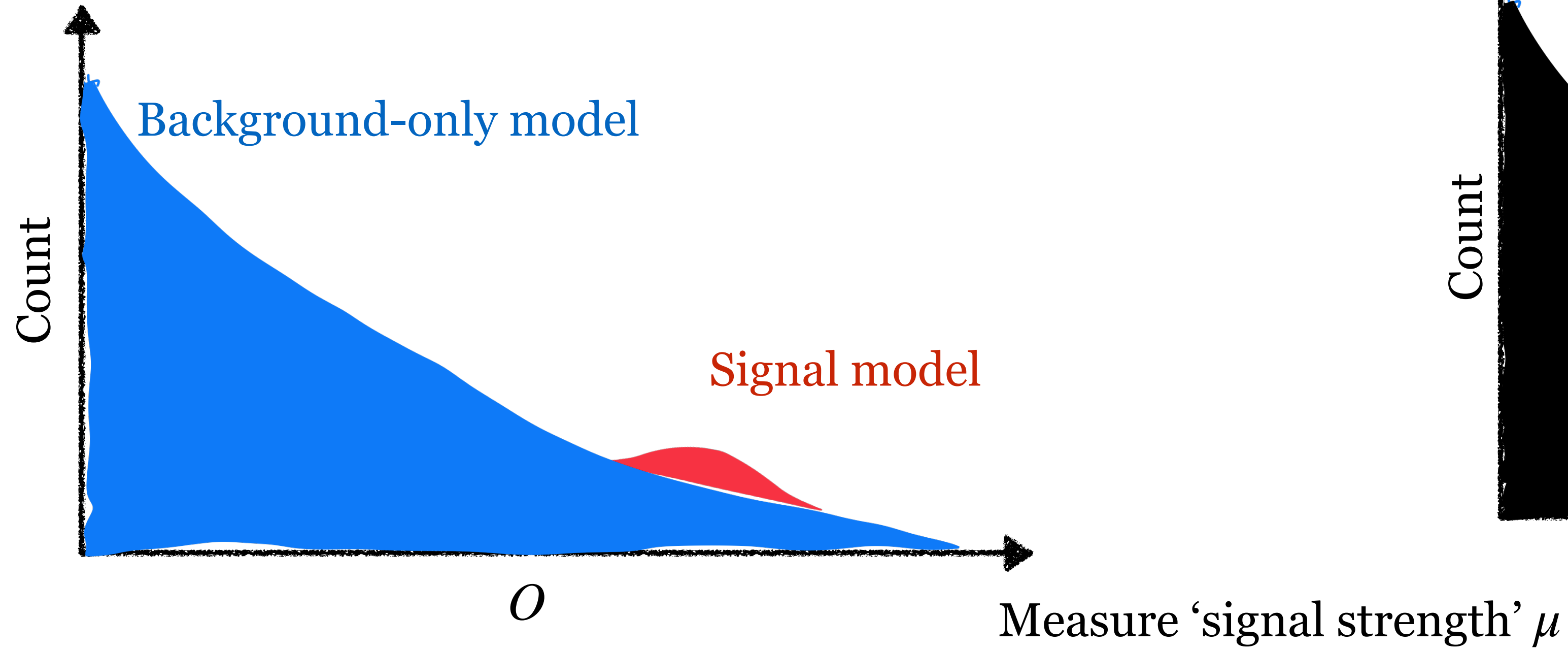


Data

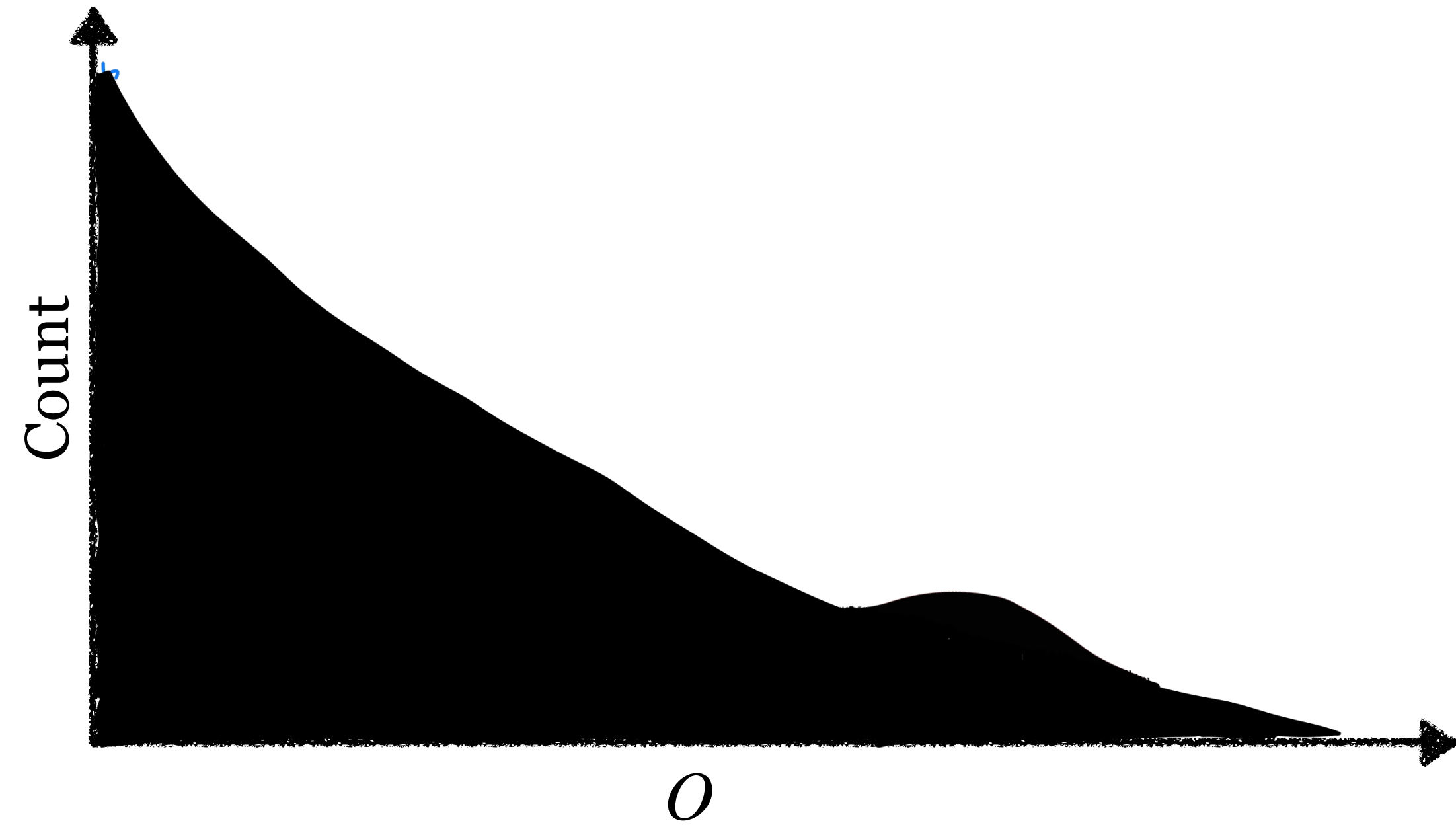


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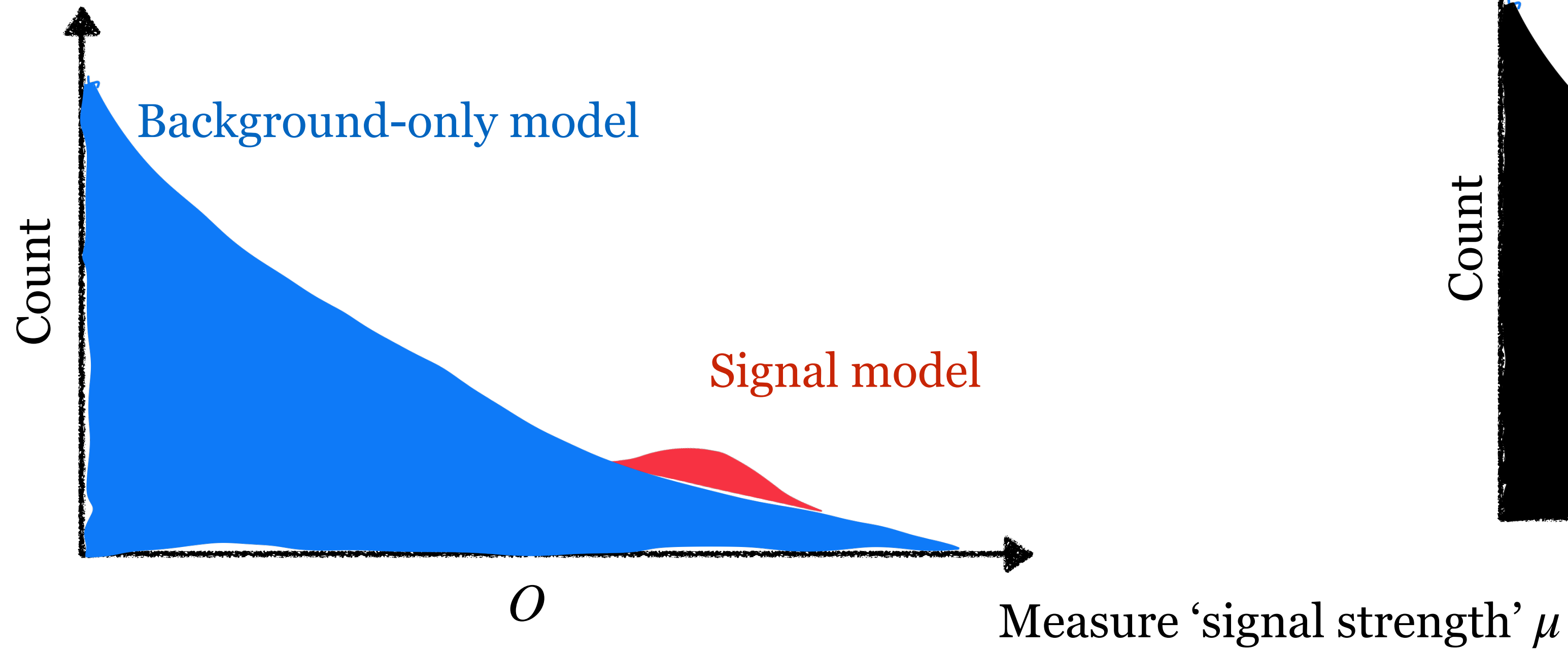


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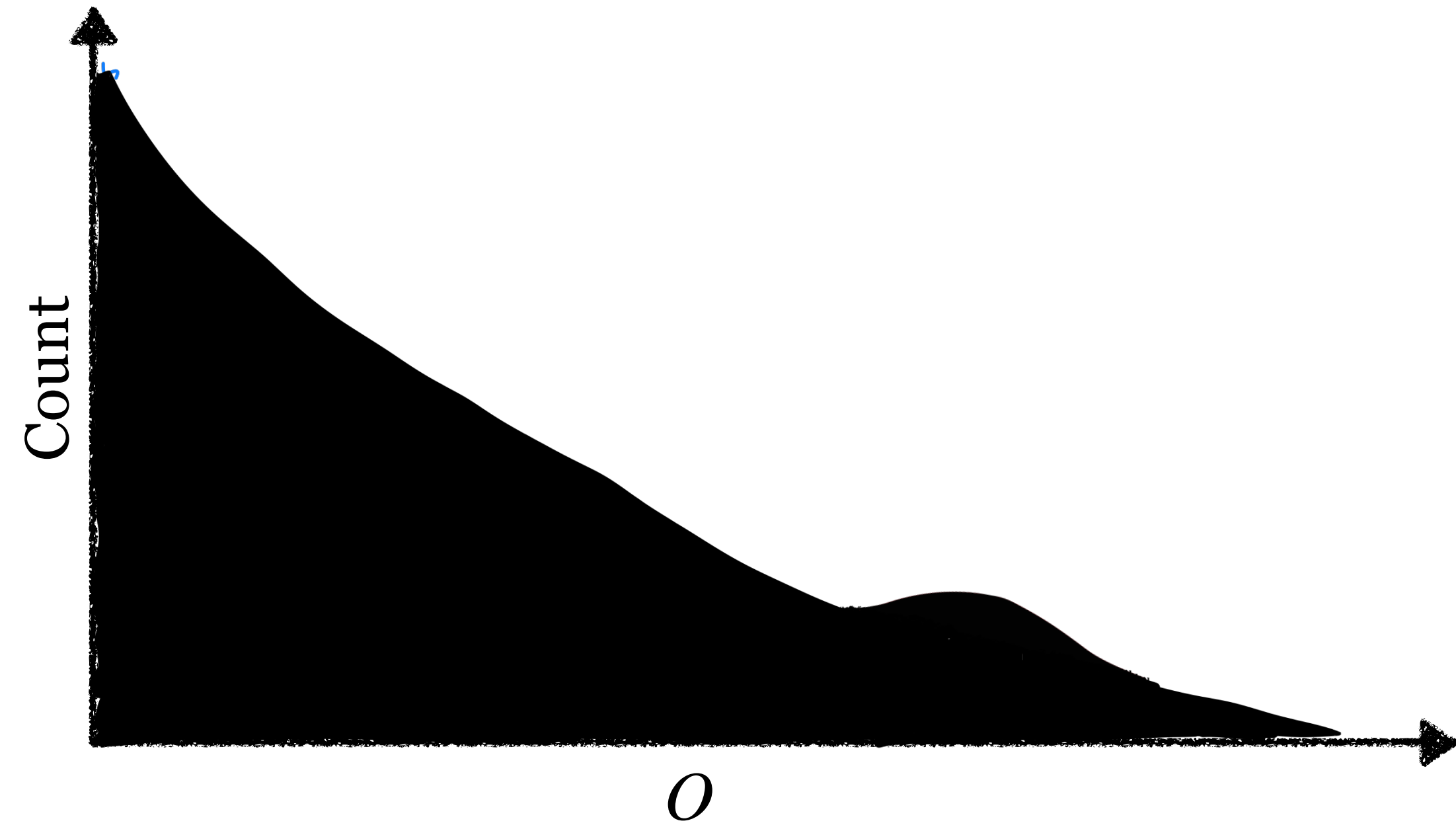


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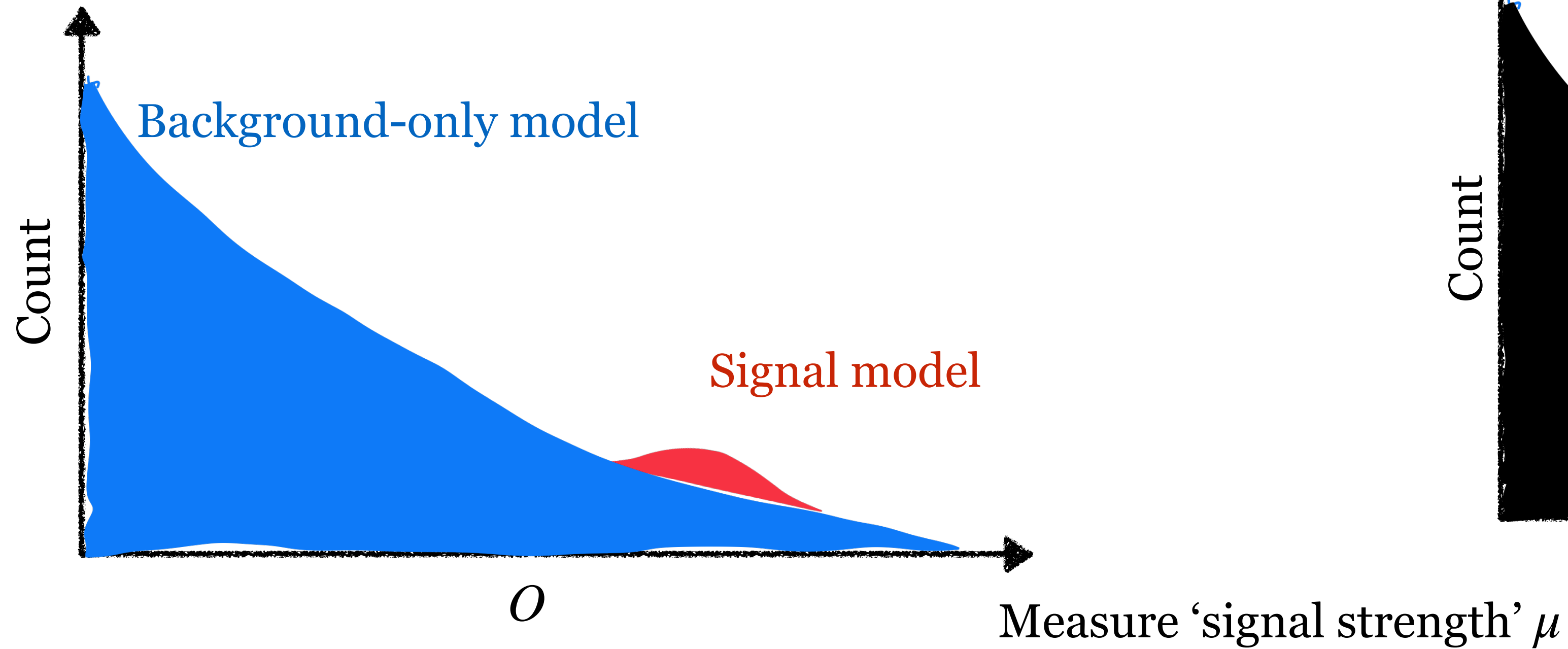


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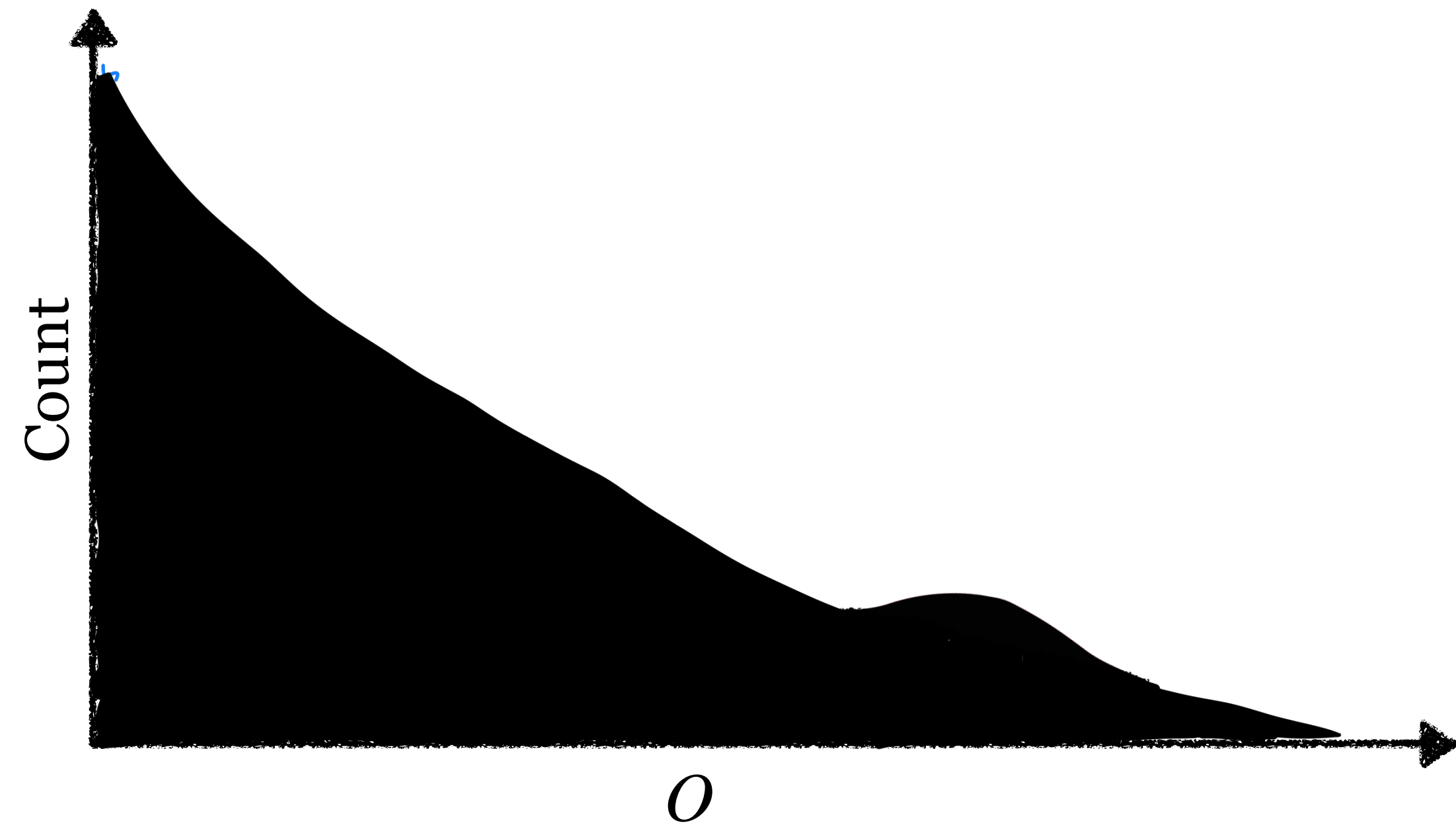


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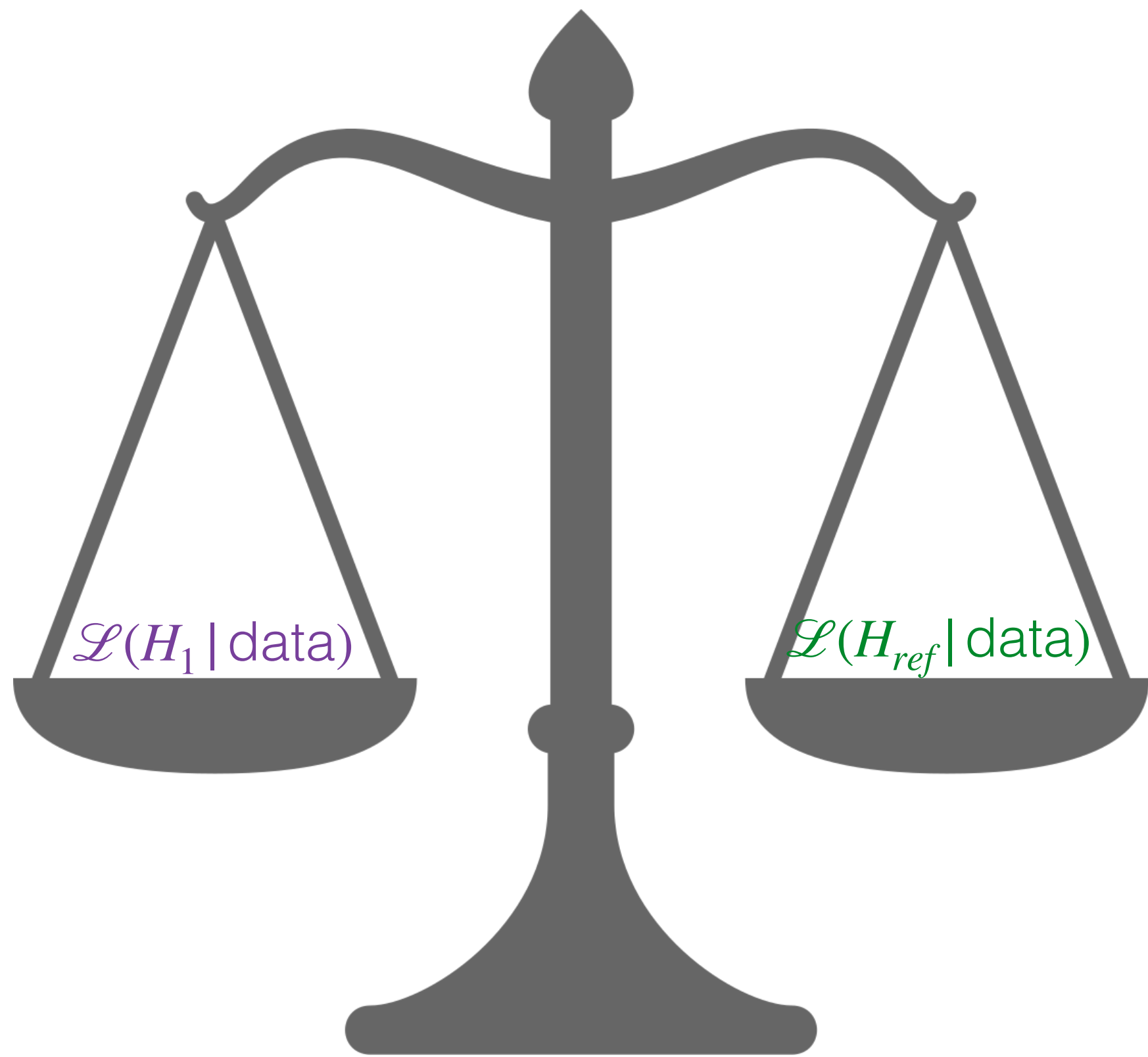
Data



With histograms we can ask “Given the data, what is the likelihood of  $\mu = 1$  hypothesis vs  $\mu = 2$  hypothesis?”

# Hypothesis tests

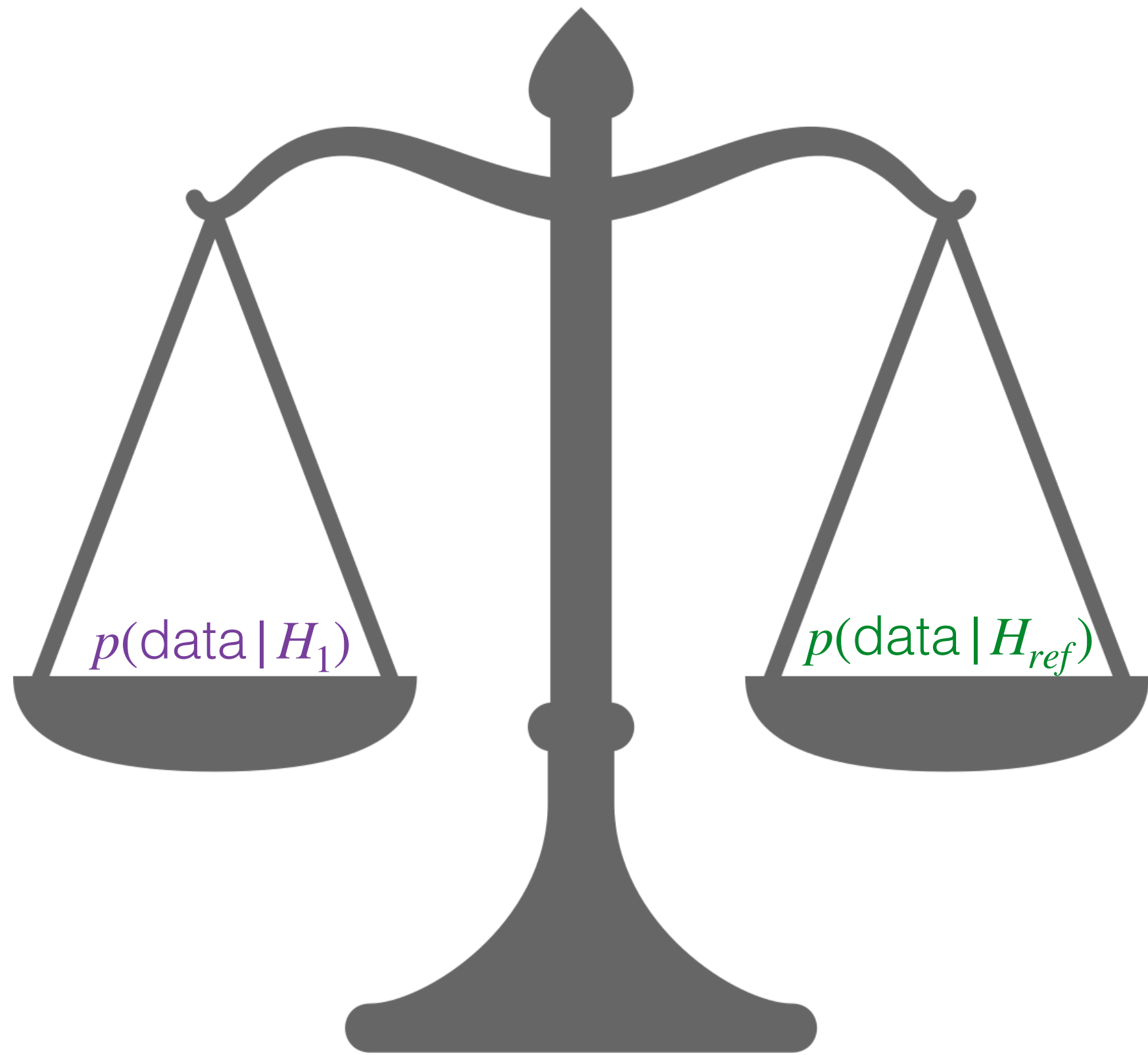
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# Hypothesis tests

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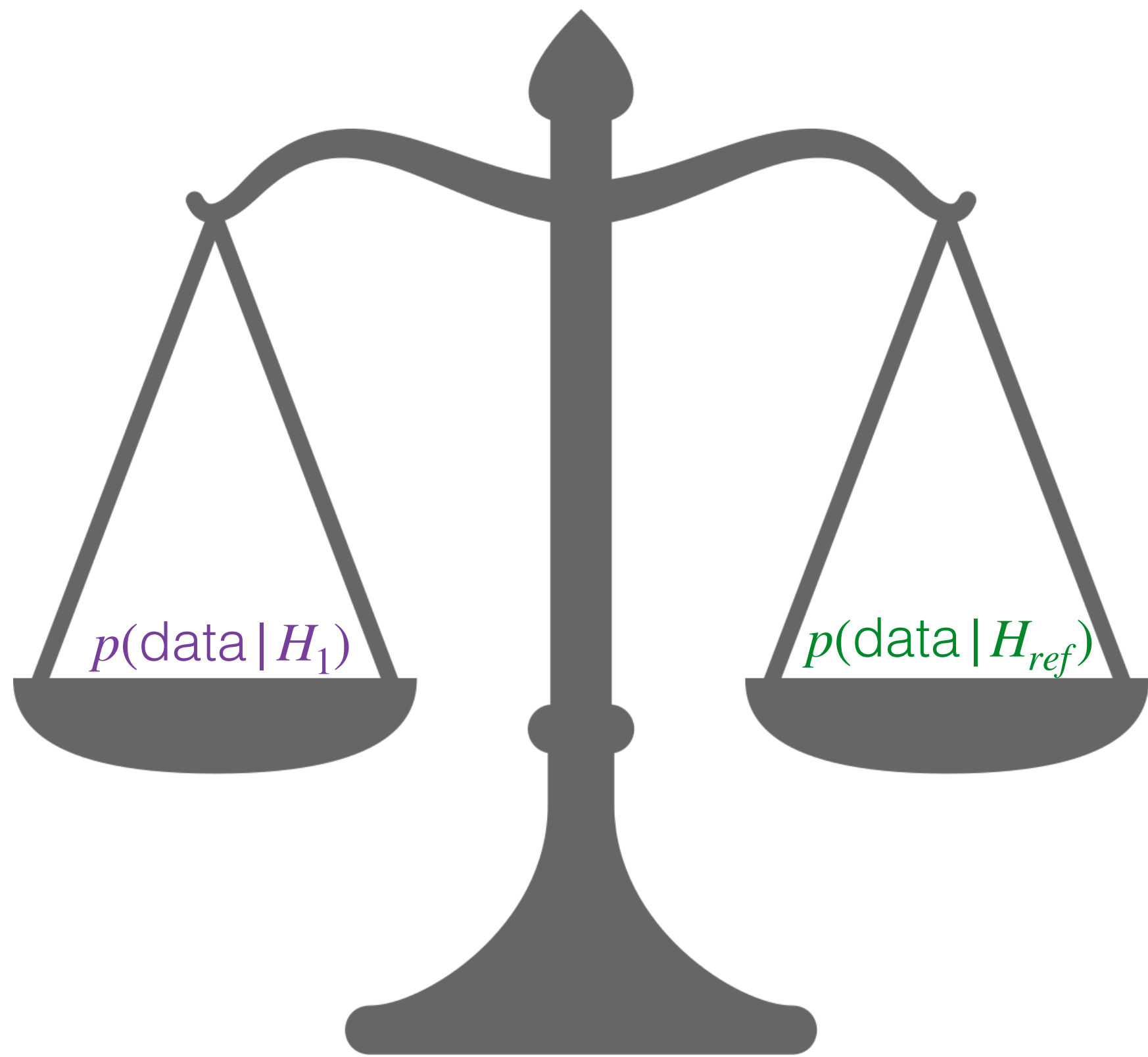
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# Hypothesis tests

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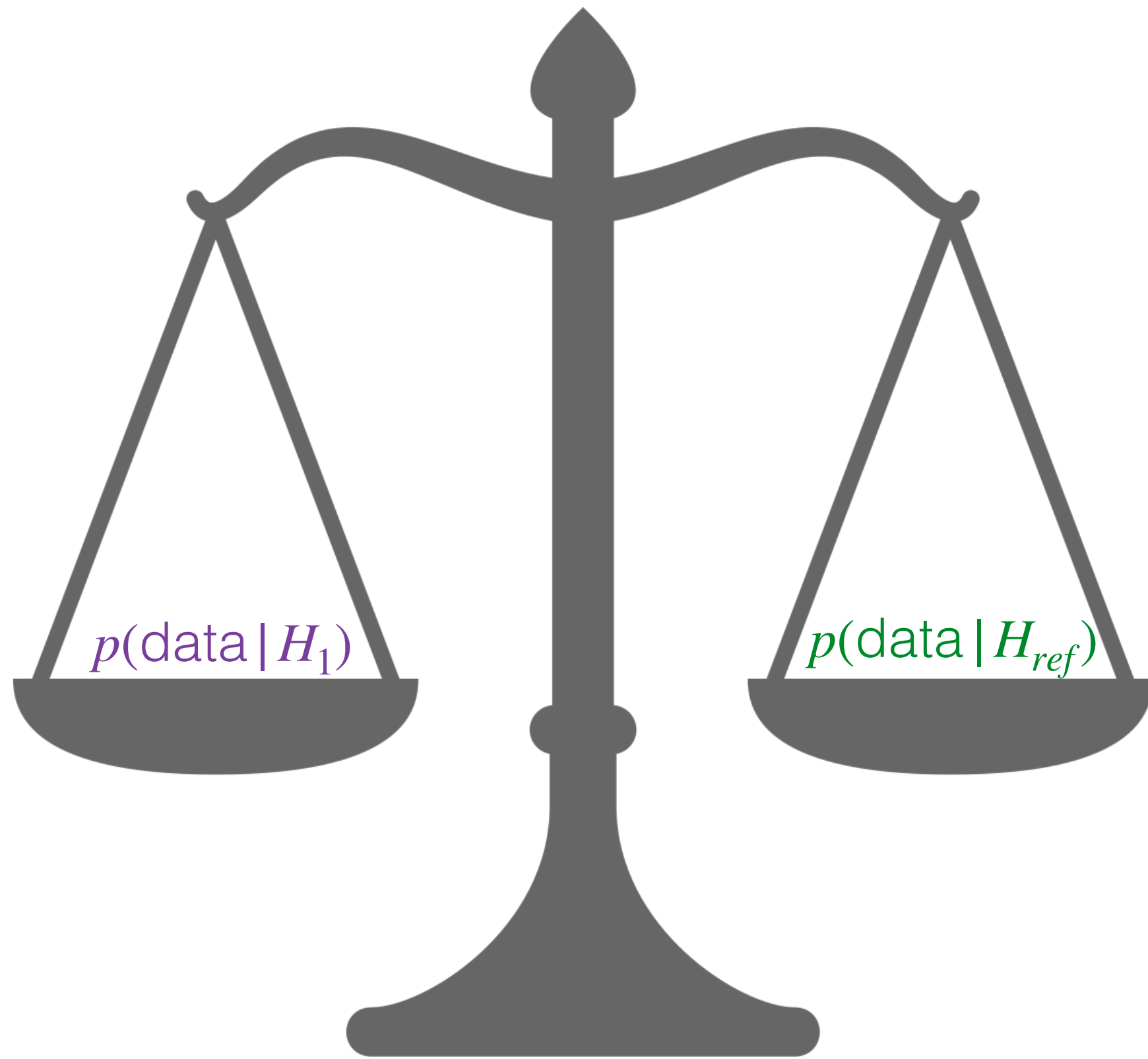
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When comparing 2 hypotheses, LR guaranteed to be optimal test by Neyman-Person lemma

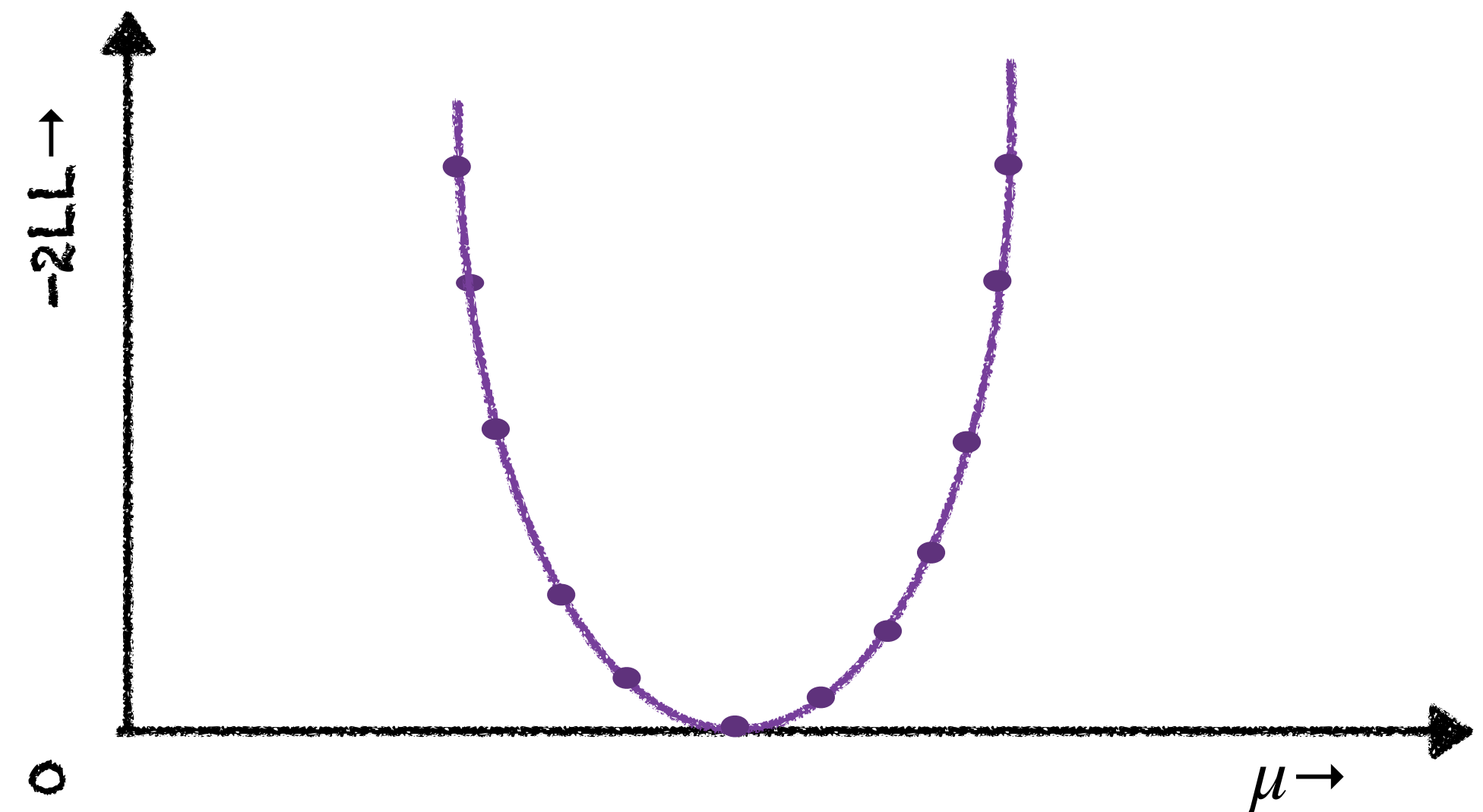
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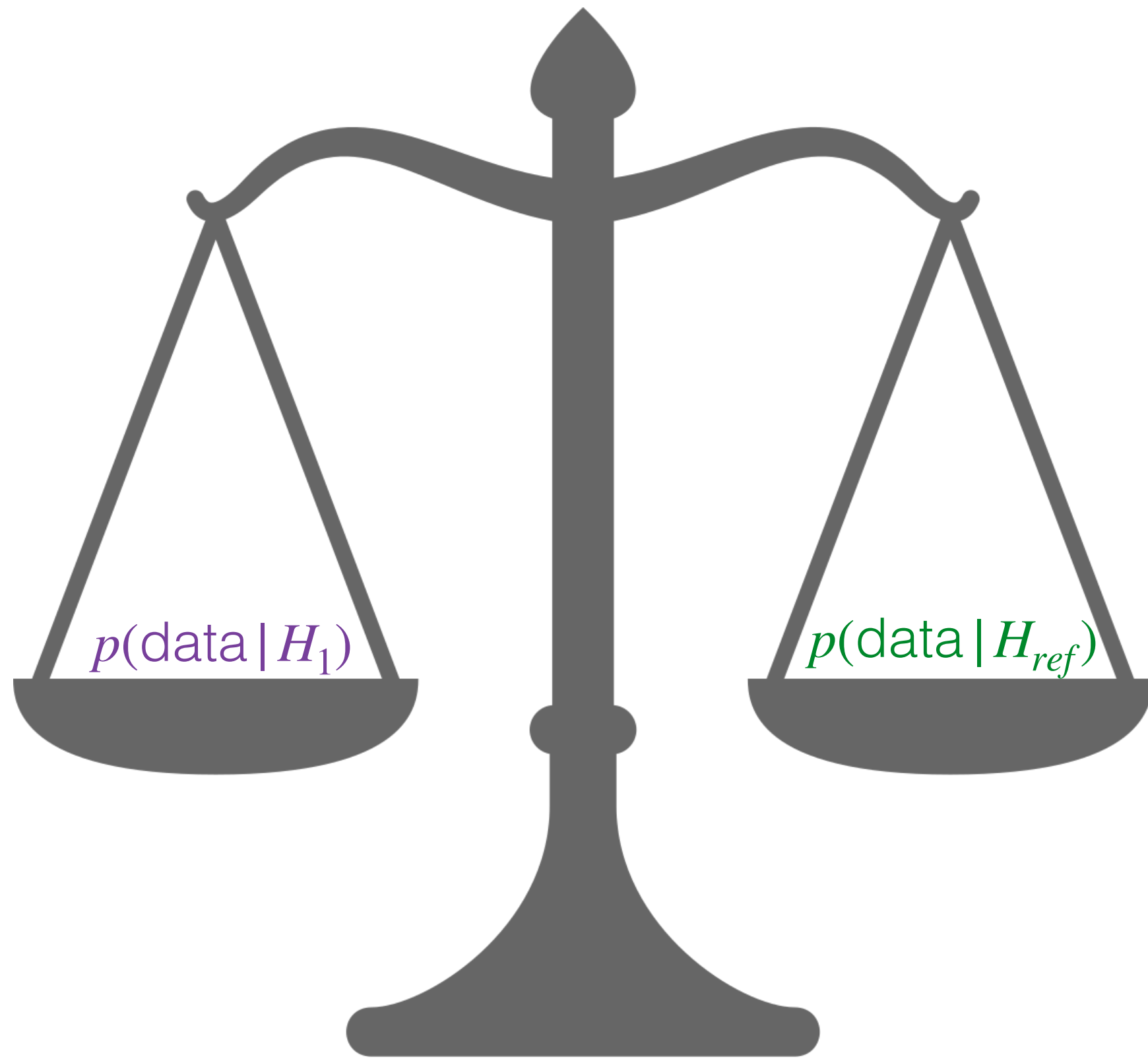
Invert the test for confidence intervals



Parameter estimation (infinite hypotheses)

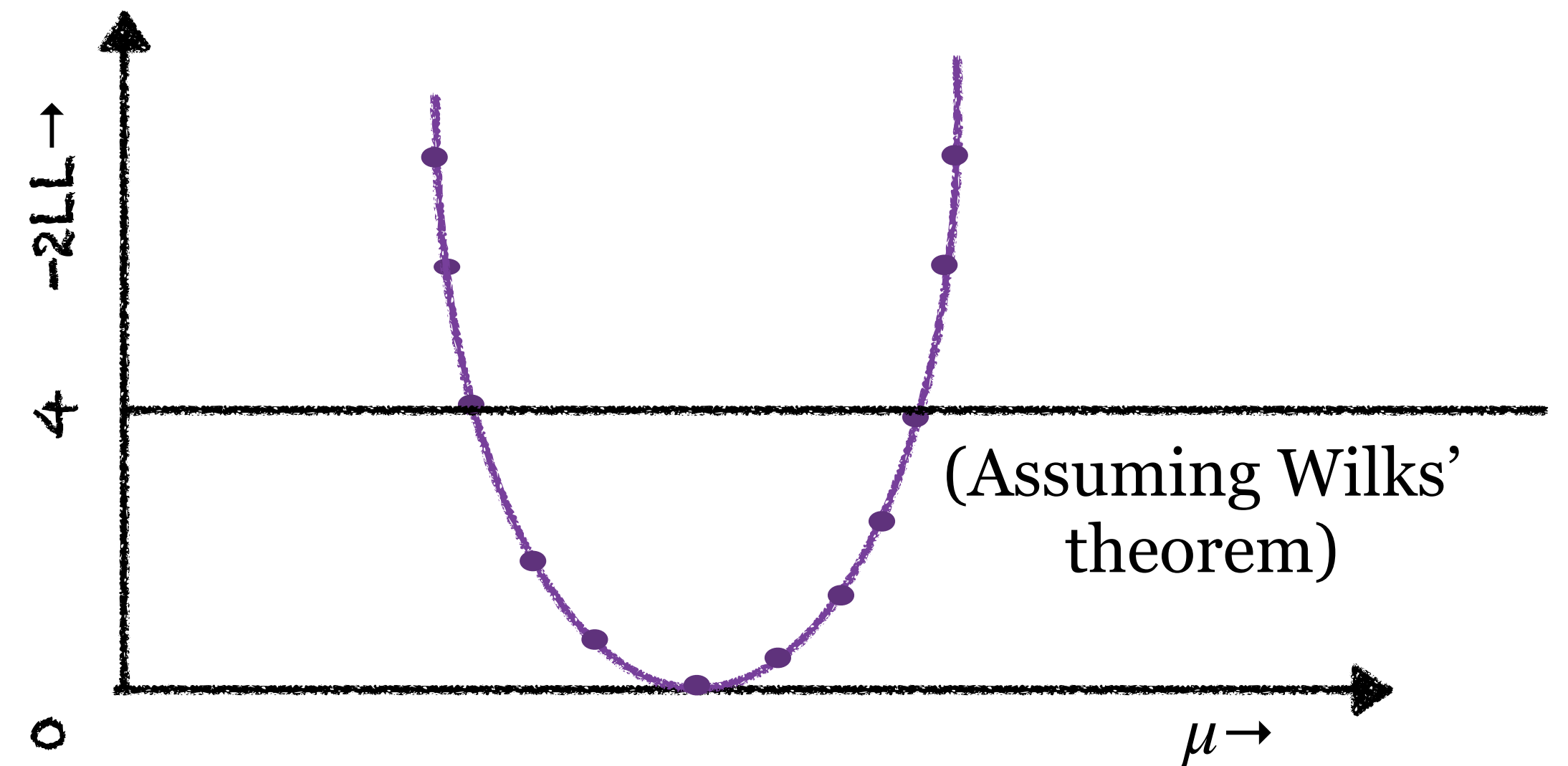
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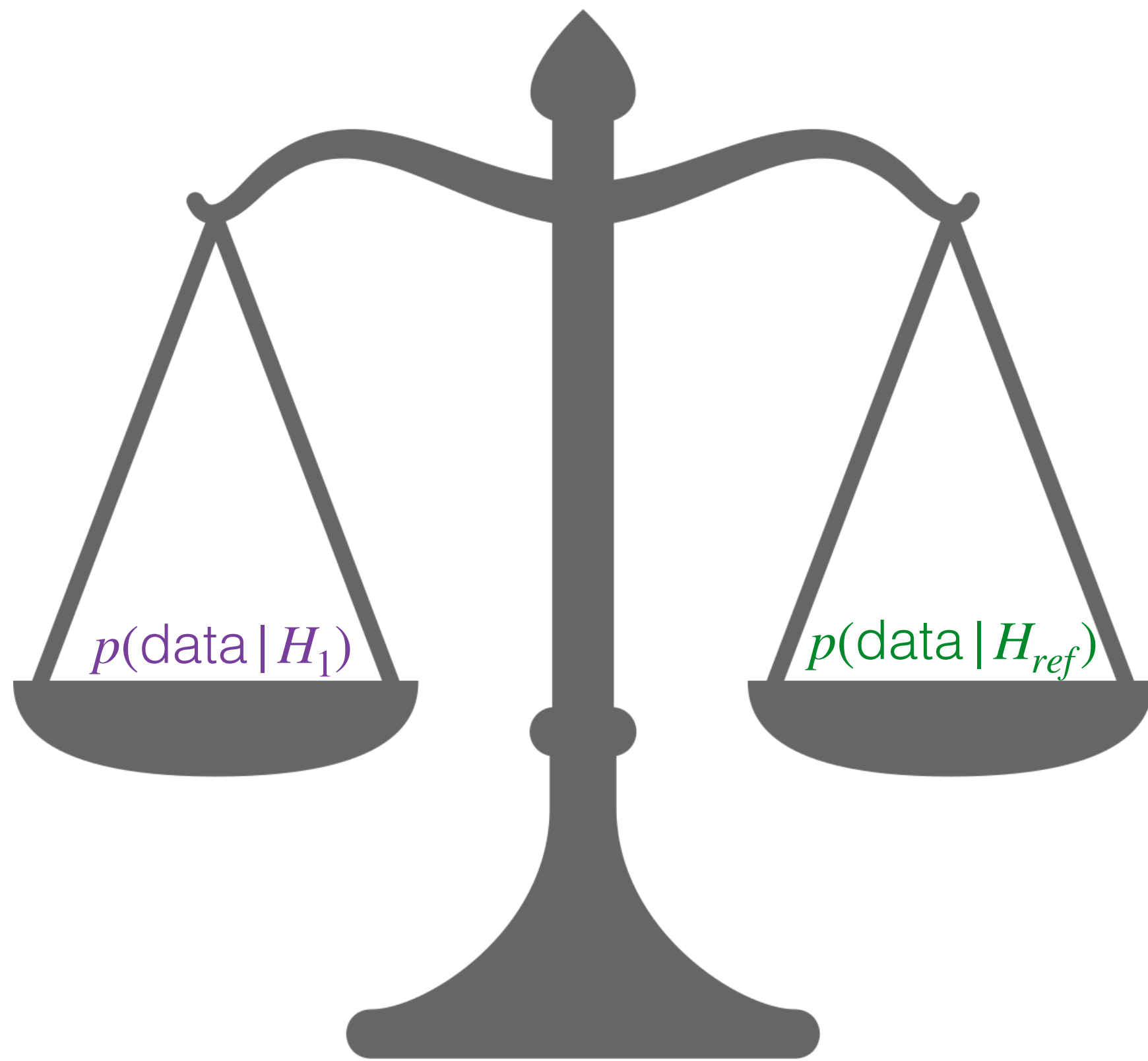
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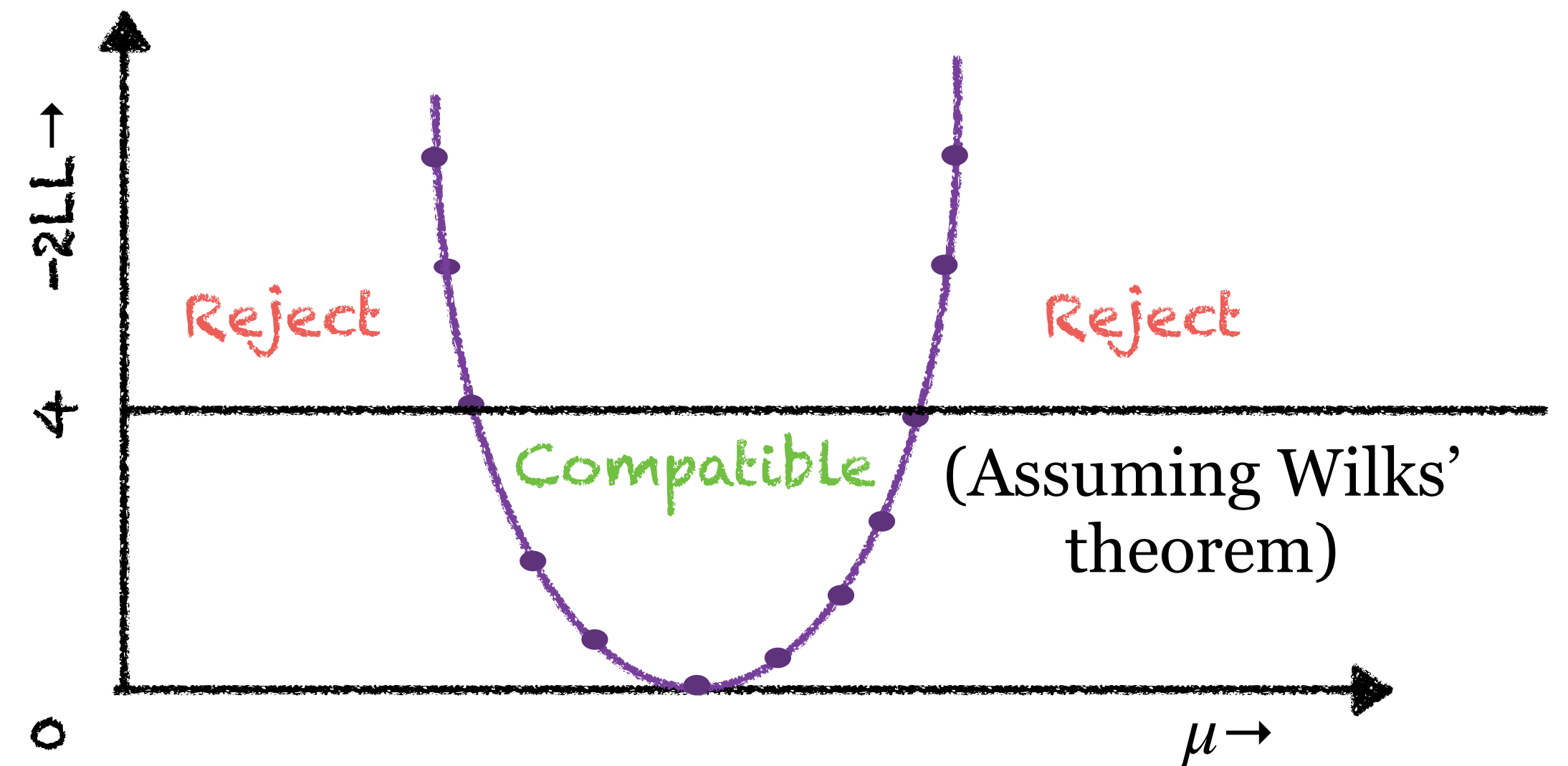
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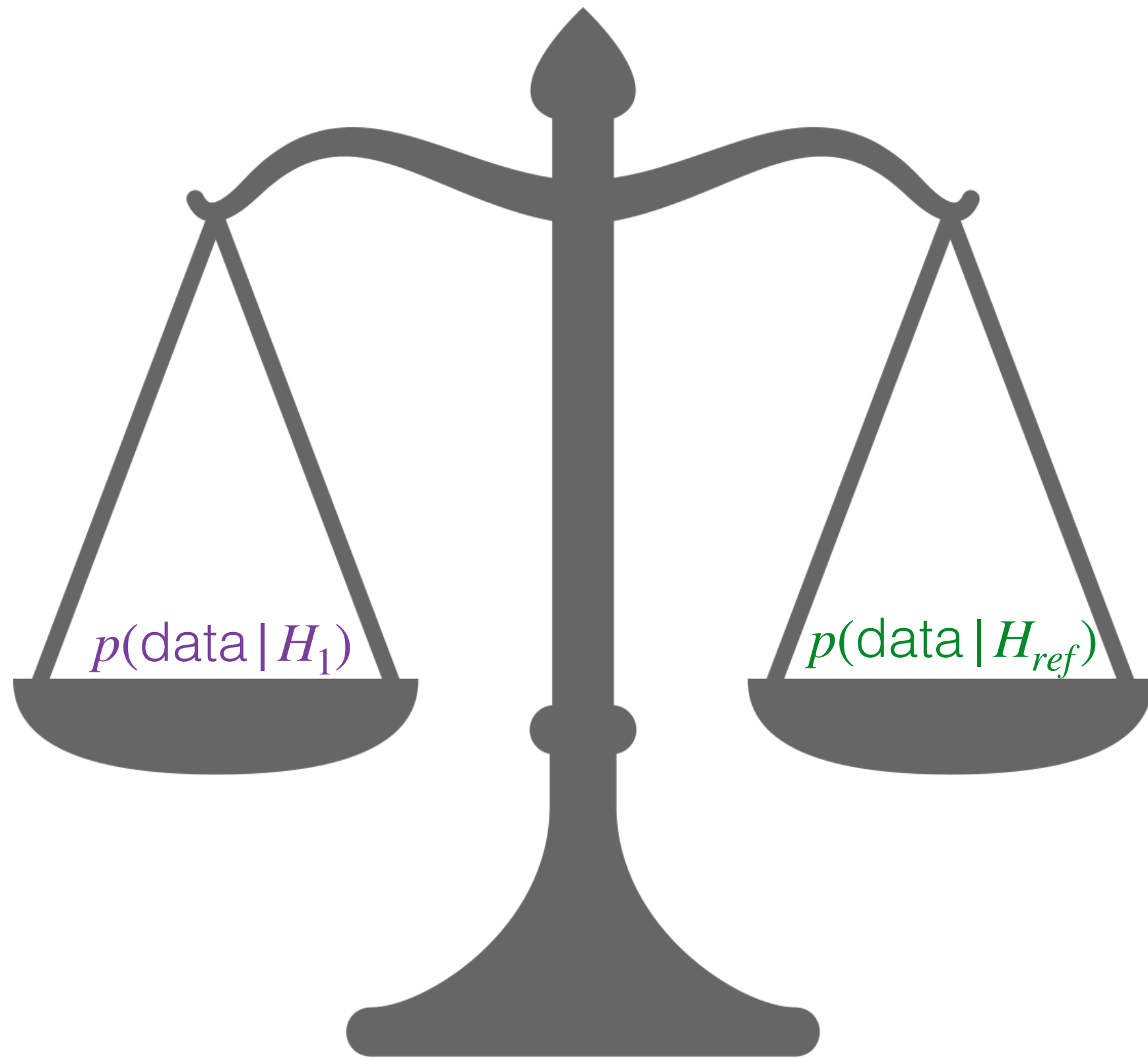
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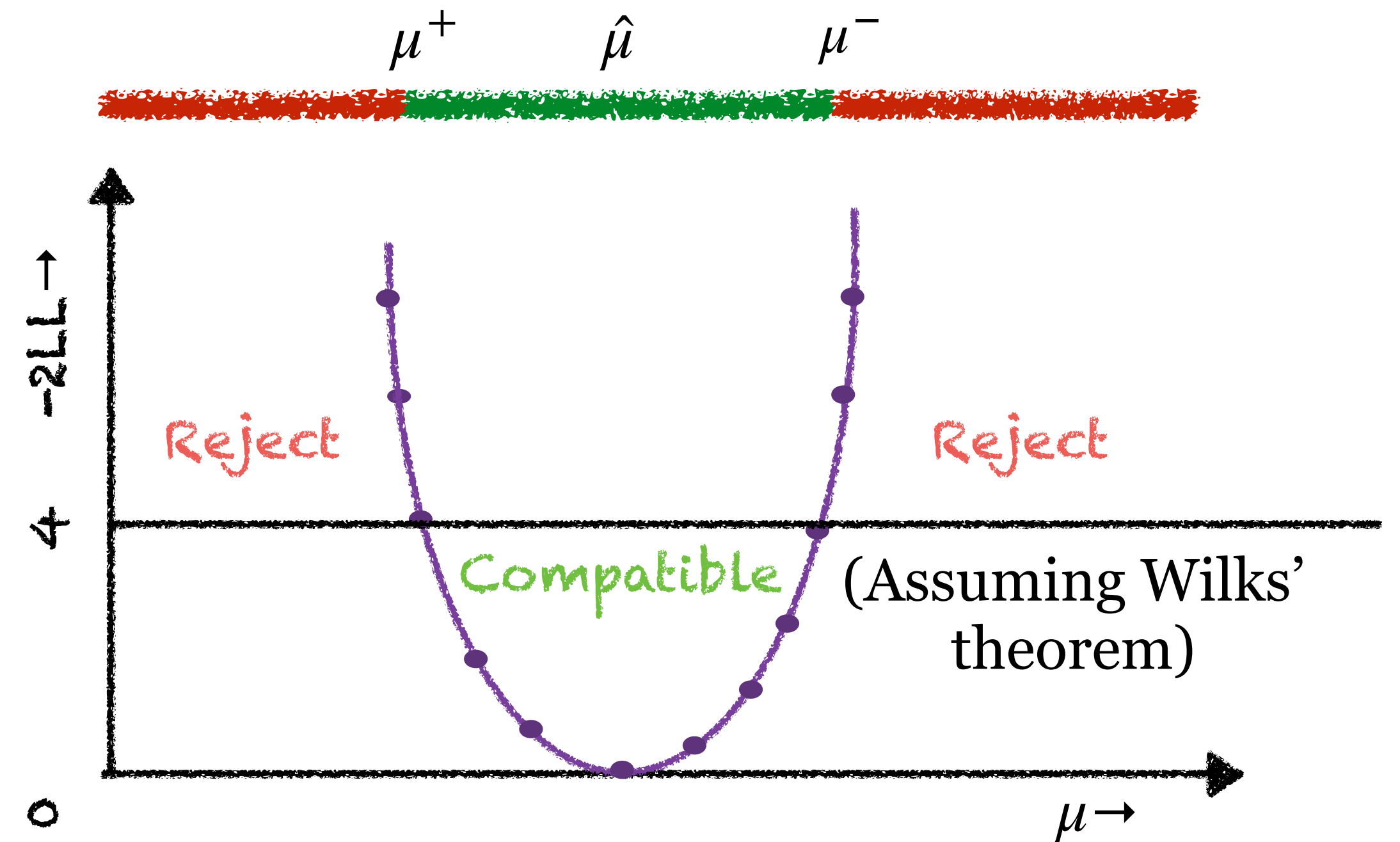
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# “Well justified assumptions”

## Claim

- You can summarise data into one variable
- You can rely on asymptotic approximations
- Impact of nuisance parameters can be factored
- We have thousands of samples, live in asymptopia, and likelihood ratio test is optimal

## Self-fulfilment

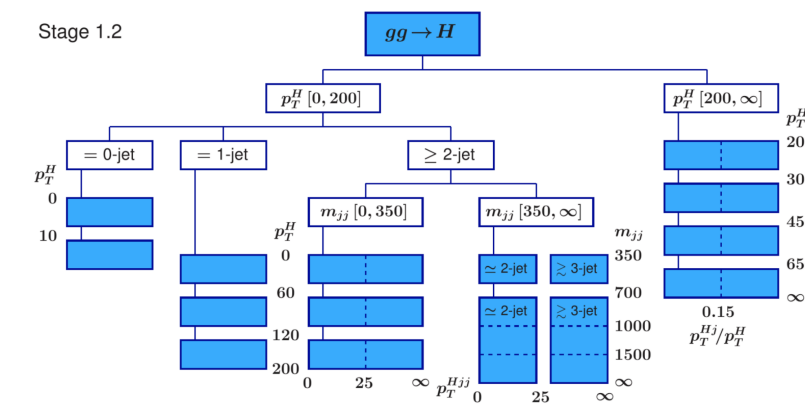
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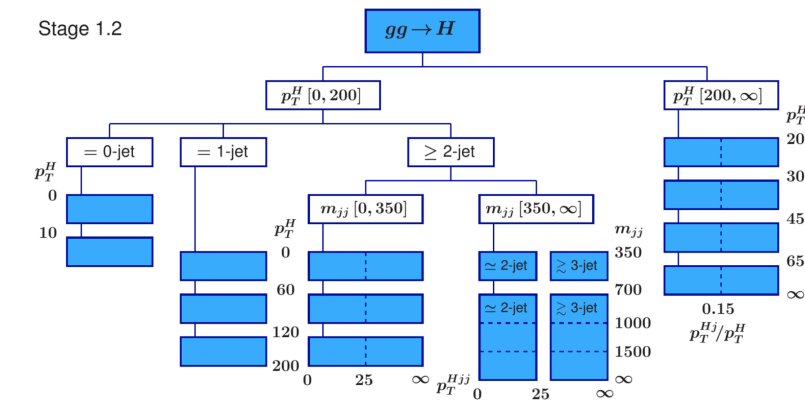
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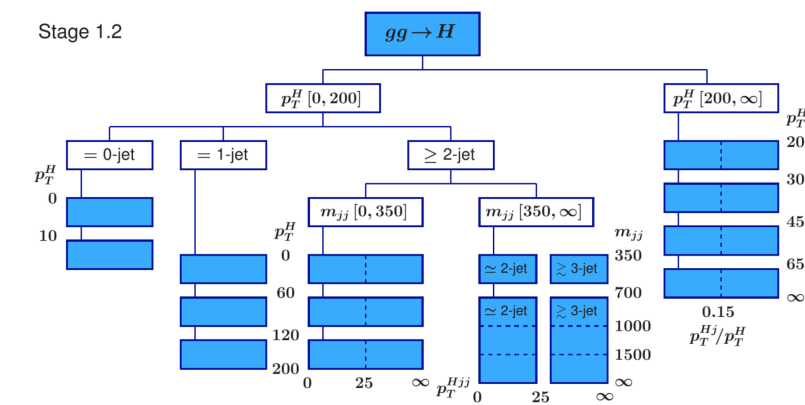
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- We re-defined nuisances in the dirtiest way to satisfy this claim (sacrificing power, or worse)





## **Open problems to neural SBI application to LHC data:**

- Robustness: Design and validation
- Systematic Uncertainties: Propagate them through network
- Neyman Construction: Generating pseudo-experiments to invert the test

# Open problems to neural SBI application to LHC data:

- Robustness: Design and validation
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→ Addressed with semi-parametric approach and lots of compute!

Reports on Progress in **Physics**

**PAPER • OPEN ACCESS**

An implementation of neural simulation-based inference for parameter estimation

To cite this article: The ATLAS Collaboration 2025 *Rep. Prog. Phys.* **88** 057803

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ATLAS collaboration (incl. Ghosh)

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**PAPER • OPEN ACCESS**

Measurement of off-shell Higgs boson production in the  $H^* \rightarrow ZZ \rightarrow 4\ell$  decay channel using a neural simulation-based inference technique in 13 TeV  $pp$  collisions with the ATLAS detector

To cite this article: The ATLAS Collaboration 2025 *Rep. Prog. Phys.* **88** 057803

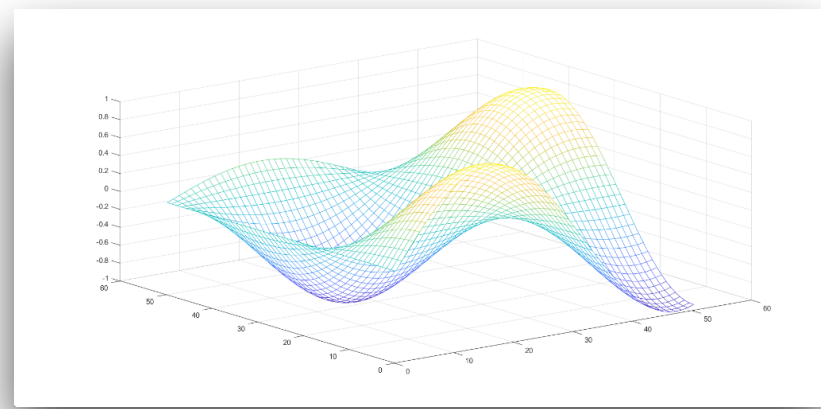
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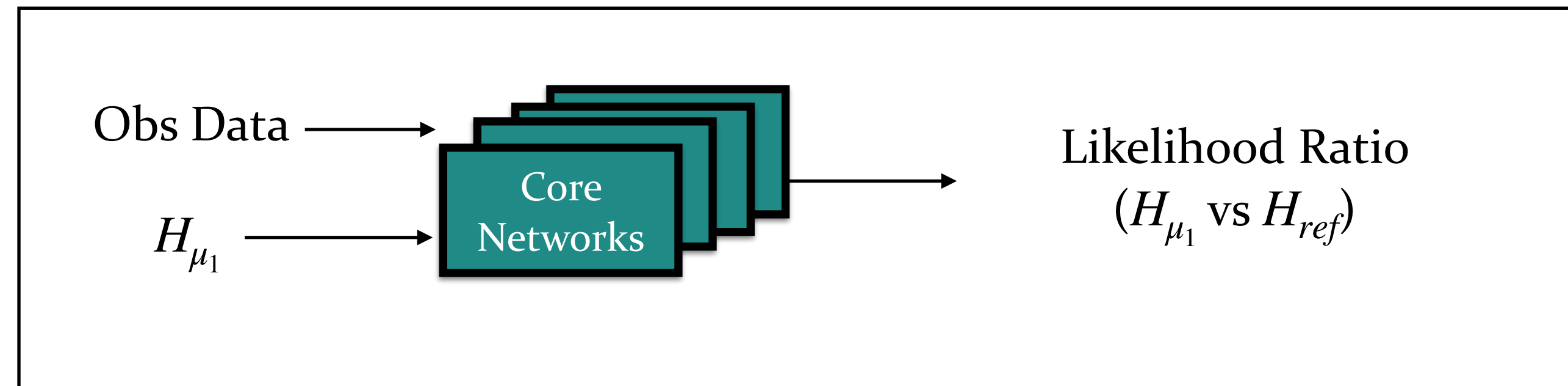
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- [70 years of hyperon spectroscopy: a review of strange, charmed and bottom baryons](#)  
Volker Crede and John Yelton

And took 6 years of my own life to make it a reality..

# Semi-parametric NSBI implementation

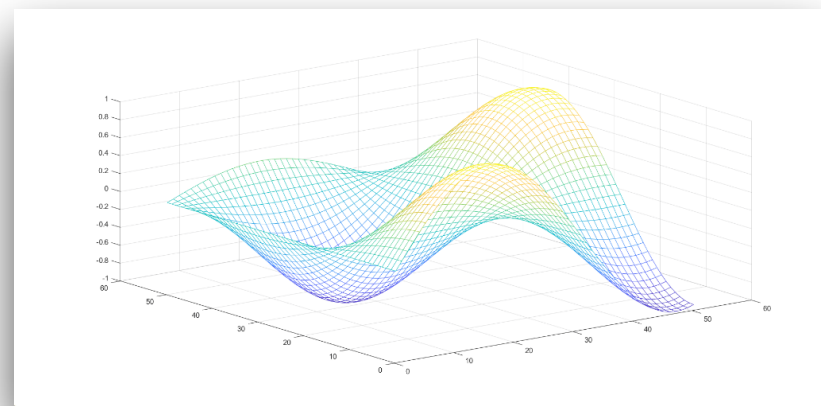


$O(16)$  observables

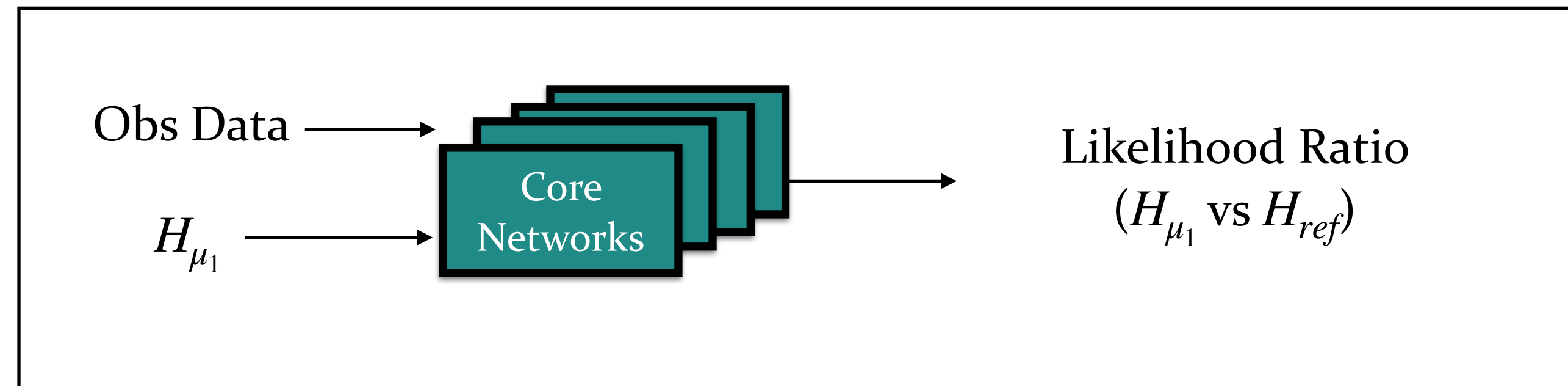


$$\frac{p(x_i | \mu, \alpha)}{p_{ref}(x_i)} = \frac{1}{\nu(\mu, \alpha)} \sum_j^C f_j(\mu) \cdot \nu_j \cdot \frac{p_j(x_i)}{p_{ref}(x_i)} \cdot \prod_k^{N_{syst}} G_j(\alpha_k) \cdot \frac{p_j(x_i, \alpha_k)}{p_j(x_i)}$$

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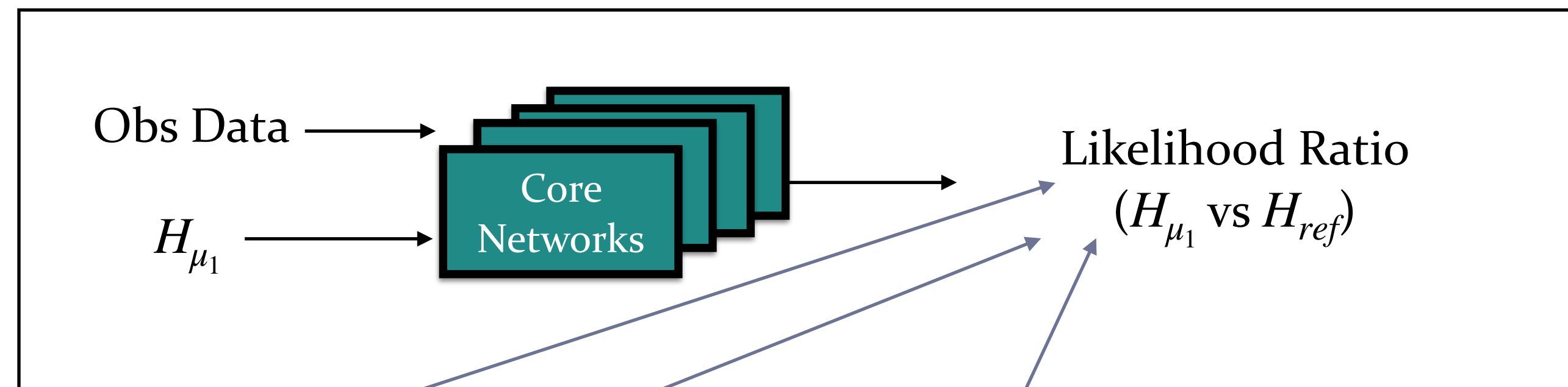


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# Semi-parametric NSBI implementation



Syst\_o  
Network

Syst\_1  
Network

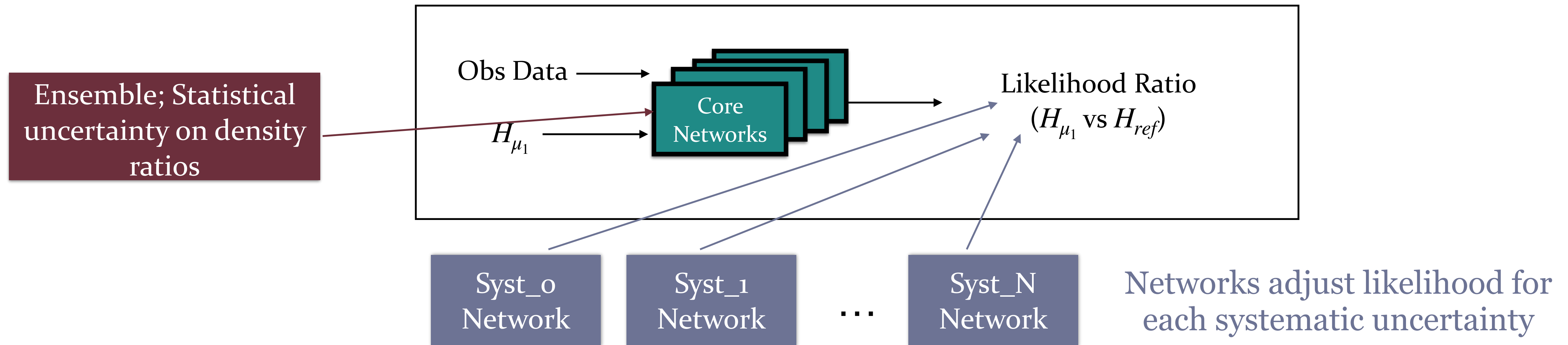
...

Syst\_N  
Network

Networks adjust likelihood for each systematic uncertainty

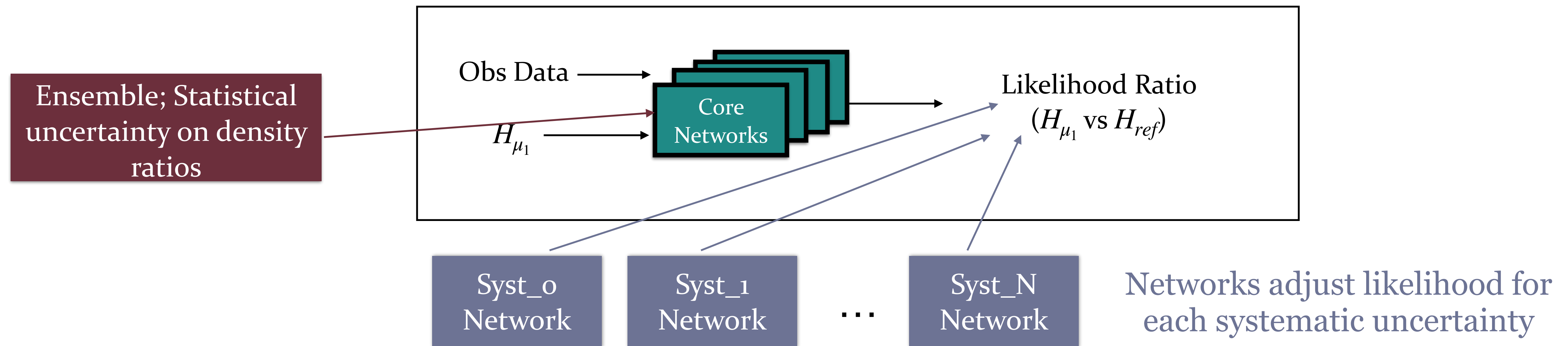
$$\frac{p(x_i | \mu, \alpha)}{p_{ref}(x_i)} = \frac{1}{\nu(\mu, \alpha)} \sum_j^C f_j(\mu) \cdot \nu_j \cdot \frac{p_j(x_i)}{p_{ref}(x_i)} \cdot \prod_k^{N_{syst}} G_j(\alpha_k) \cdot \frac{p_j(x_i, \alpha_k)}{p_j(x_i)}$$

# Semi-parametric NSBI implementation



$$\frac{p(x_i | \mu, \alpha)}{p_{ref}(x_i)} = \frac{1}{\nu(\mu, \alpha)} \sum_j^C f_j(\mu) \cdot \nu_j \cdot \frac{p_j(x_i)}{p_{ref}(x_i)} \cdot \prod_k^{N_{syst}} G_j(\alpha_k) \cdot \frac{p_j(x_i, \alpha_k)}{p_j(x_i)}$$

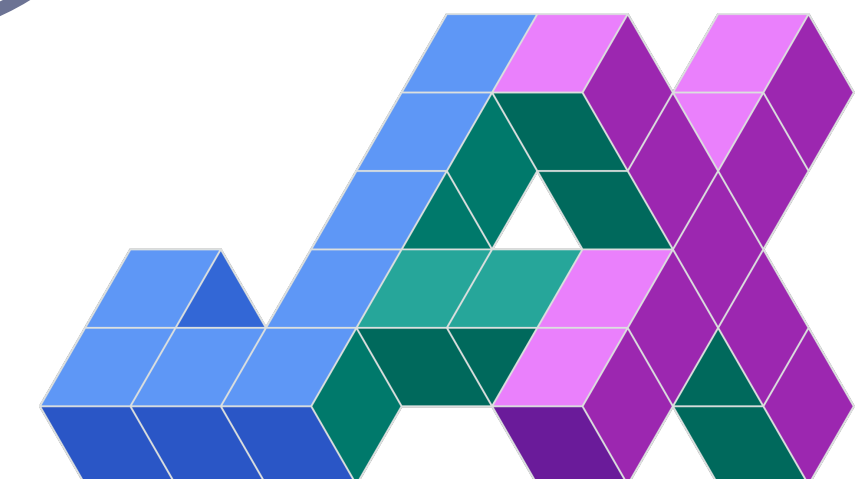
# Semi-parametric NSBI implementation



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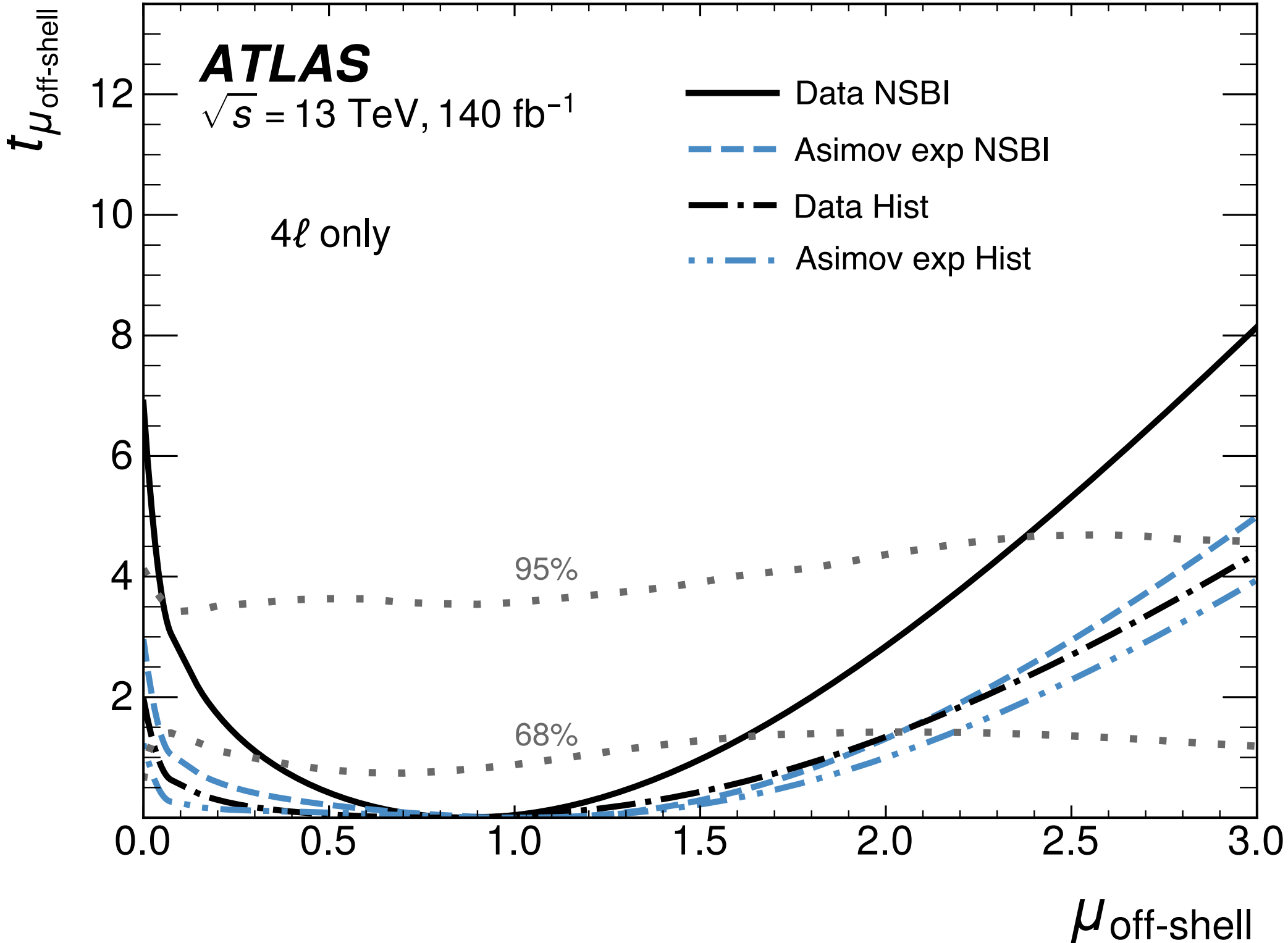
Training details

- ◆ Train  $O(10^4)$  networks on TensorFlow
- ◆ Computing resources provided by Google, SMU, other HPC clusters
- ◆ Fits with JAX



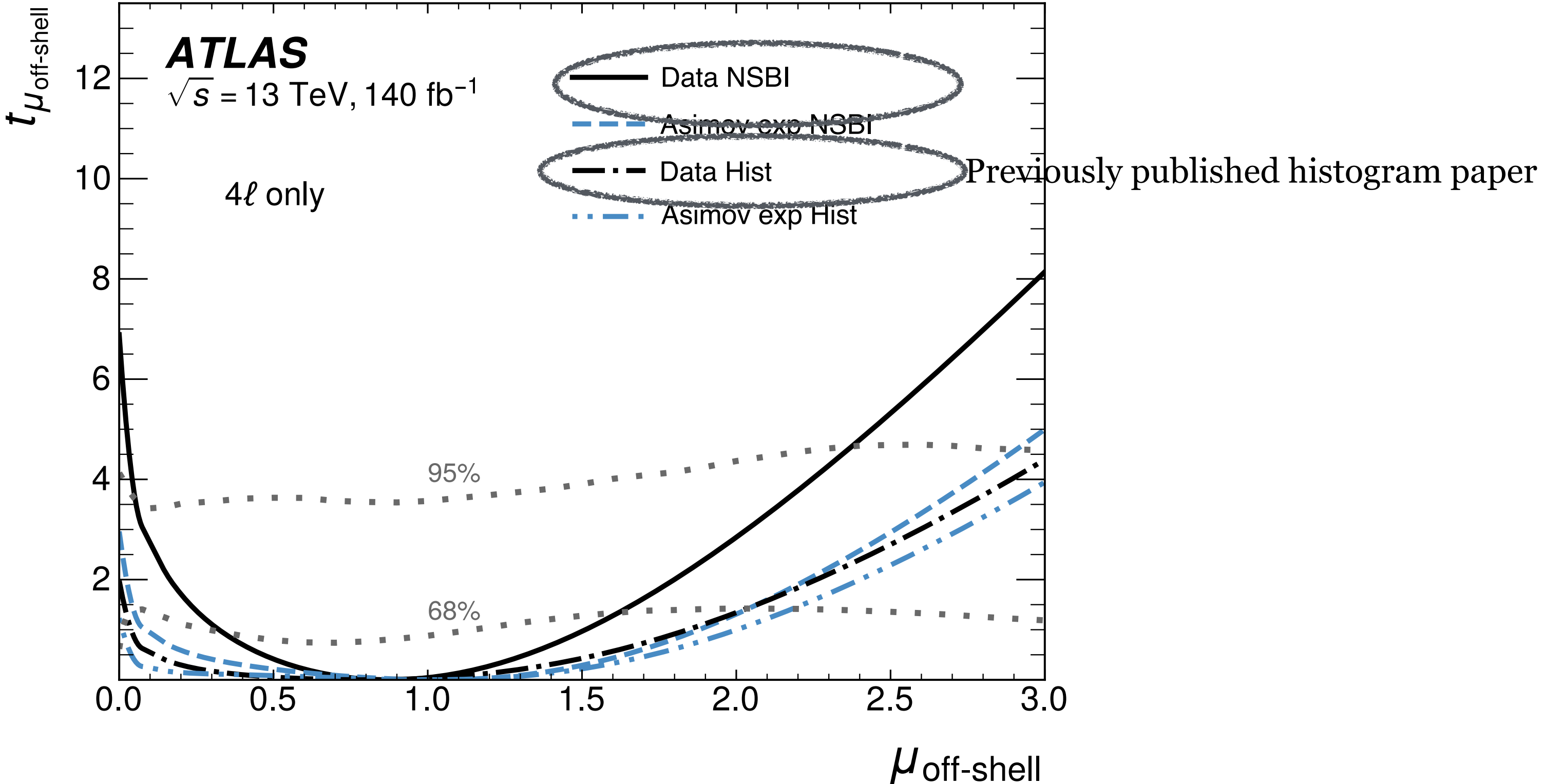
# Application to a flagship Higgs measurement at LHC

## NSBI vs (published) histogram analysis



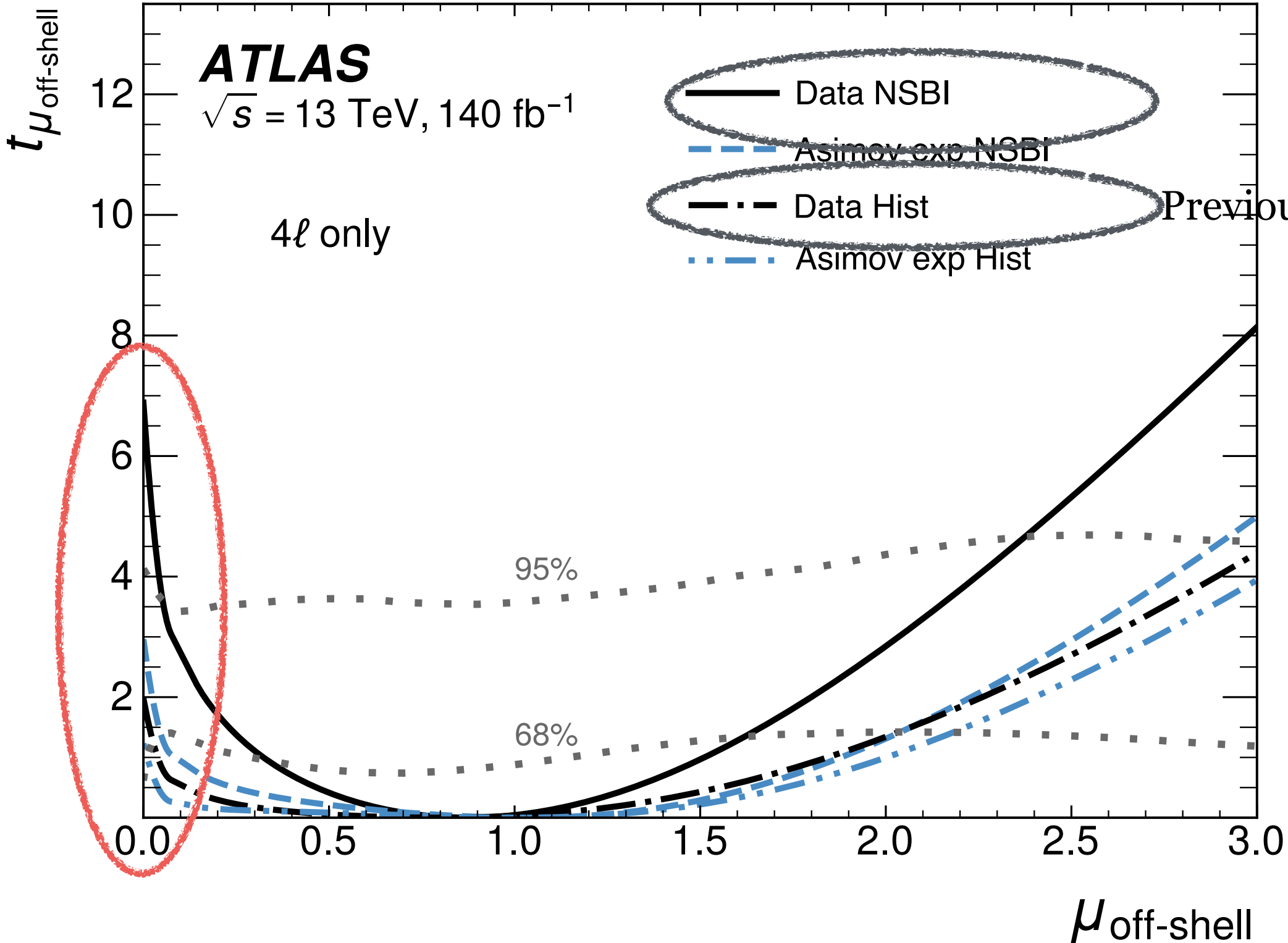
# Application to a flagship Higgs measurement at LHC

## NSBI vs (published) histogram analysis



# Application to a flagship Higgs measurement at LHC

## NSBI vs (published) histogram analysis



Previously published histogram paper

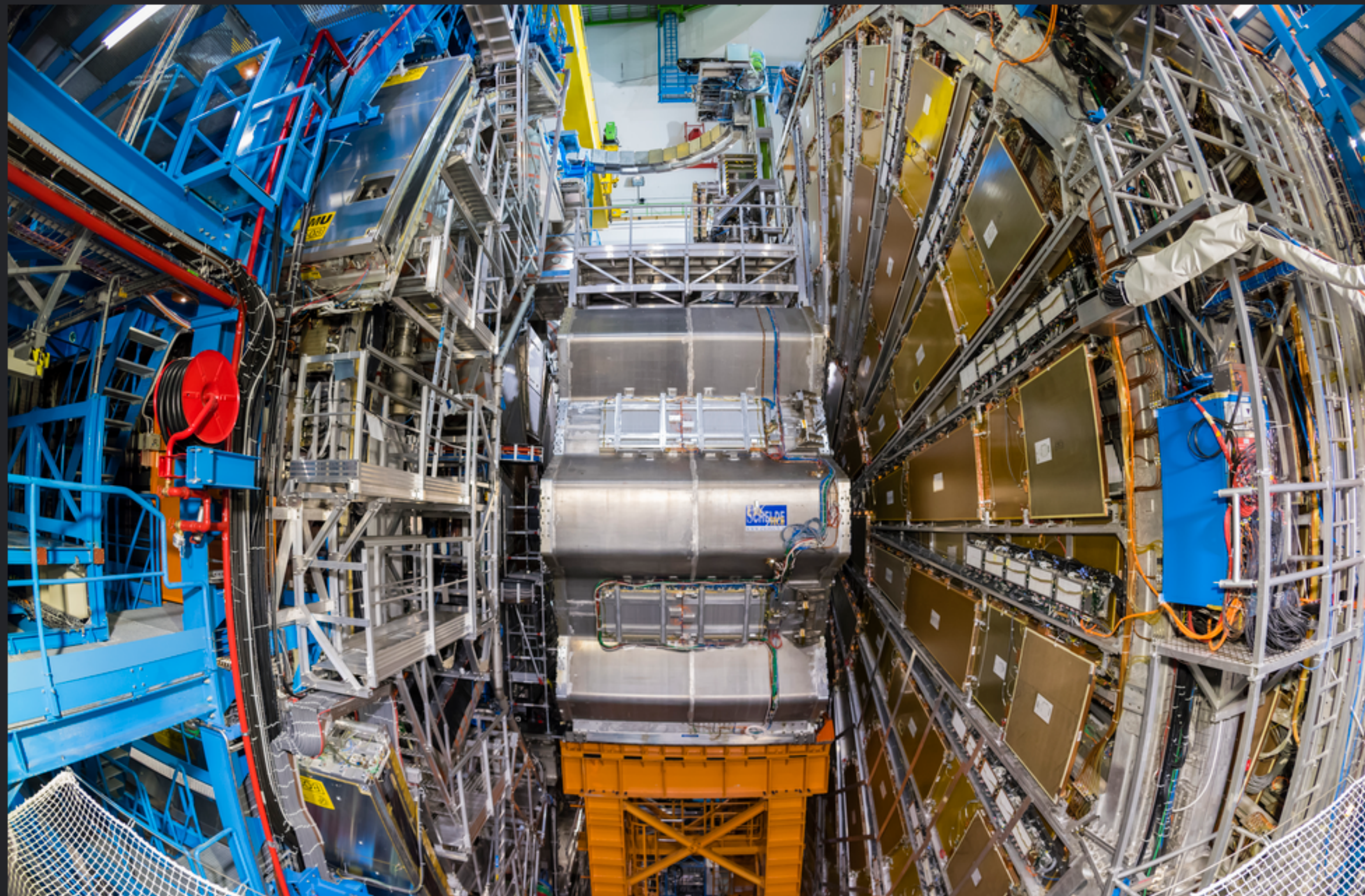
Unprecedented improvement in ability to reject null hypothesis!

A VERY FUN PROBLEM

# How a grad student got LHC data to play nice with quantum interference

New approach is already having an impact on the experiment's plans for future work.

MATT VON HIPPEL - 23 JUN 2025 11:00 70



→ The ATLAS particle detector of the Large Hadron Collider (LHC) at the European Nuclear Research Center (CERN) in Geneva, Switzerland. Credit:

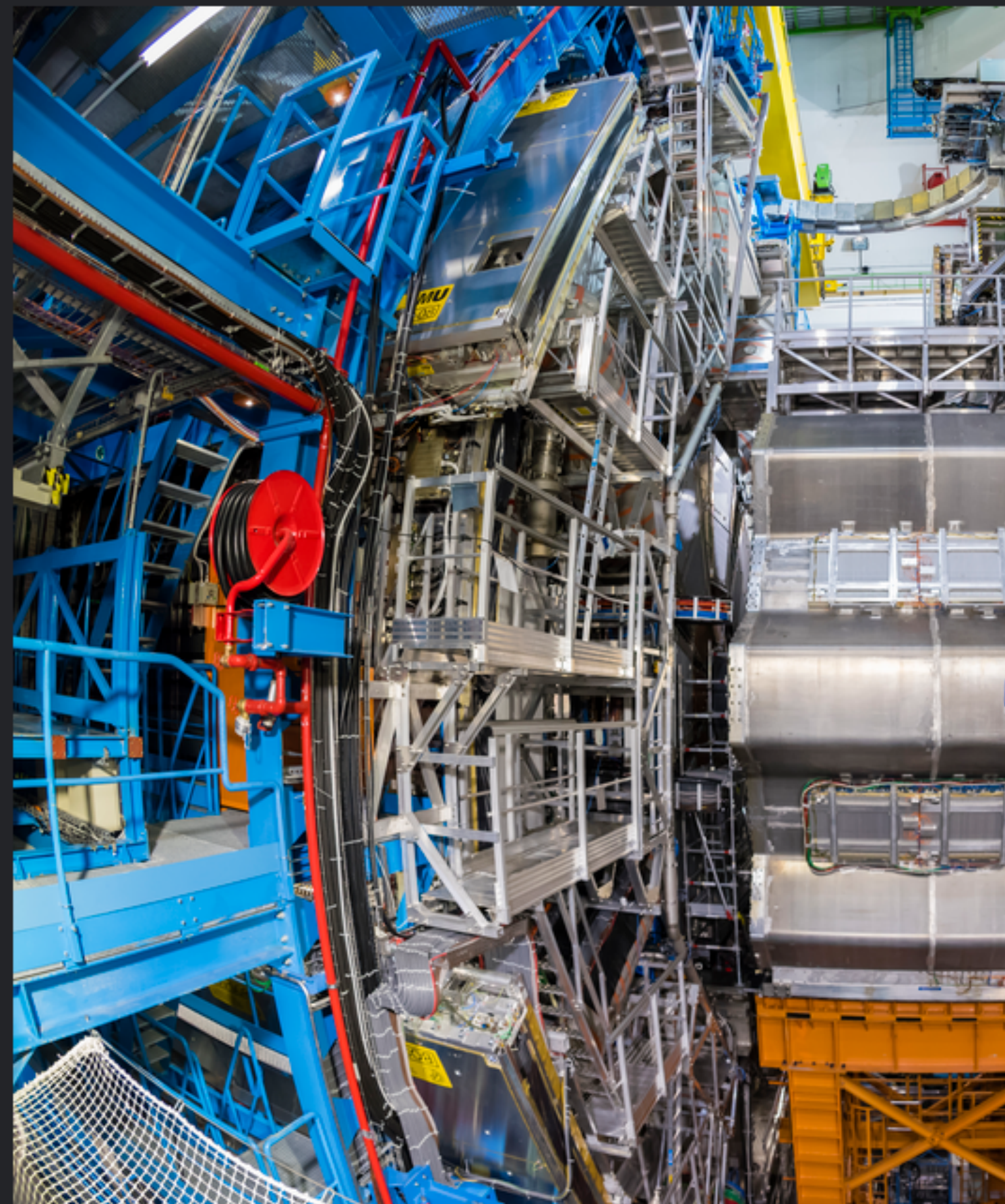
EThamPhoto/Getty Images

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EThamPhoto/Getty Images



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Updates > Briefing > Cracking open the Higgs shell: new ATLAS measurement of “off-shell” production uses AI techniques

## Physics Briefing

Tags:  
Higgs boson,  
physics results

# Cracking open the Higgs shell: new ATLAS measurement of “off-shell” production uses AI techniques

6 November 2024 | By [ATLAS Collaboration](#)

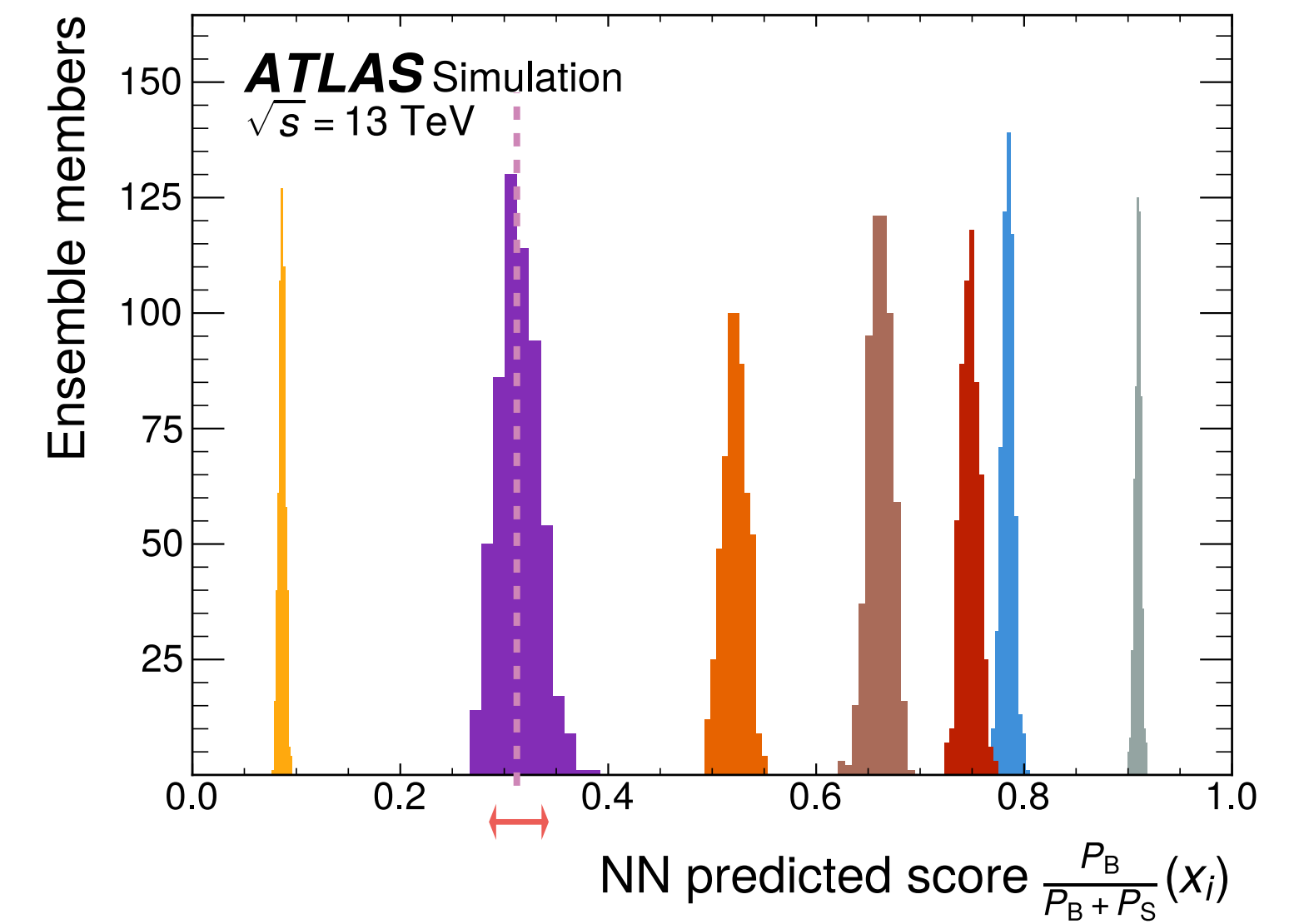
In 2012, scientists from the ATLAS and CMS Collaborations detected a “bump” in their data. This peak in the mass distribution, around 125 GeV, revealed the presence of the Higgs boson. While most Higgs bosons are observed at the LHC with this “on shell” mass, about 15% of Higgs bosons have a “virtual” mass well above 125 GeV – known as the “off shell” mass. This is one of the quirks of quantum mechanics, which allows particles to fluctuate their mass for an extremely short time.

By comparing the production rates of on-shell and off-shell Higgs bosons, physicists can study key properties of the Higgs boson, such as the width of the peak in the mass distribution. This width is related to the Higgs boson’s lifetime and its decay rate to other particles. As it is predicted to be just 4 MeV – more than a hundred times smaller than the resolution of the ATLAS detector – this is the only way to set constraints on its value at the LHC. Additionally, the rate of off-shell Higgs-boson production is sensitive to potential contributions from Beyond the Standard Model (BSM) physics at high energies – making it an important tool for exploring new physics.

As off-shell production doesn’t result in the characteristic “bump” of on-shell production in the mass spectrum, it can be difficult to distinguish from background processes with identical signatures and larger rates. To overcome these challenges, the [initial ATLAS measurement](#) of the Higgs-boson width with LHC Run-2 data (collected in 2015-2018) used a machine-learned neural network to distinguish signal events from background. This approach, however, is not optimal for a measurement of off-shell Higgs production due to quantum interference between the signal and background.

# Uncertainties on the network estimates themselves

Distribution of NN predictions for example events



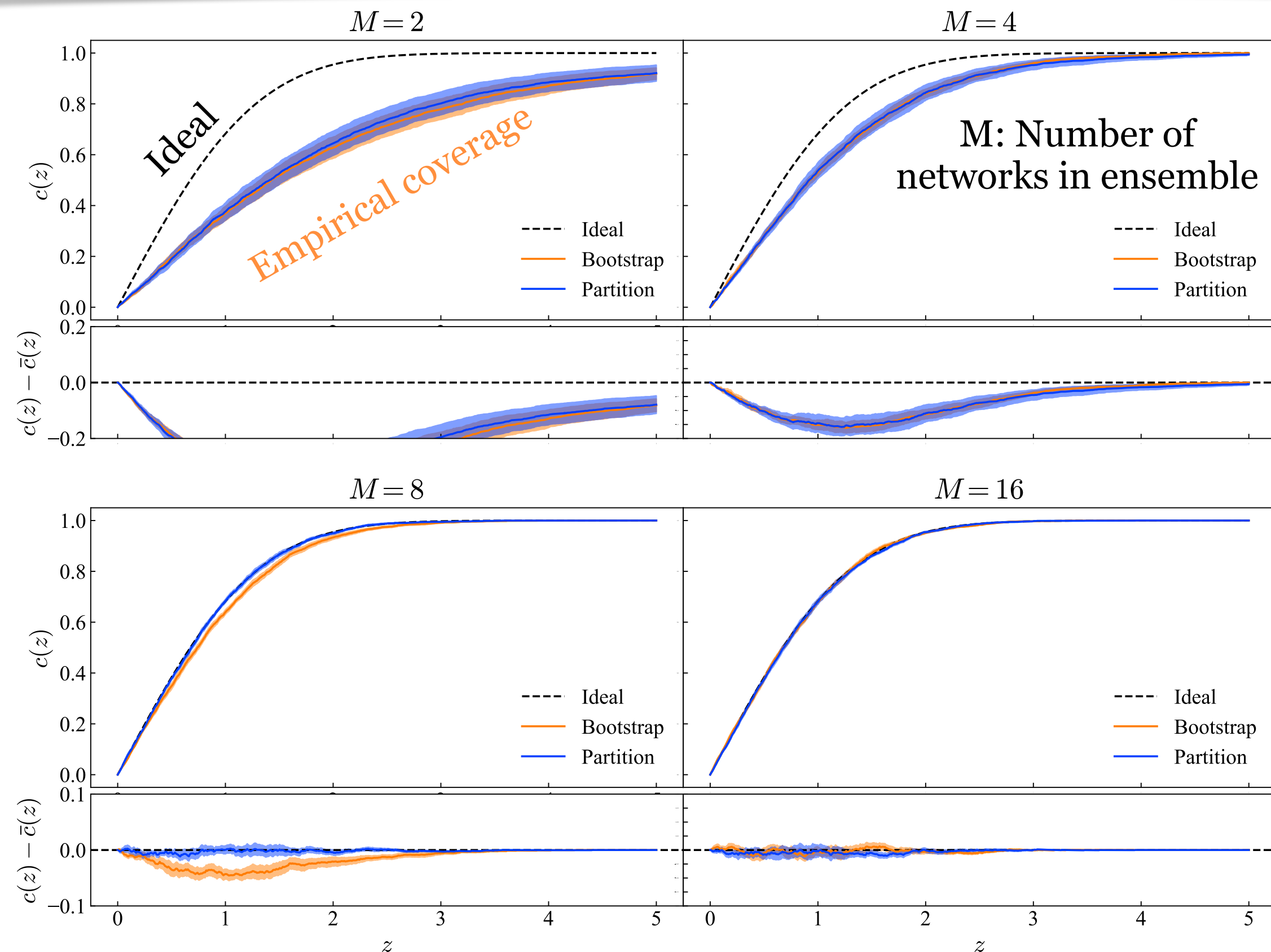
# Uncertainties on the network estimates themselves

## Frequentist Uncertainties on Neural Density Ratios with $w_i f_i$ Ensembles

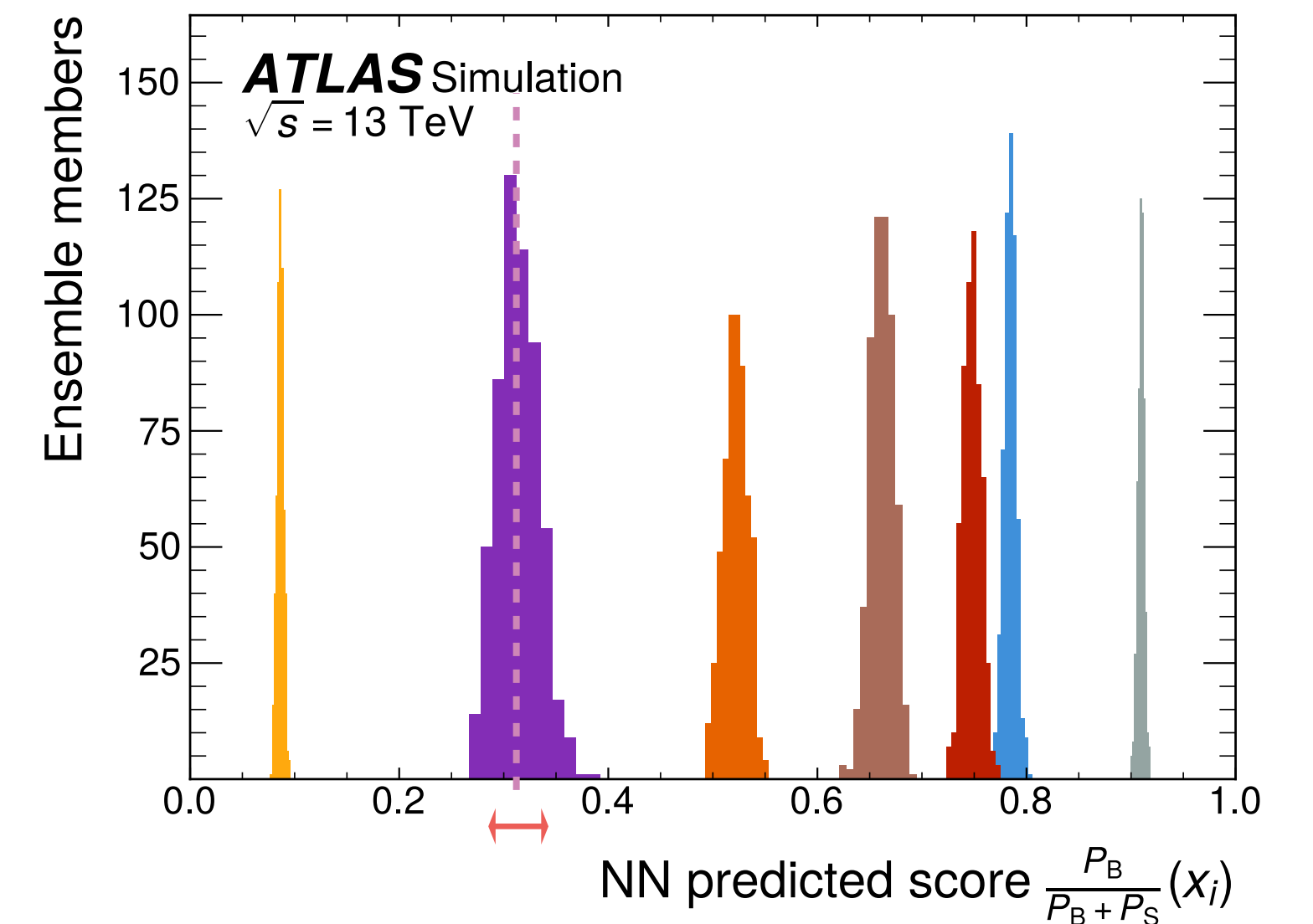
Sean Benevedes<sup>1,2,\*</sup> and Jesse Thaler<sup>1,2,†</sup>

<sup>1</sup>Center for Theoretical Physics, Massachusetts Institute of Technology,  
Cambridge, Massachusetts, United States

<sup>2</sup>The NSF AI Institute for Artificial Intelligence and Fundamental Interactions



Distribution of NN predictions for example events

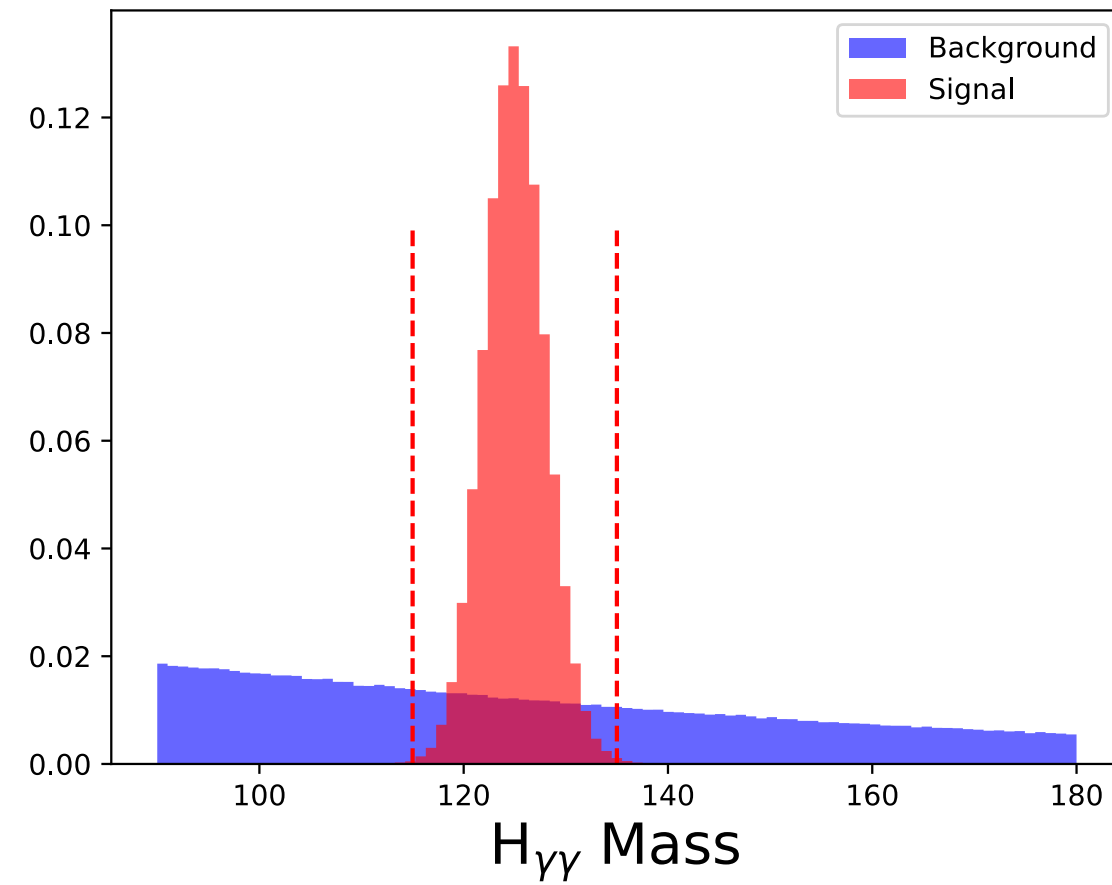


- Genuine 95% coverage, with much smaller ensembles
- Elegant, mathematically motivated method to estimate uncertainties

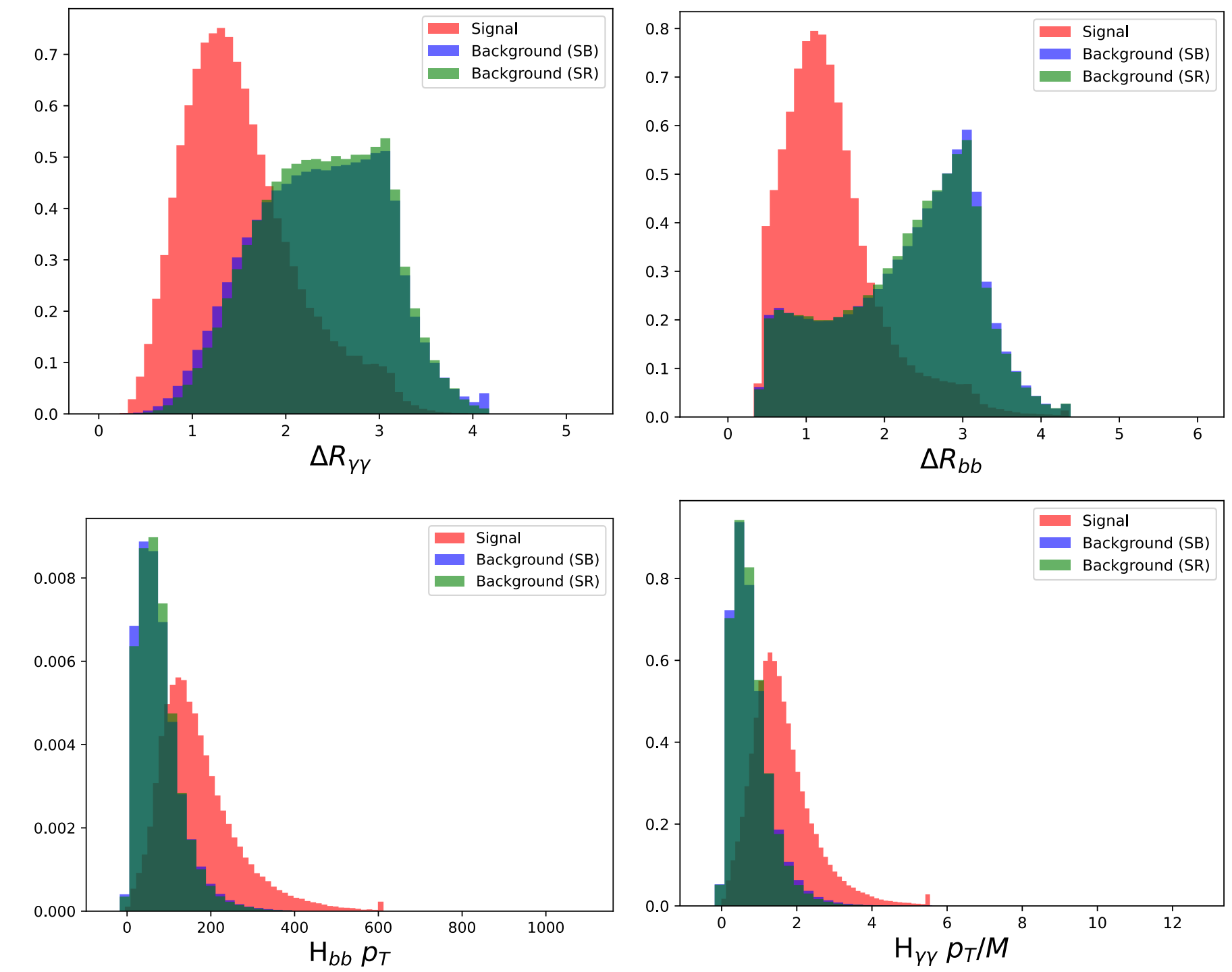
[arXiv:2506.00113](https://arxiv.org/abs/2506.00113): Benevedes & Thaler



# Unreliable simulator ?



- Training data can come from **control regions of real data**
- Data-driven background estimation techniques work with NSBI
- Amram & Szwec uses generative models instead of classifiers for NSBI

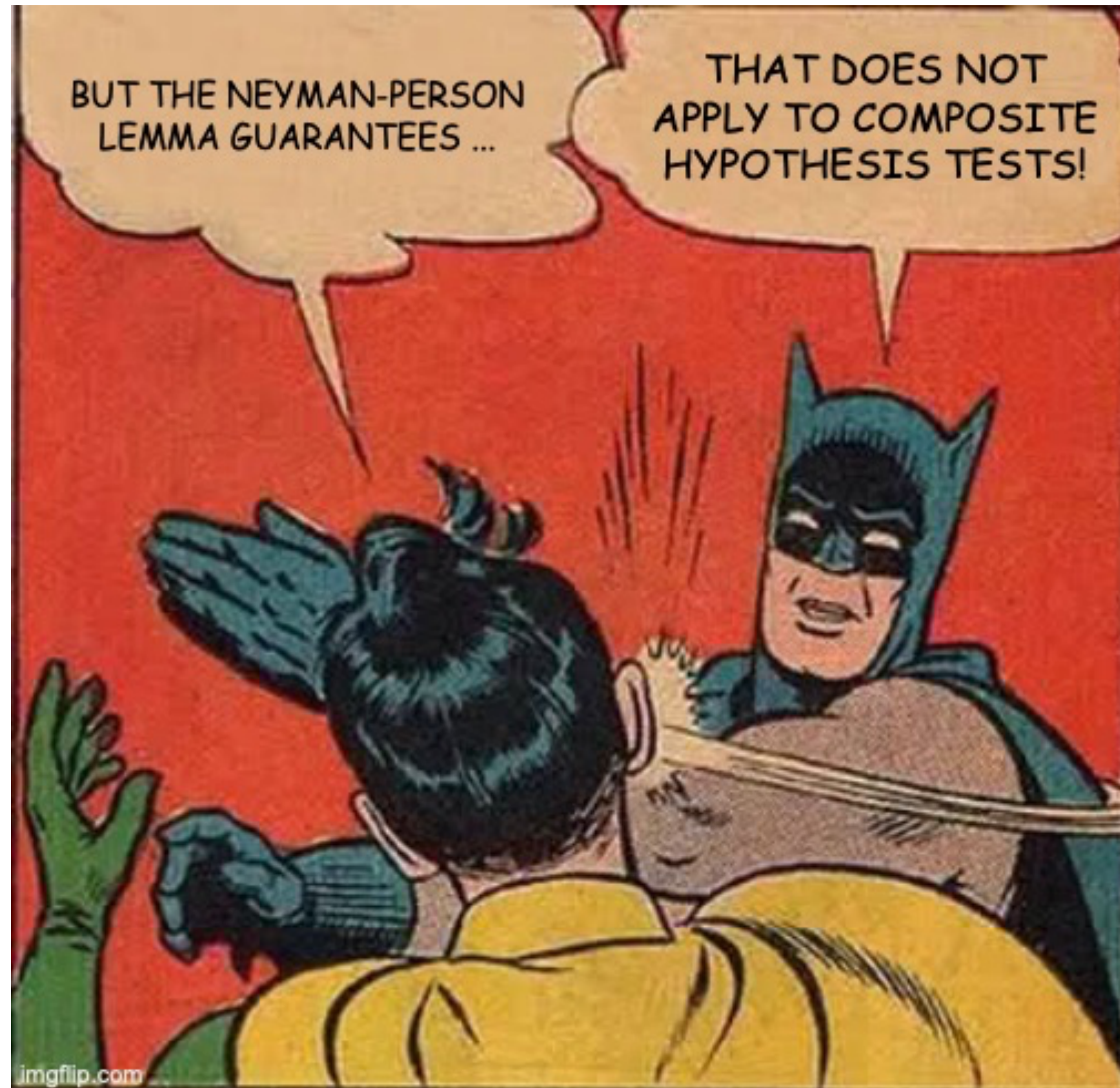


Until now, we have replaced individual pieces with ML in age-old likelihood ratio test

Do we dare question the test itself?

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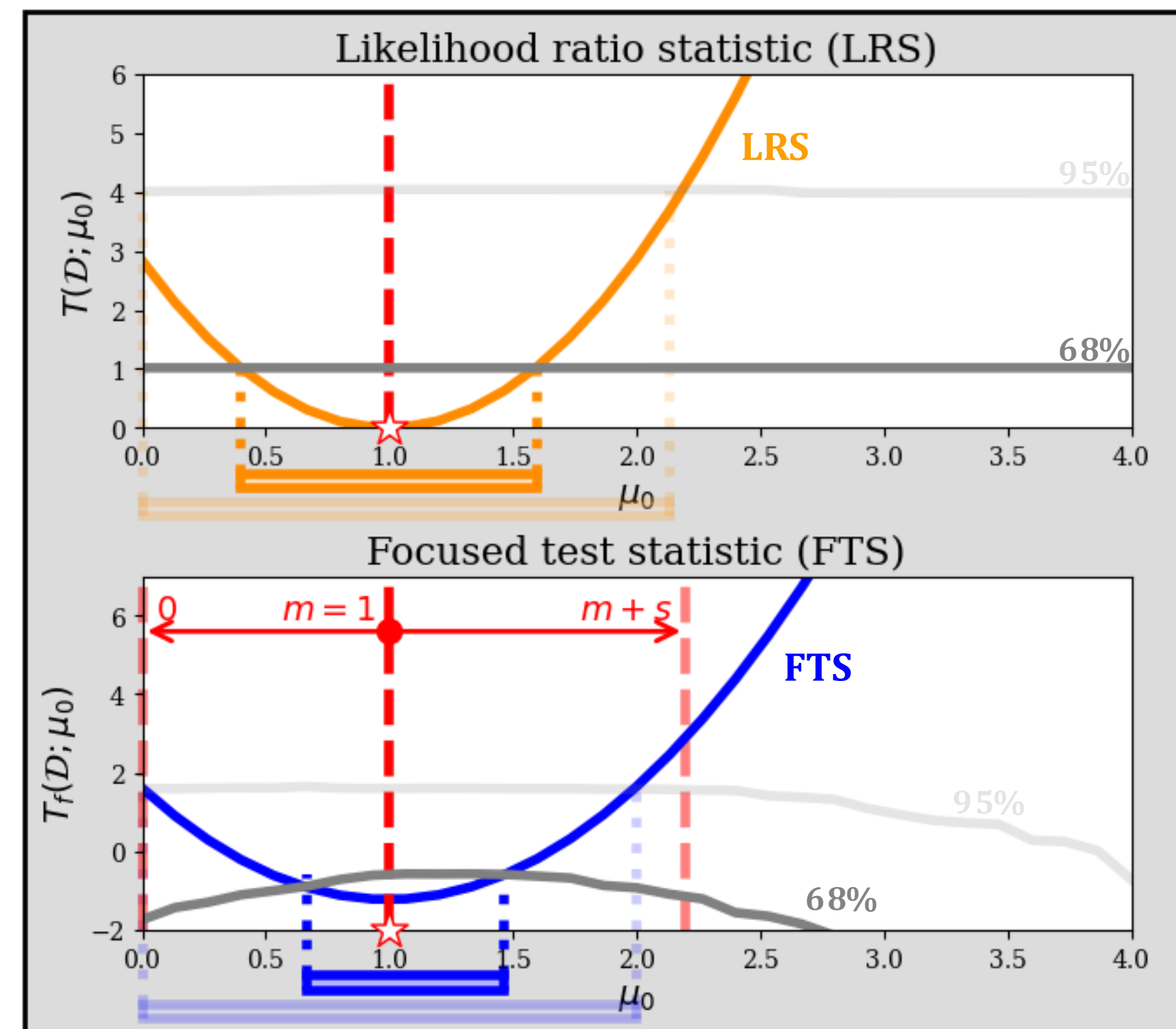
Do we dare question the test itself?



# Challenging a deeply-held belief in particle physics

$$LRS(\mathcal{D}; \mu_0) = -2 \log \left( \frac{p(\mathcal{D} | \mu_0)}{\sup_{\mu \in \Theta} p(\mathcal{D} | \mu)} \right) \longrightarrow FTS(\mathcal{D}; \mu_0) = -2 \log \left( \frac{p(\mathcal{D} | \mu_0)}{\int_{\Theta} p(\mathcal{D} | \mu) f(\mu) d\mu} \right)$$

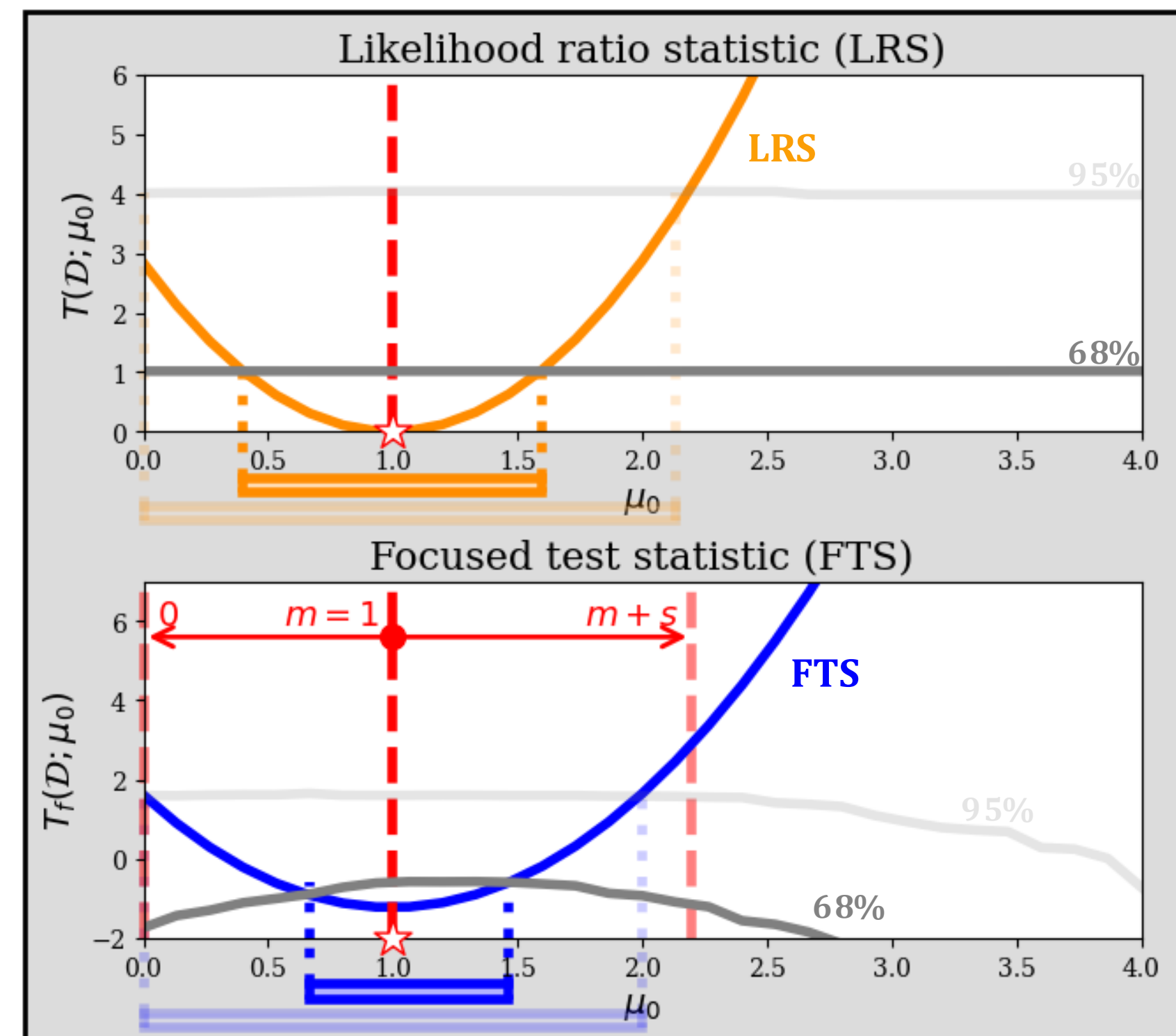
- Denominator in ‘Focused Test Statistic’ (FTS) knows about all alternate hypotheses



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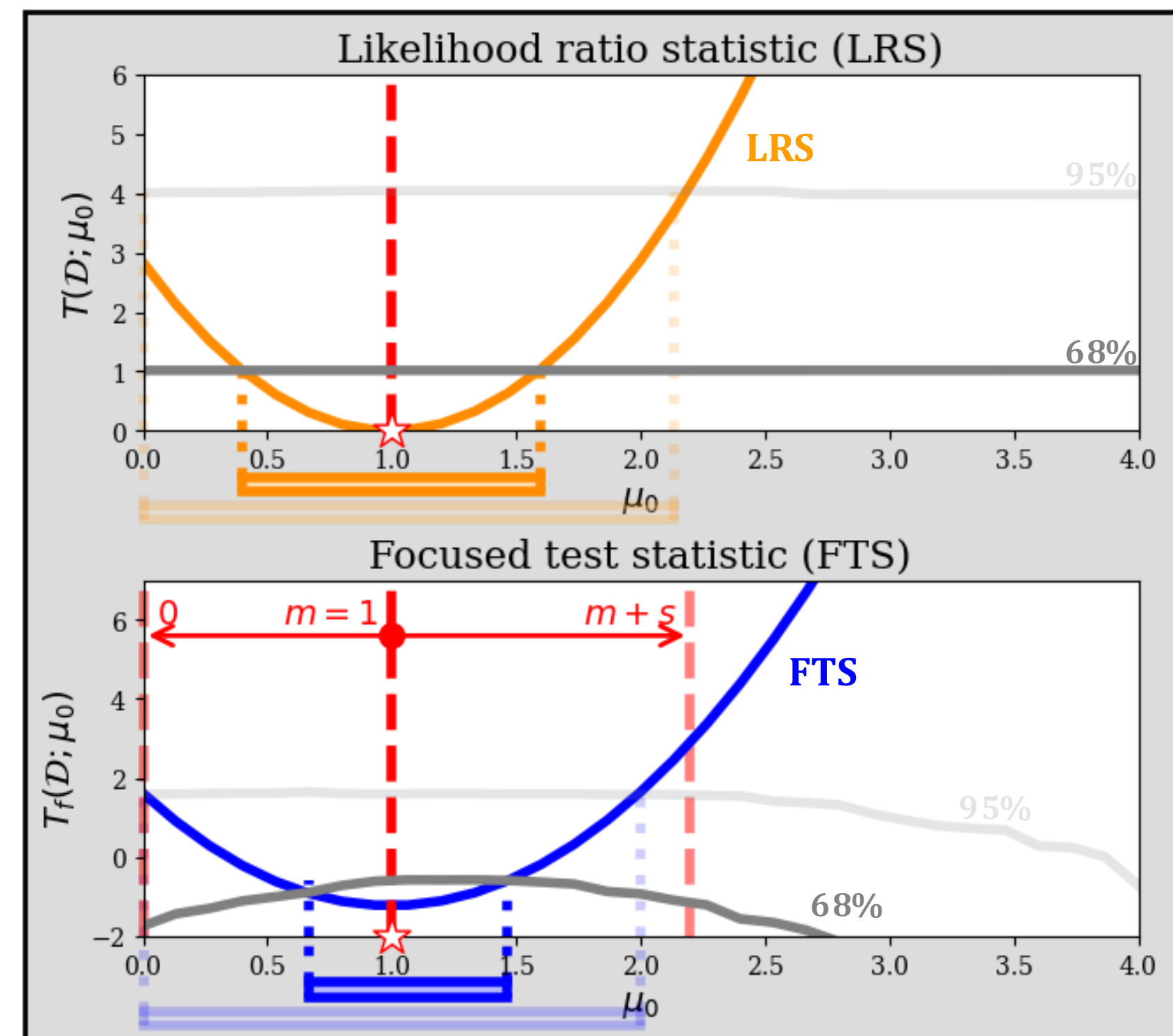


Shorter median length for confidence intervals with FTS even in ‘asymptotic regime’ where Wilks’ applies

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- Denominator in ‘Focused Test Statistic’ (FTS) knows about all alternate hypotheses
- $f(\mu)$  focuses statistical power in meaningful regions of parameter space
  - Particularly useful in small sample / small signal regime
- Fast critical value estimation with ML

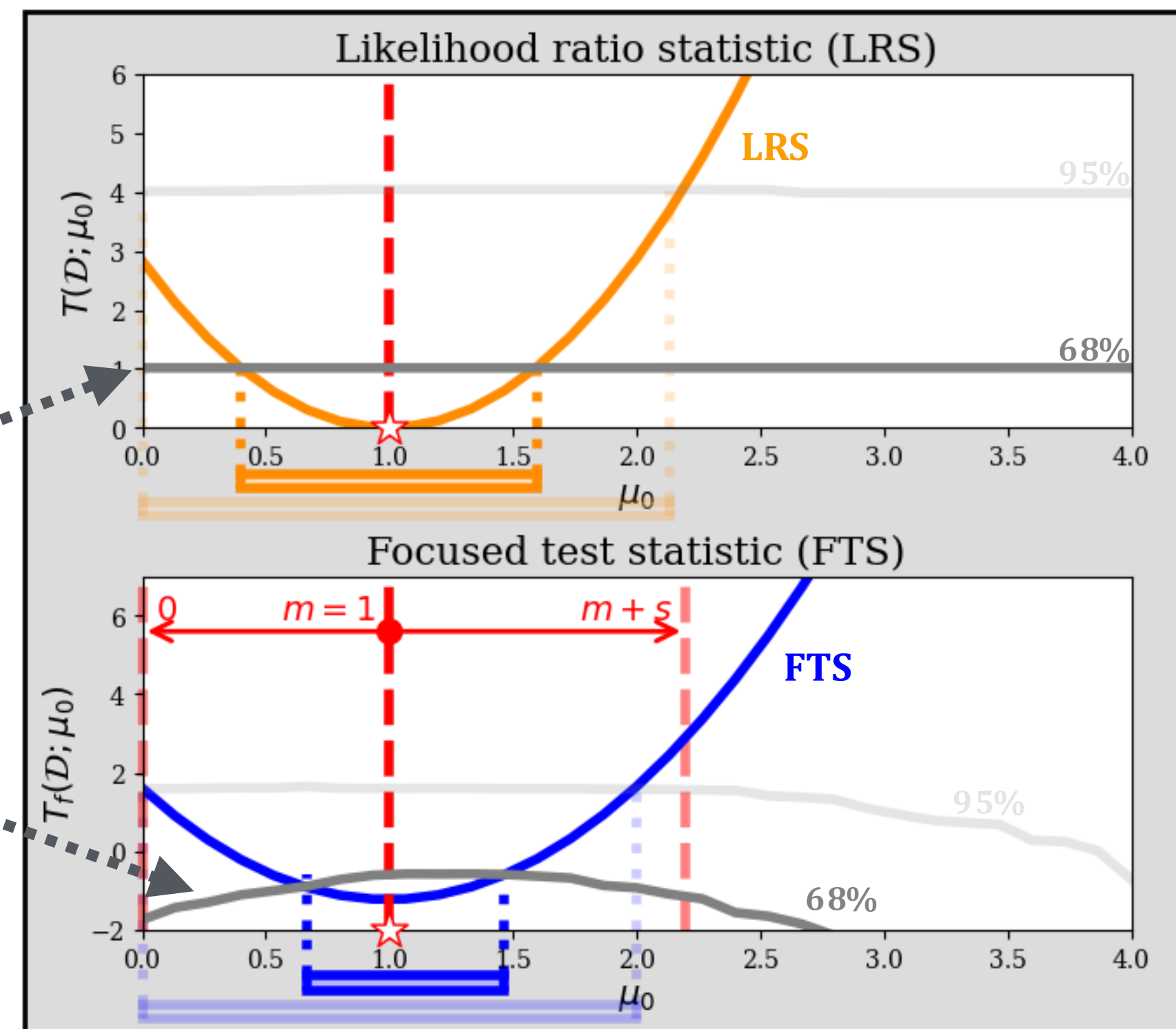


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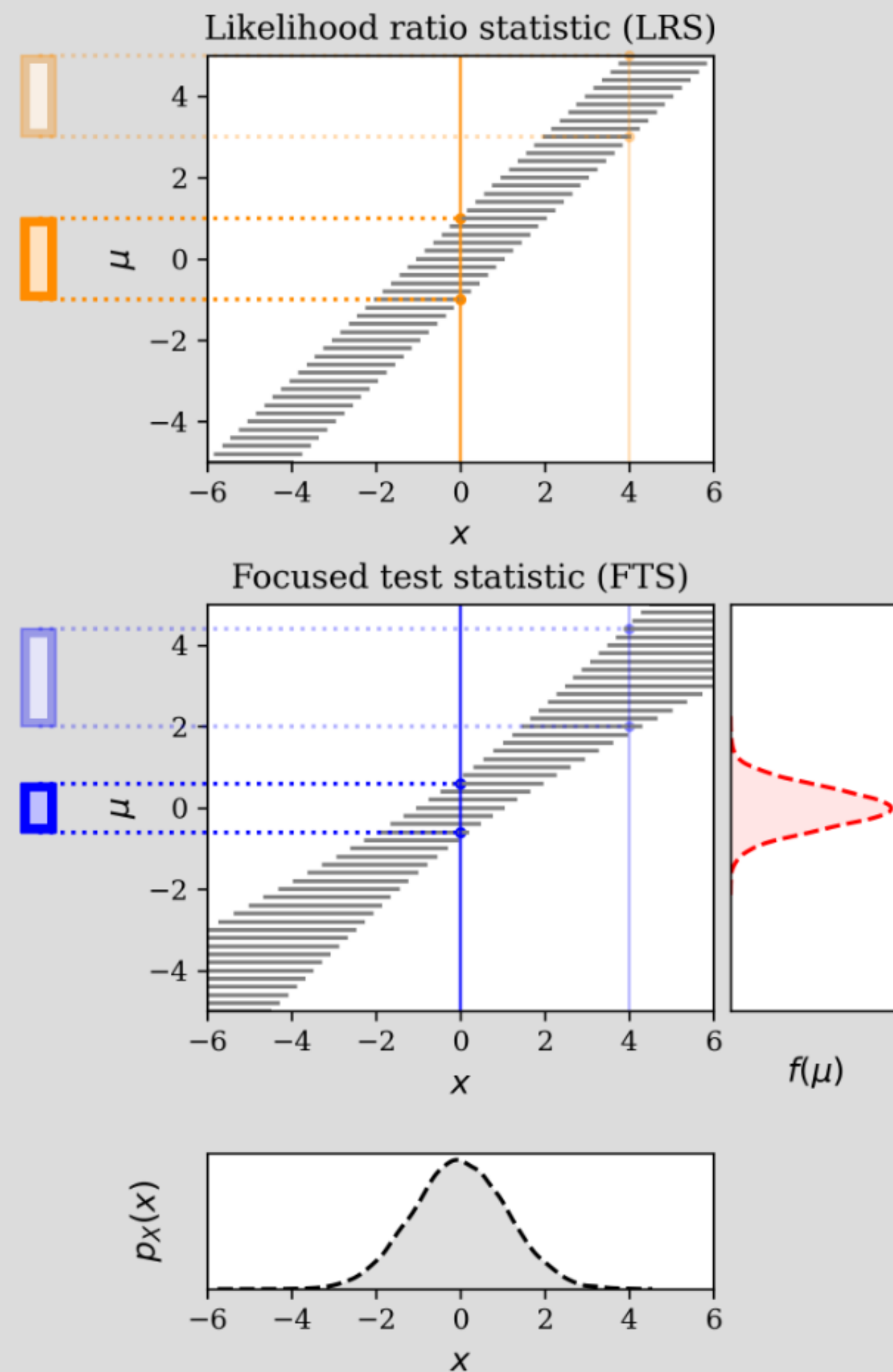
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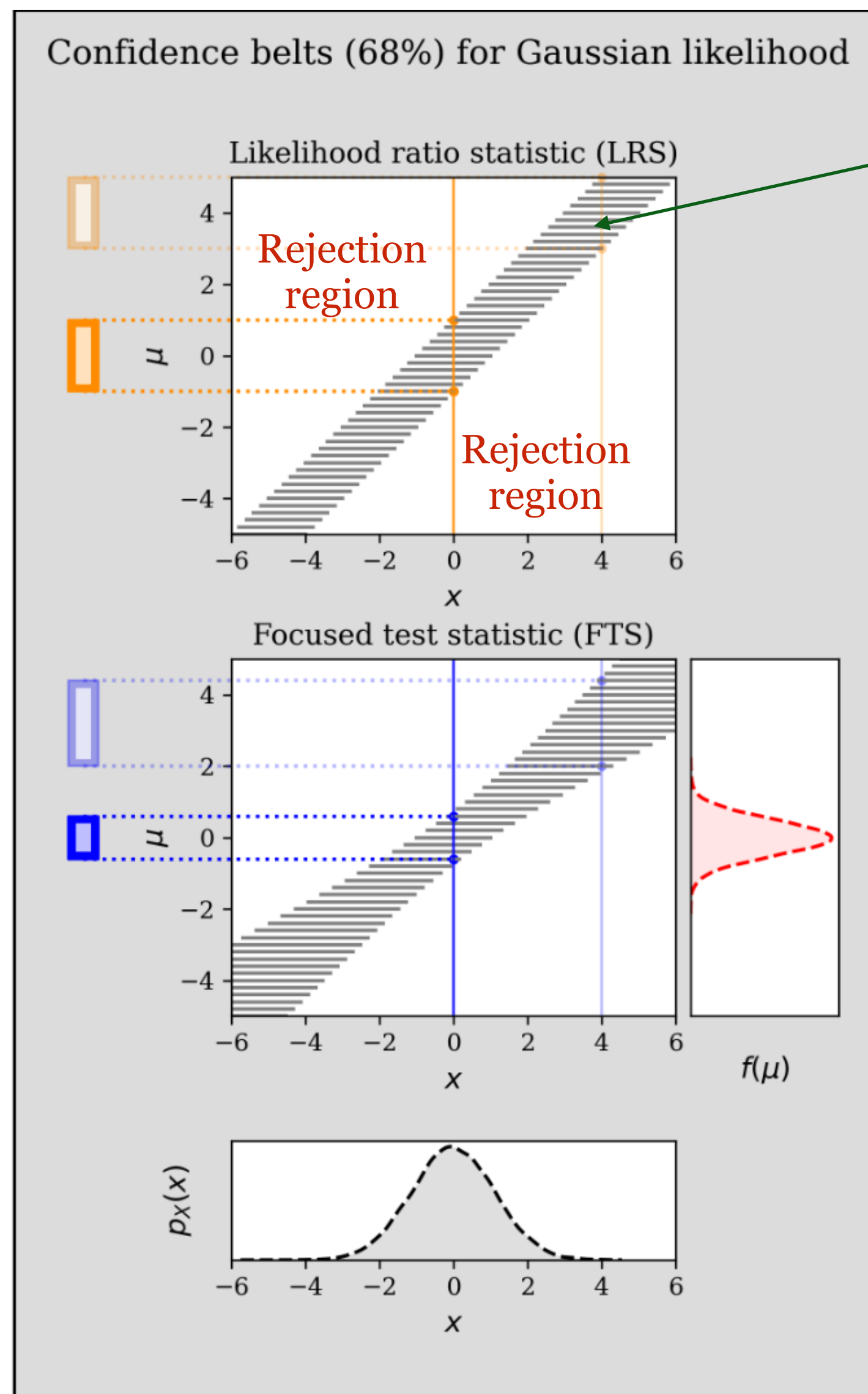
# Test inversion for simplest measurement

Confidence belts (68%) for Gaussian likelihood



Data comprises single measurement  $x$

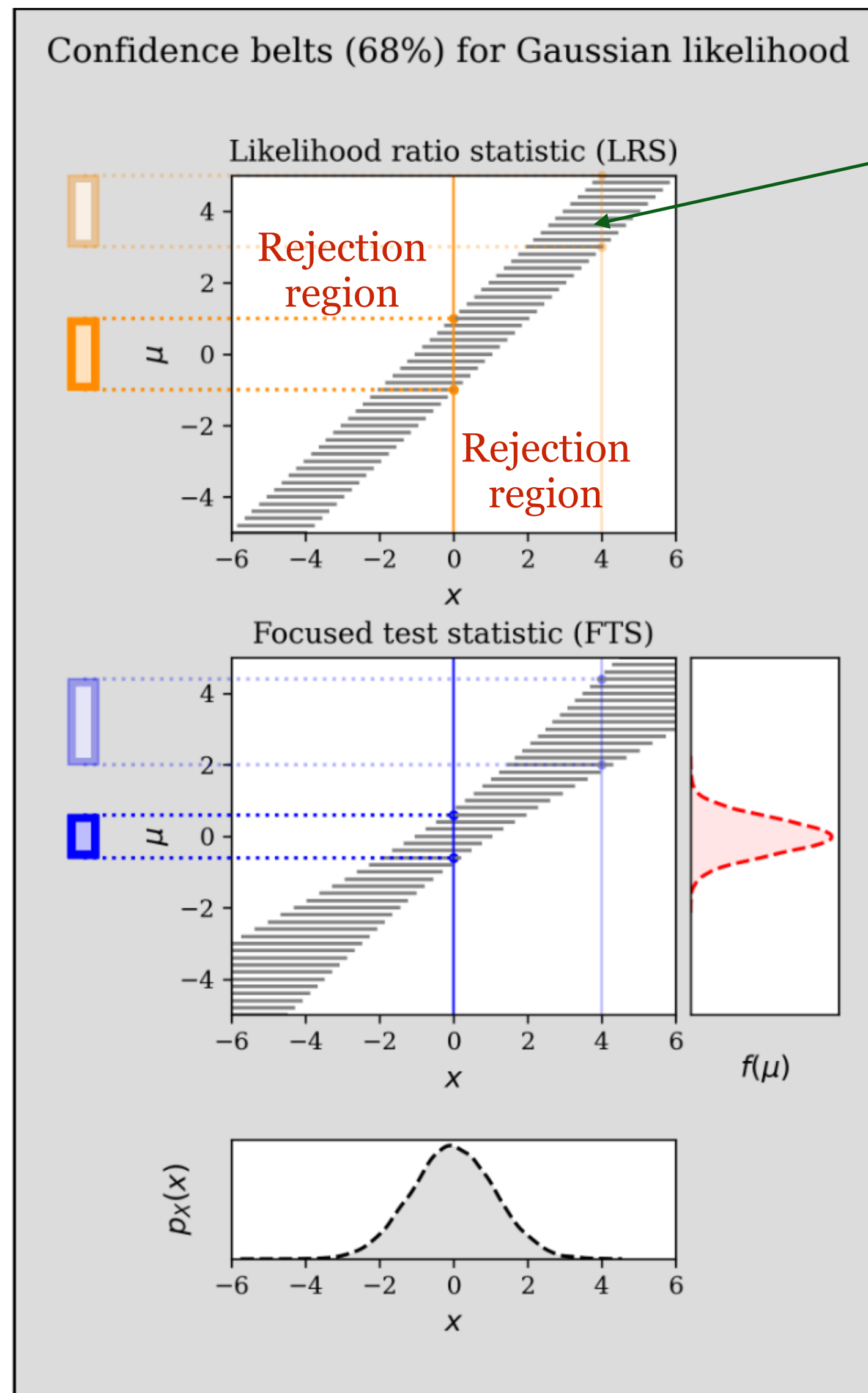
# Test inversion for simplest measurement



Acceptance region

Data comprises single measurement  $x$

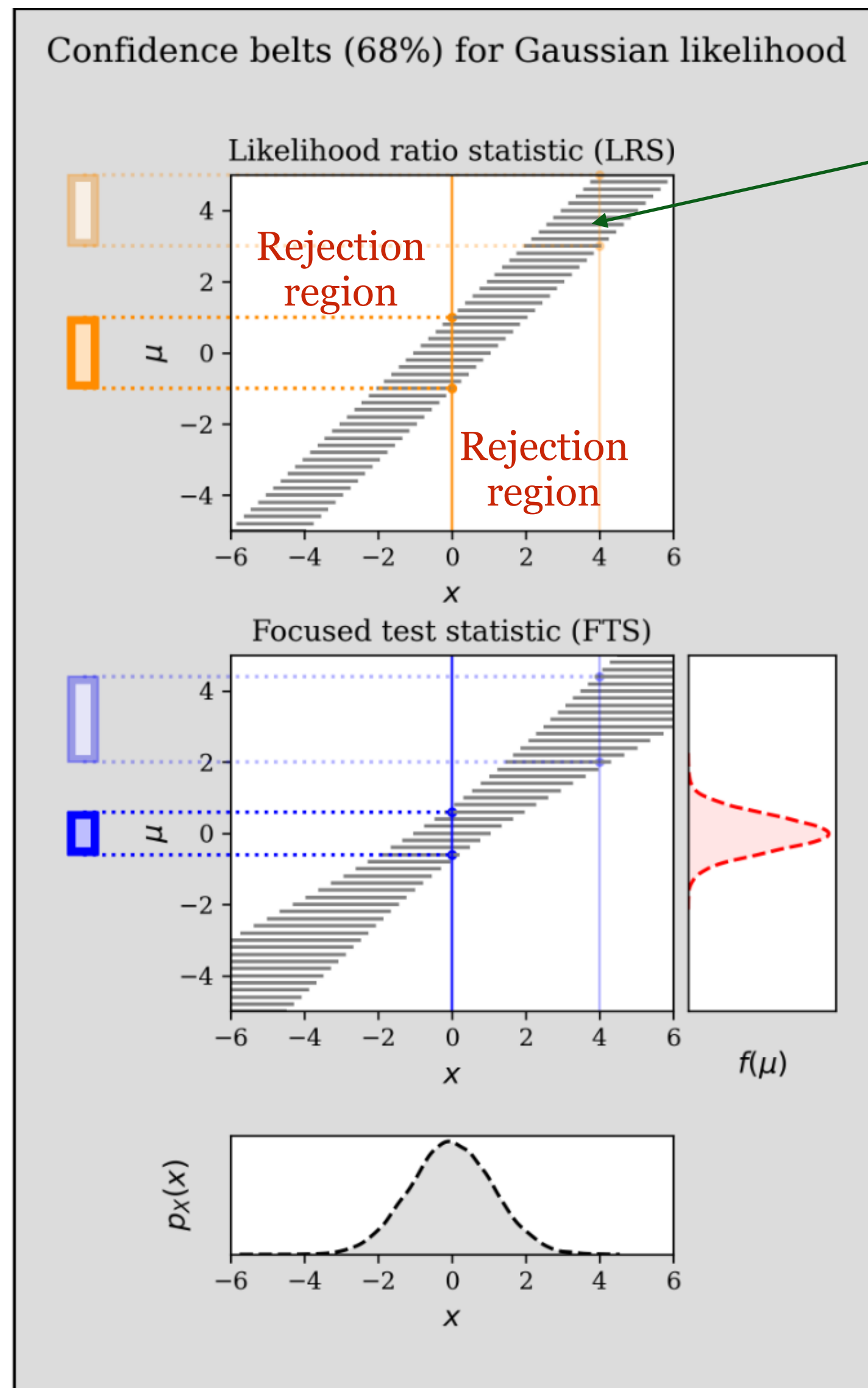
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- FTS gives narrower near focus region by sacrificing power away from focus region

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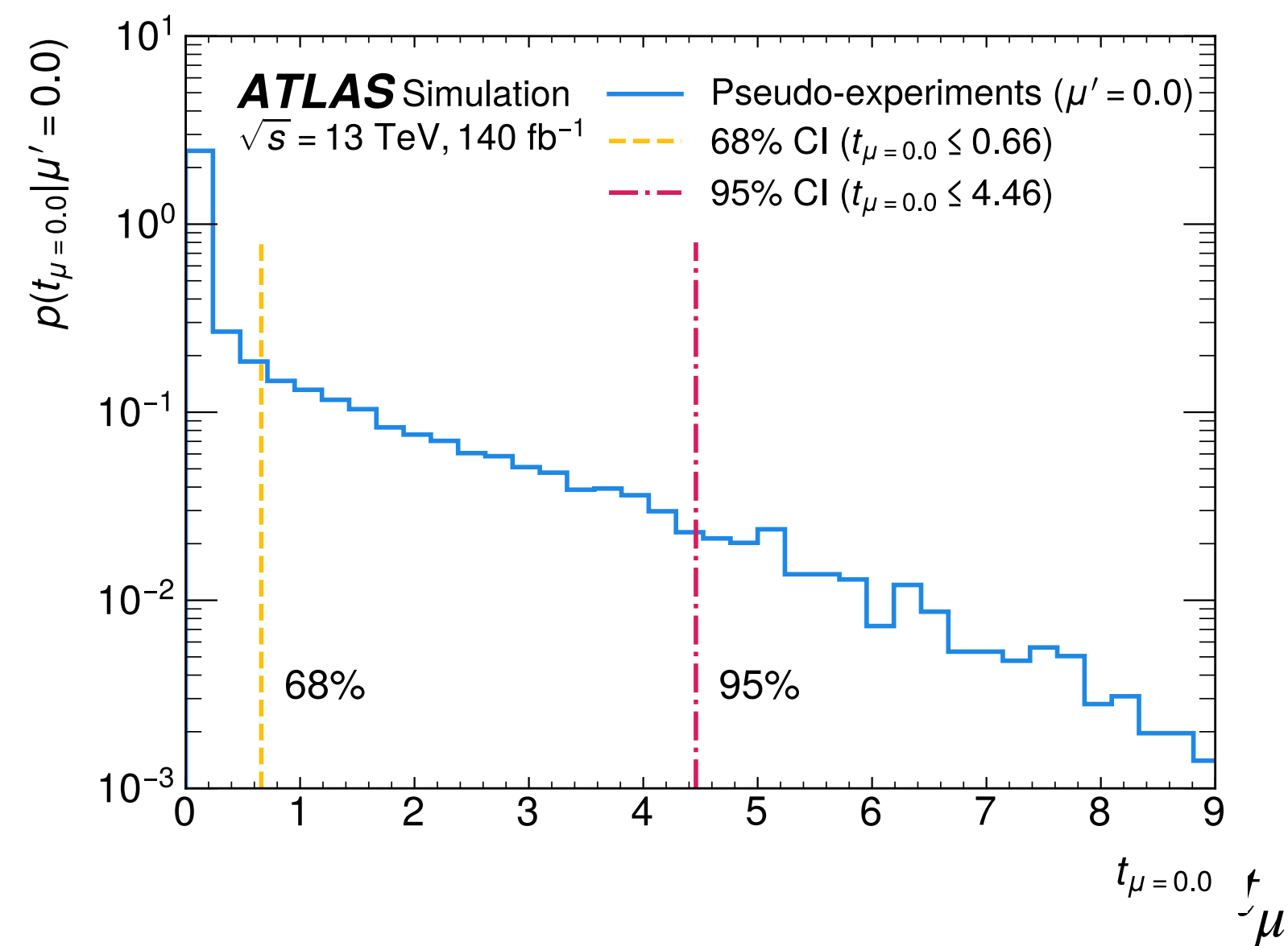
- LRS gives constant CI sizes
- FTS gives narrower near focus region by sacrificing power away from focus region
- You have full freedom to **choose any focus function** that gives you the **best expected sensitivity** in physics-motivated regions using simulated samples

# Computational Challenge: Inverting the test

- Determine 68 % & 95 % CI empirically from this distribution
- Do it for each value of  $\mu$

Distribution of test statistic  $t_\mu$  over thousands of simulated pseudo-experiments

True  $\mu = 0$



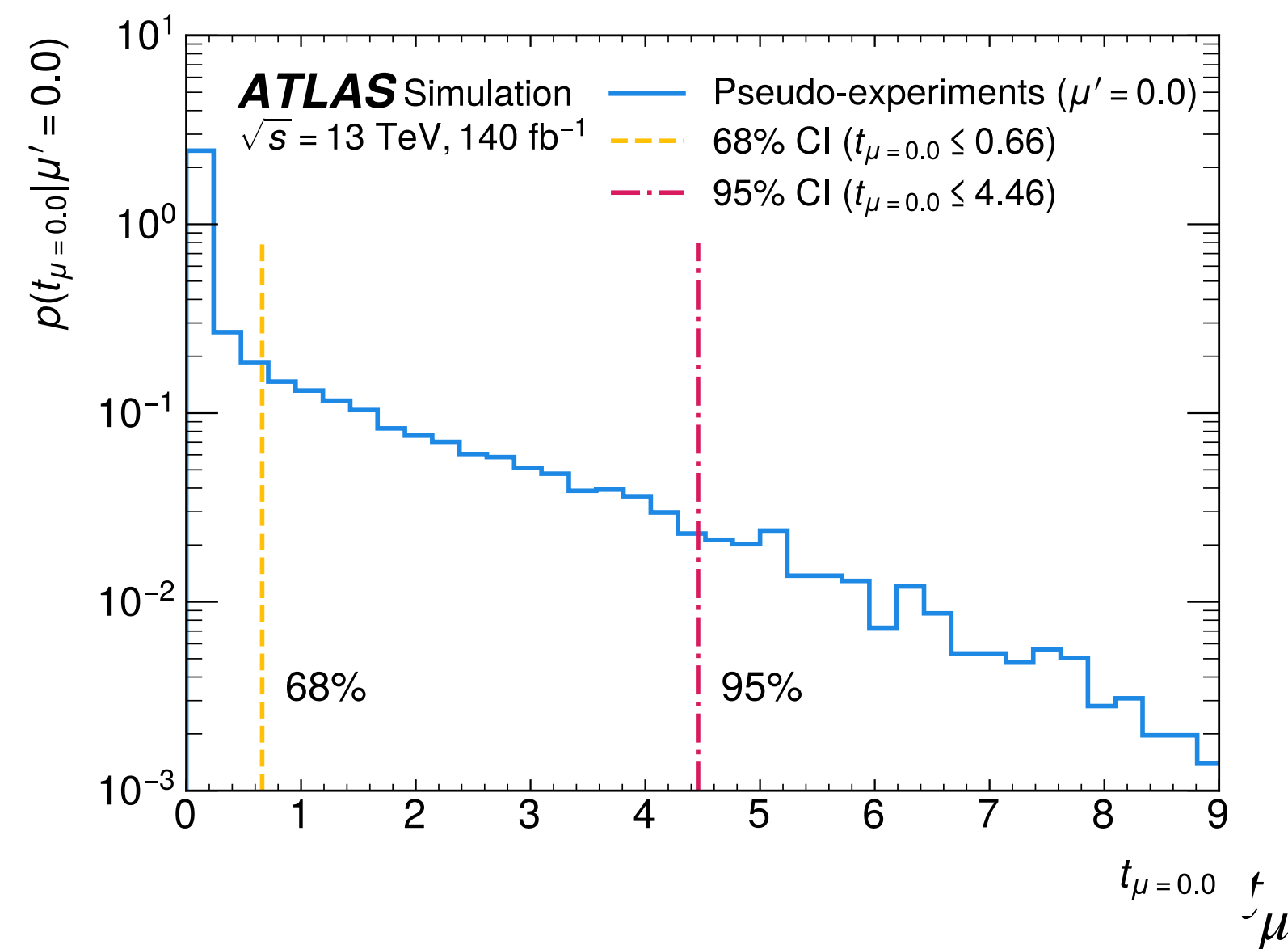
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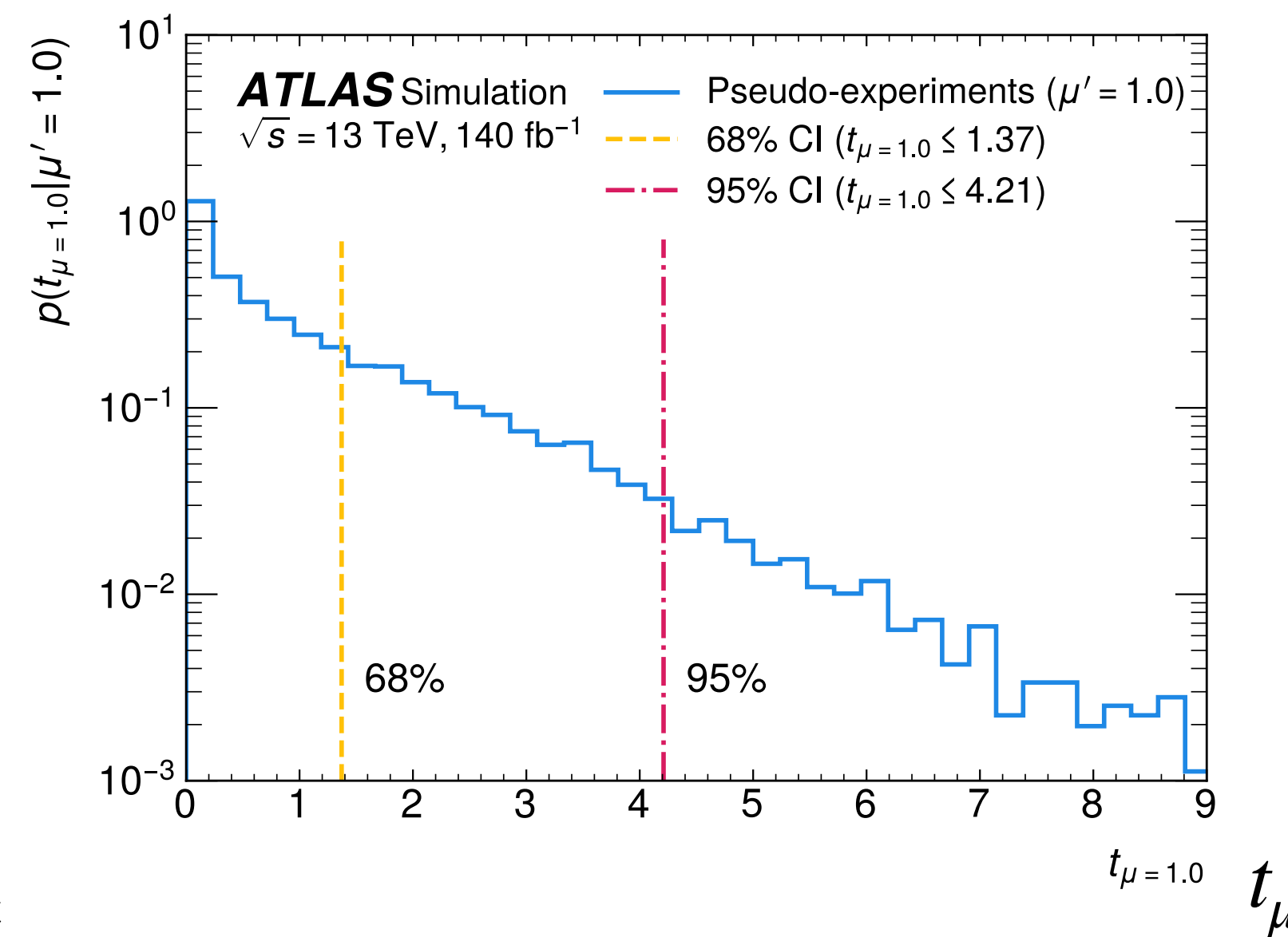
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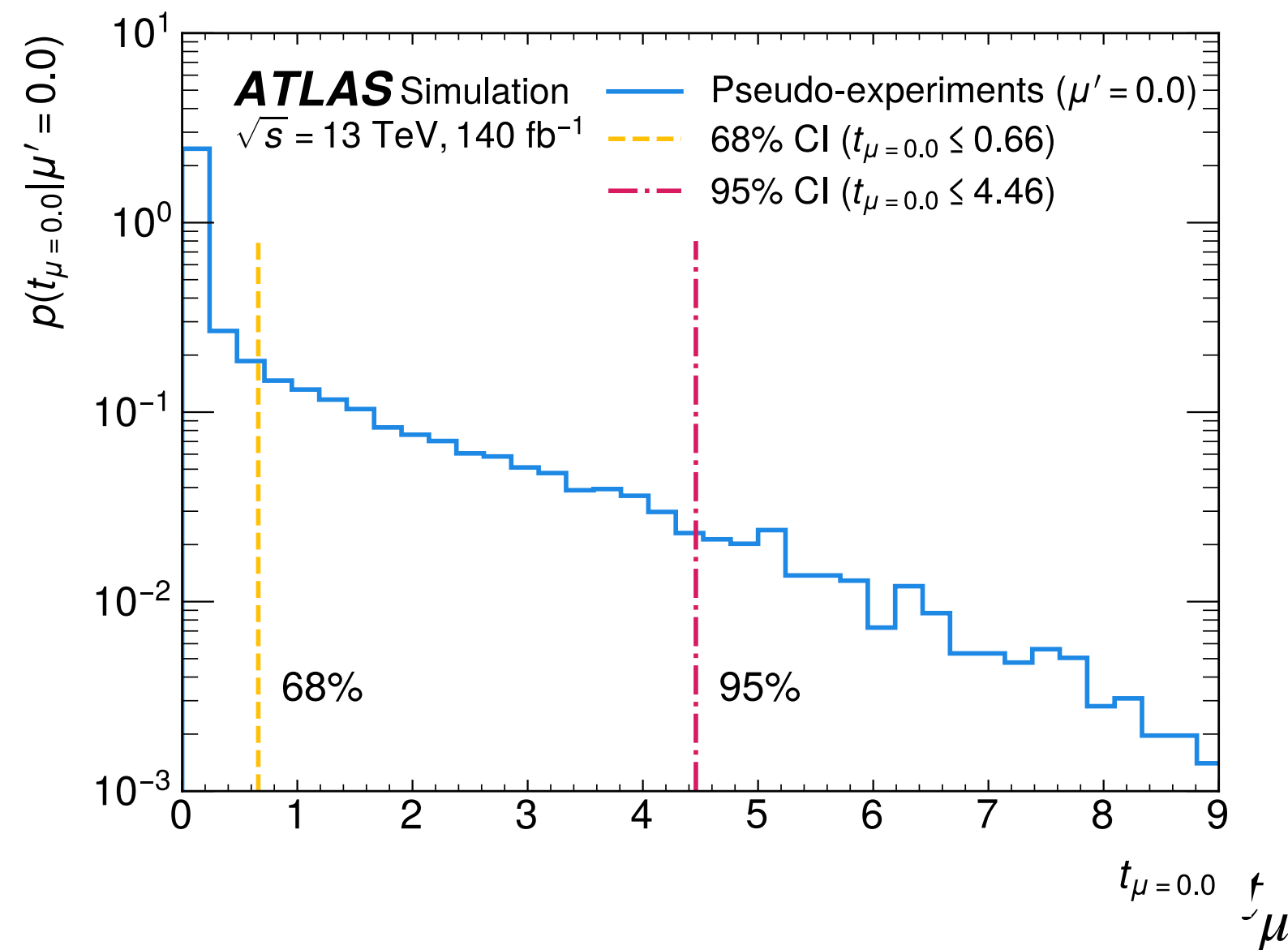


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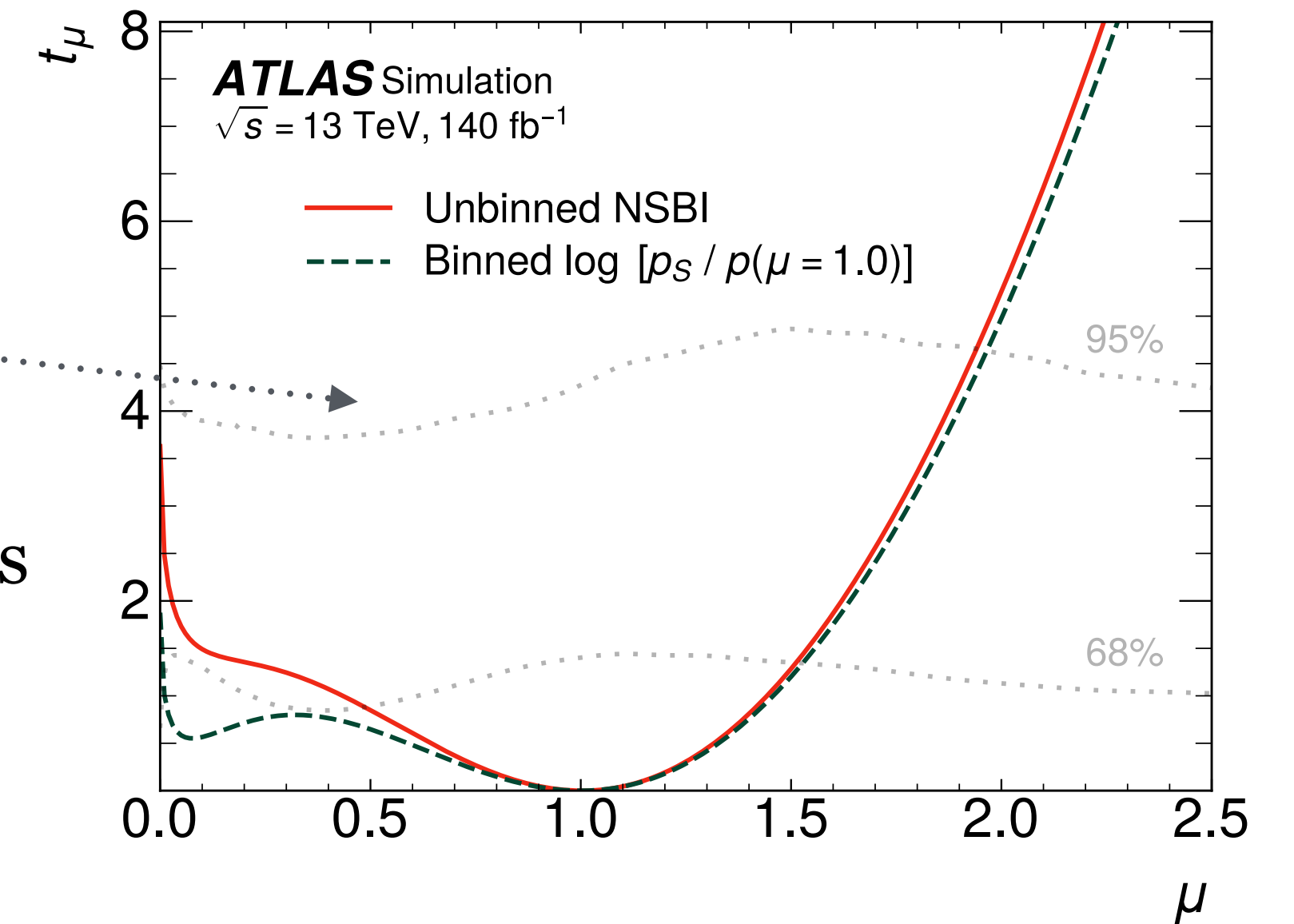
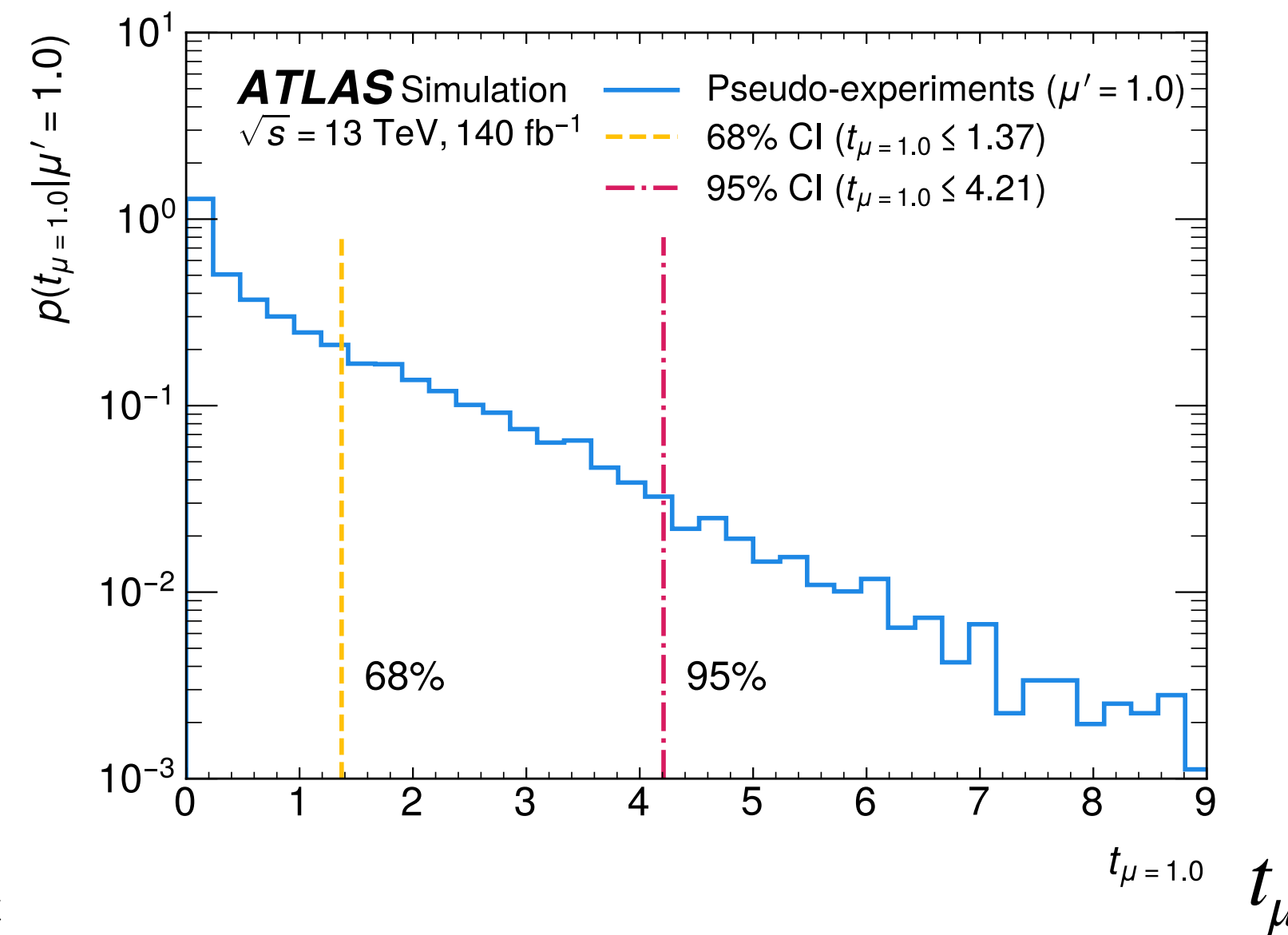
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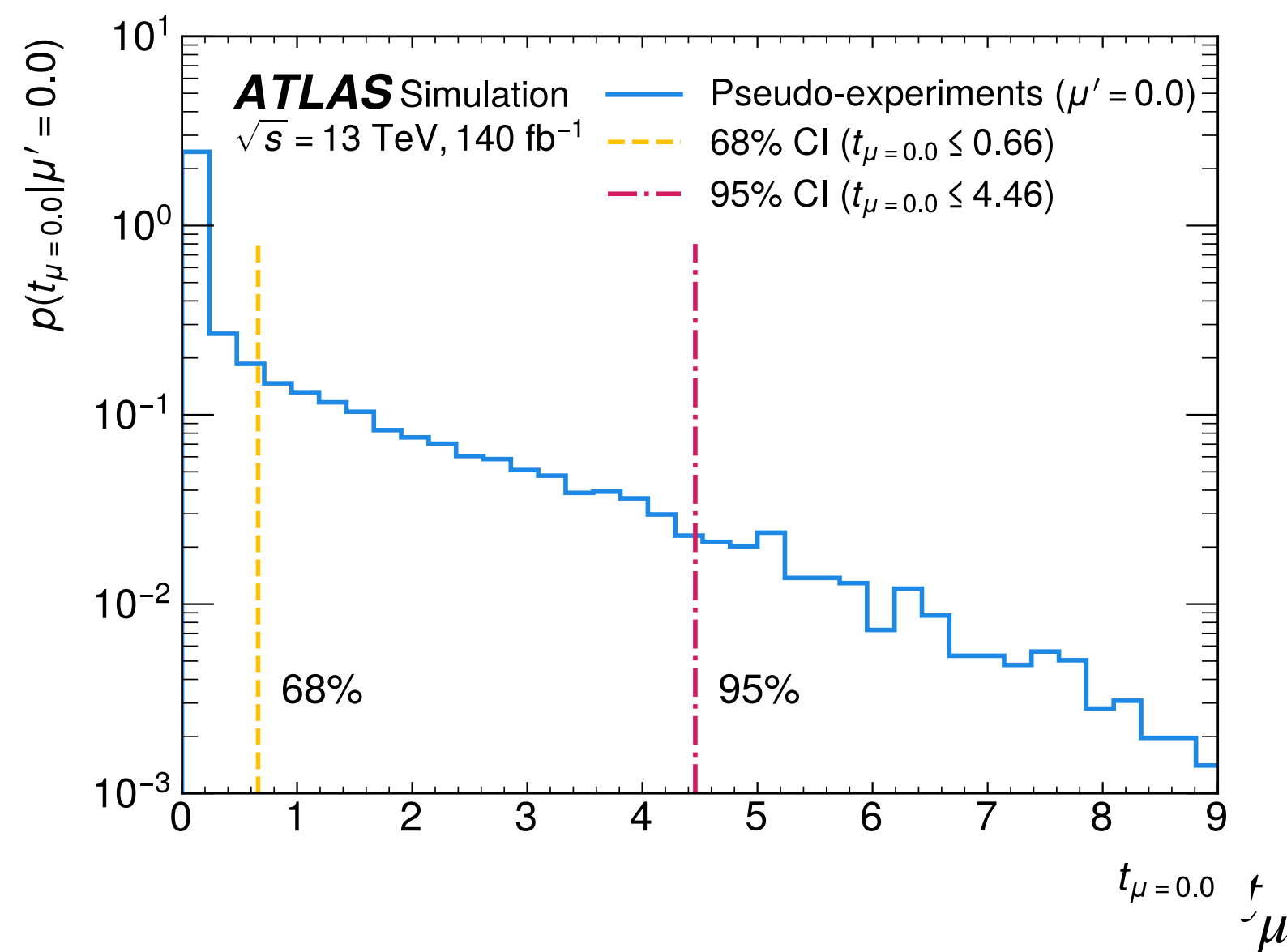


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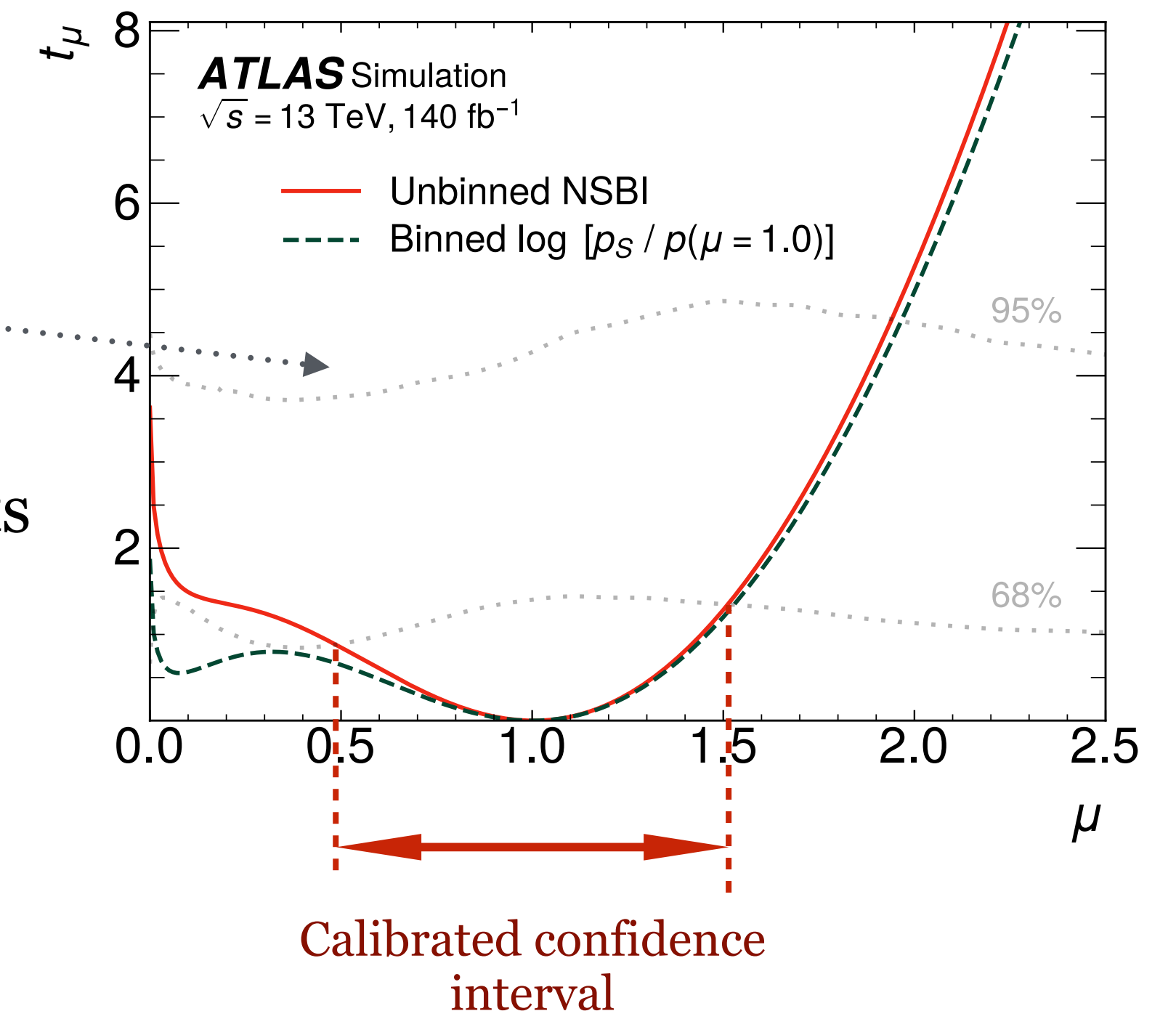
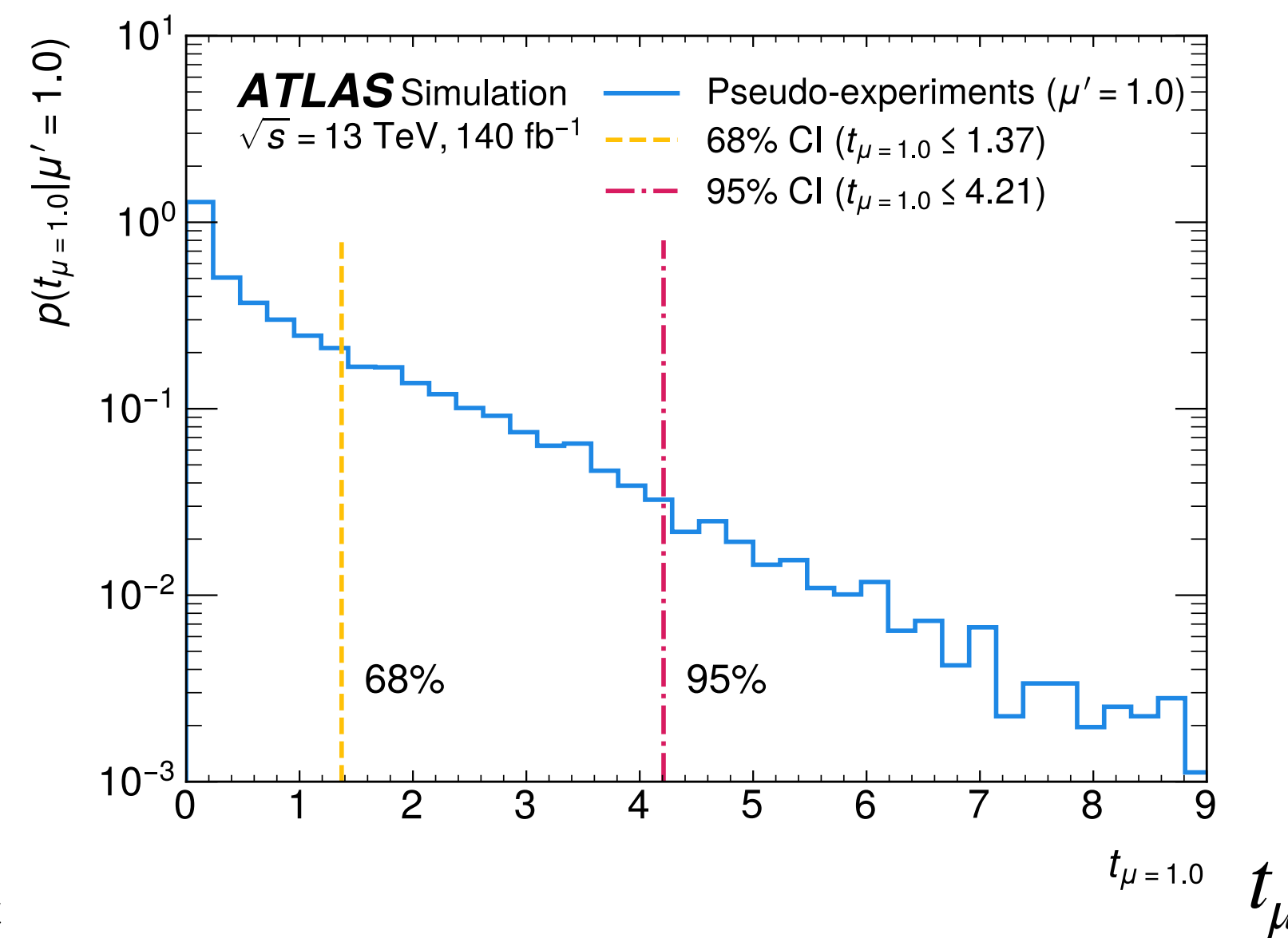
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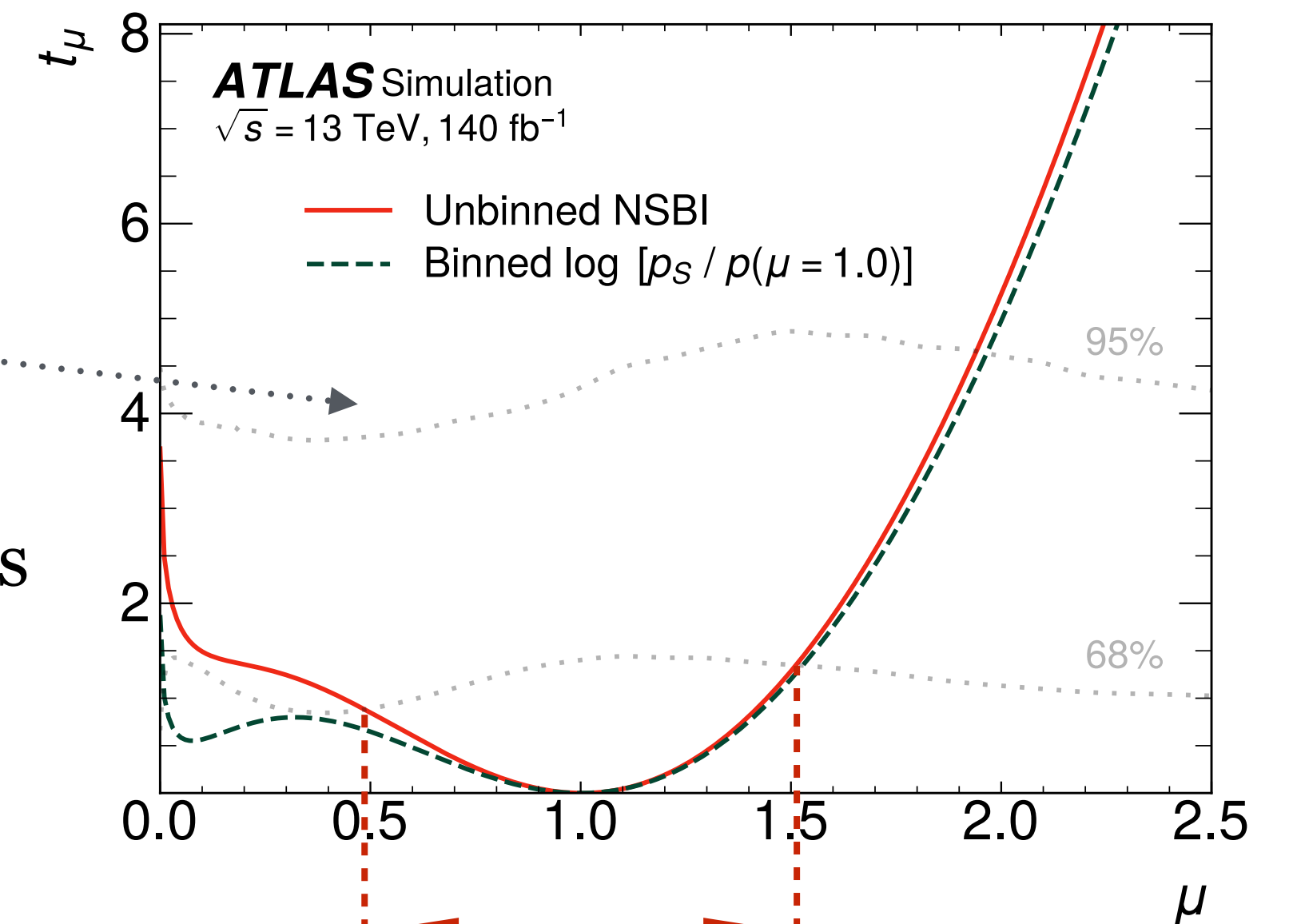
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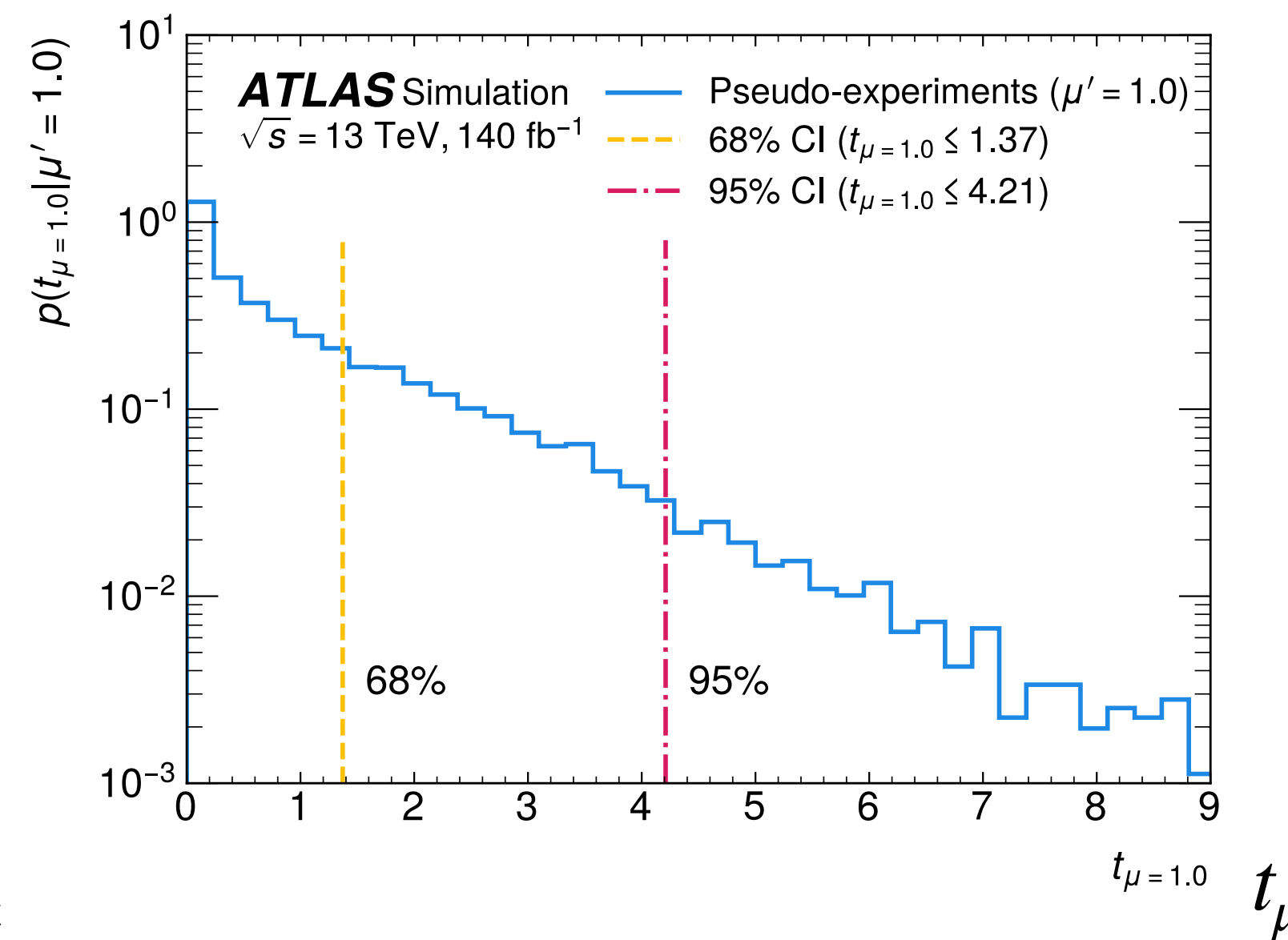
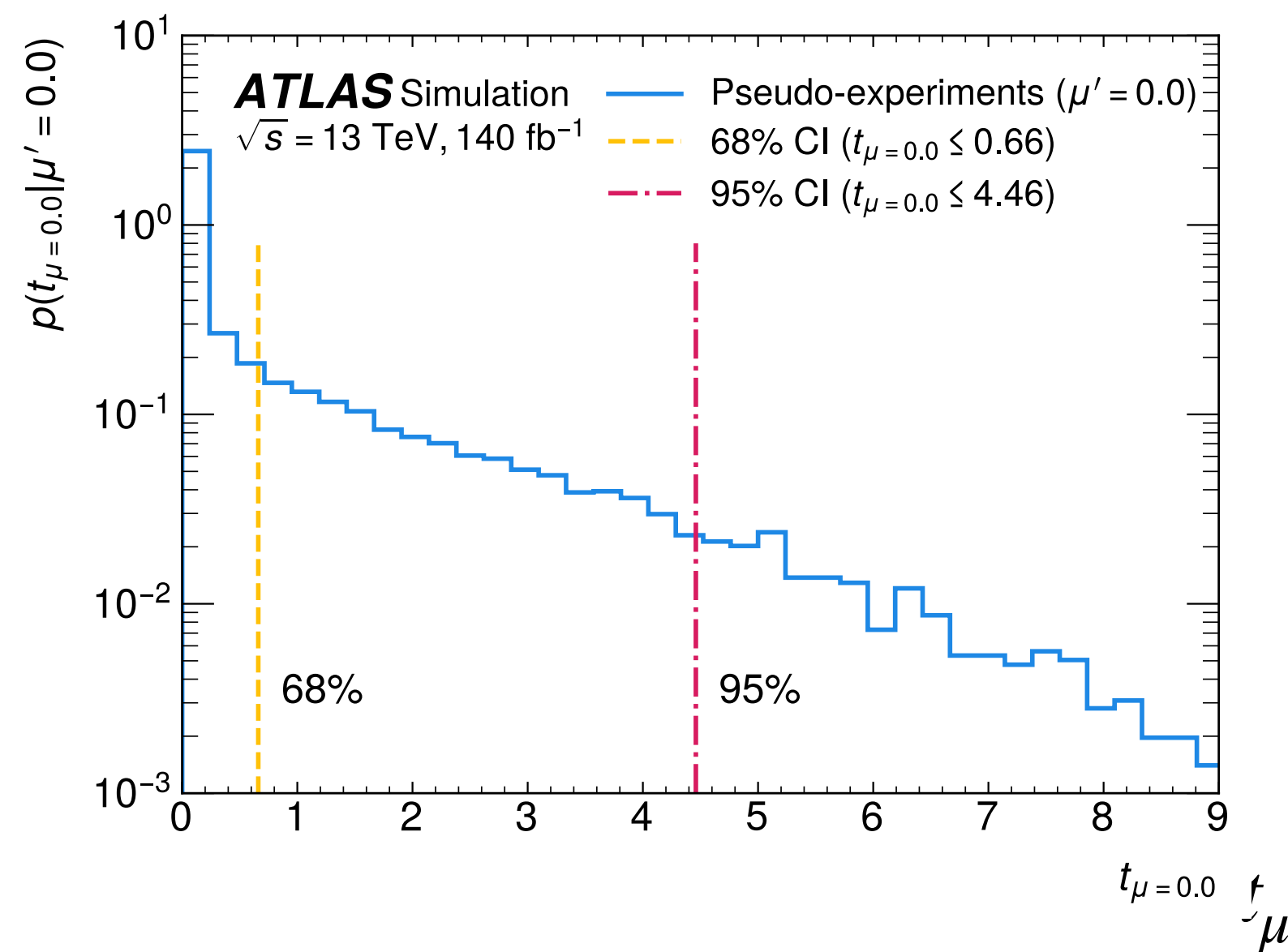
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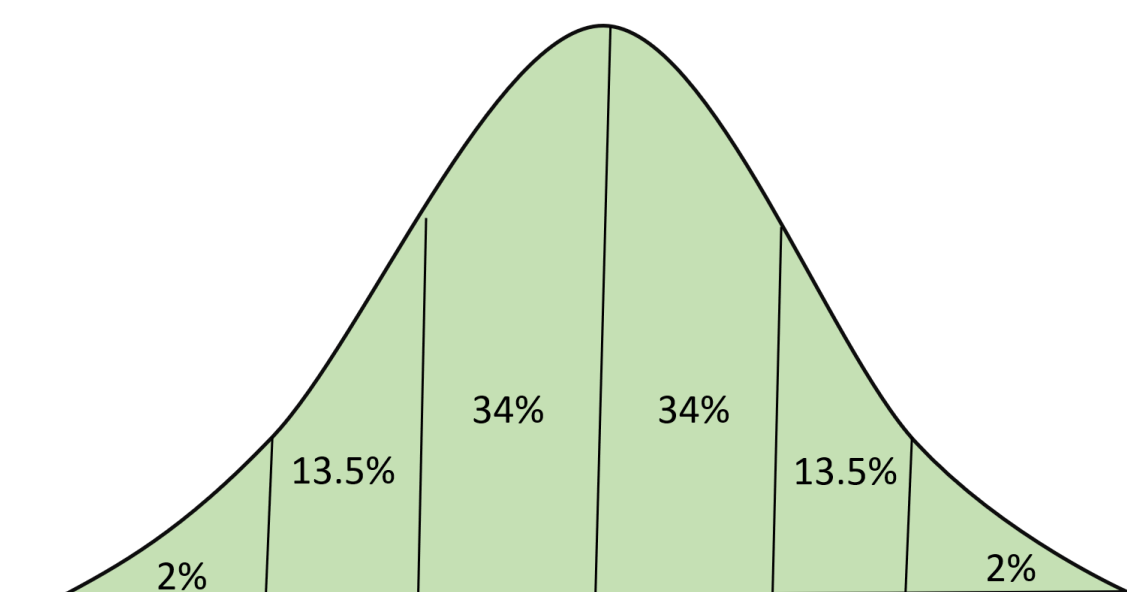


True  $\mu = 0$

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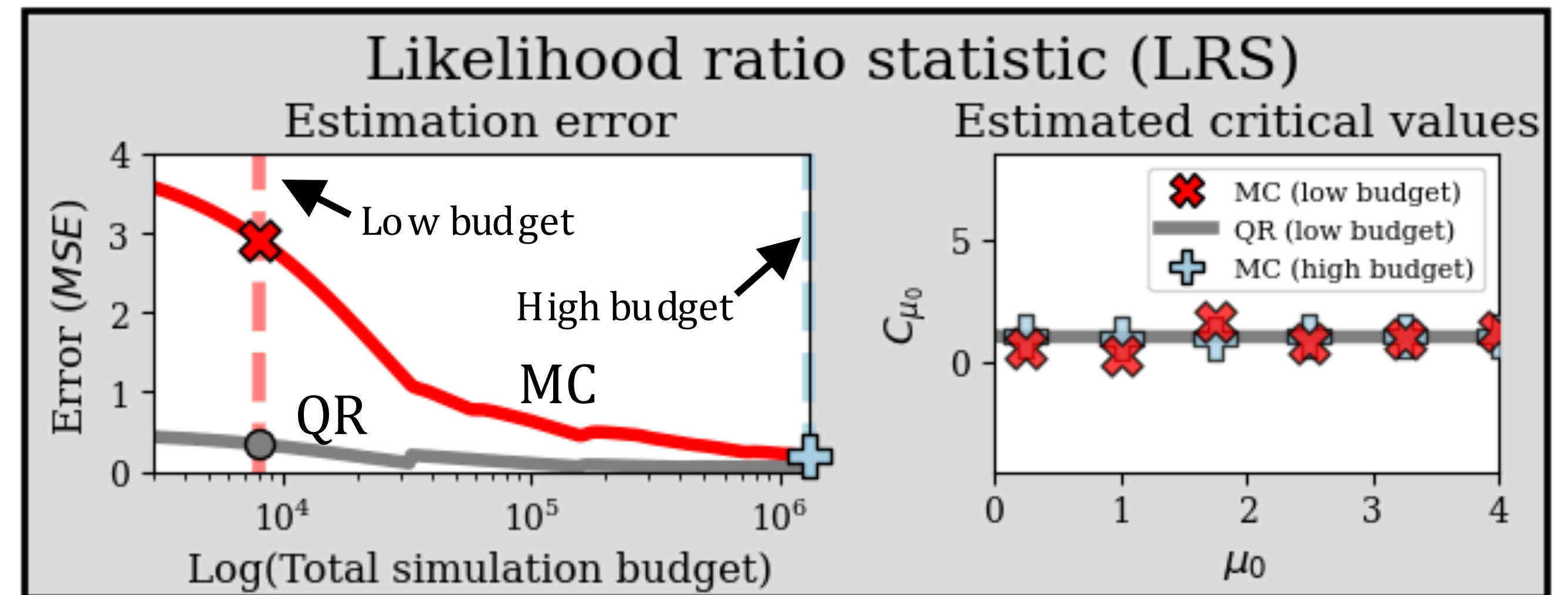


Task is to determine quantiles



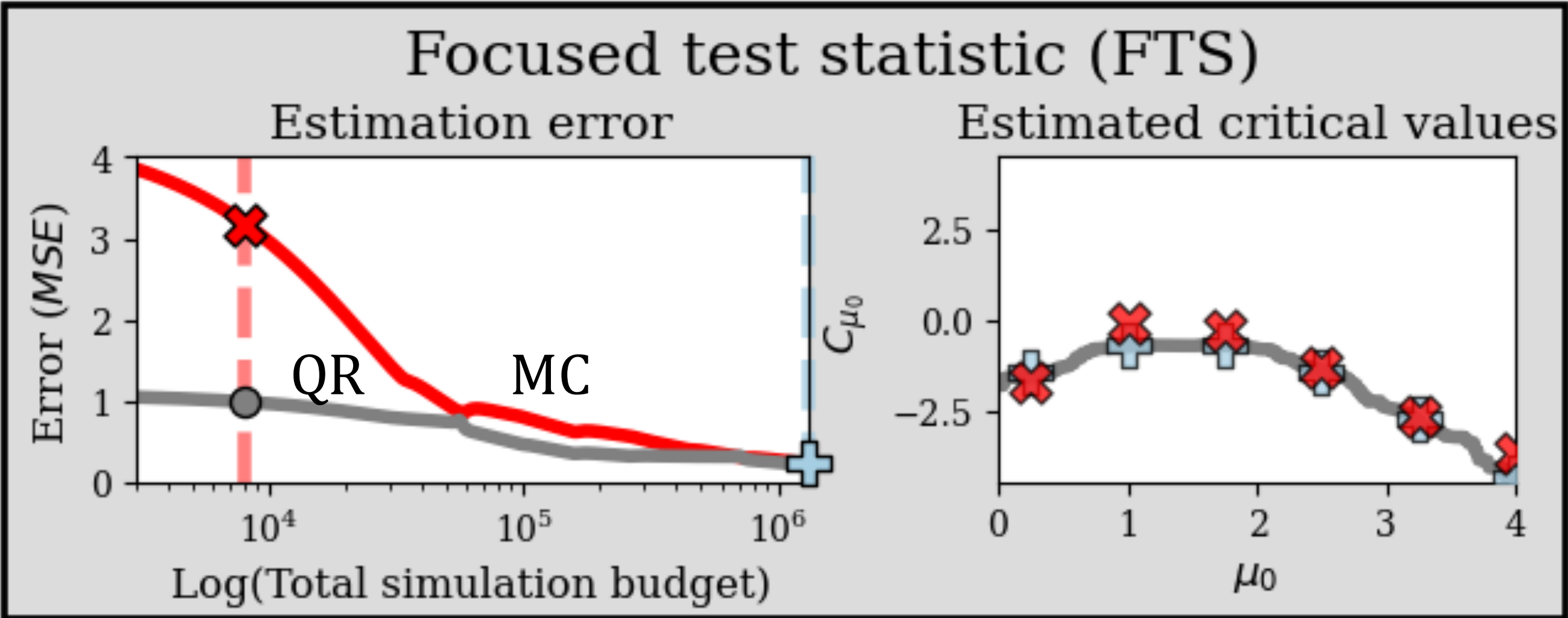
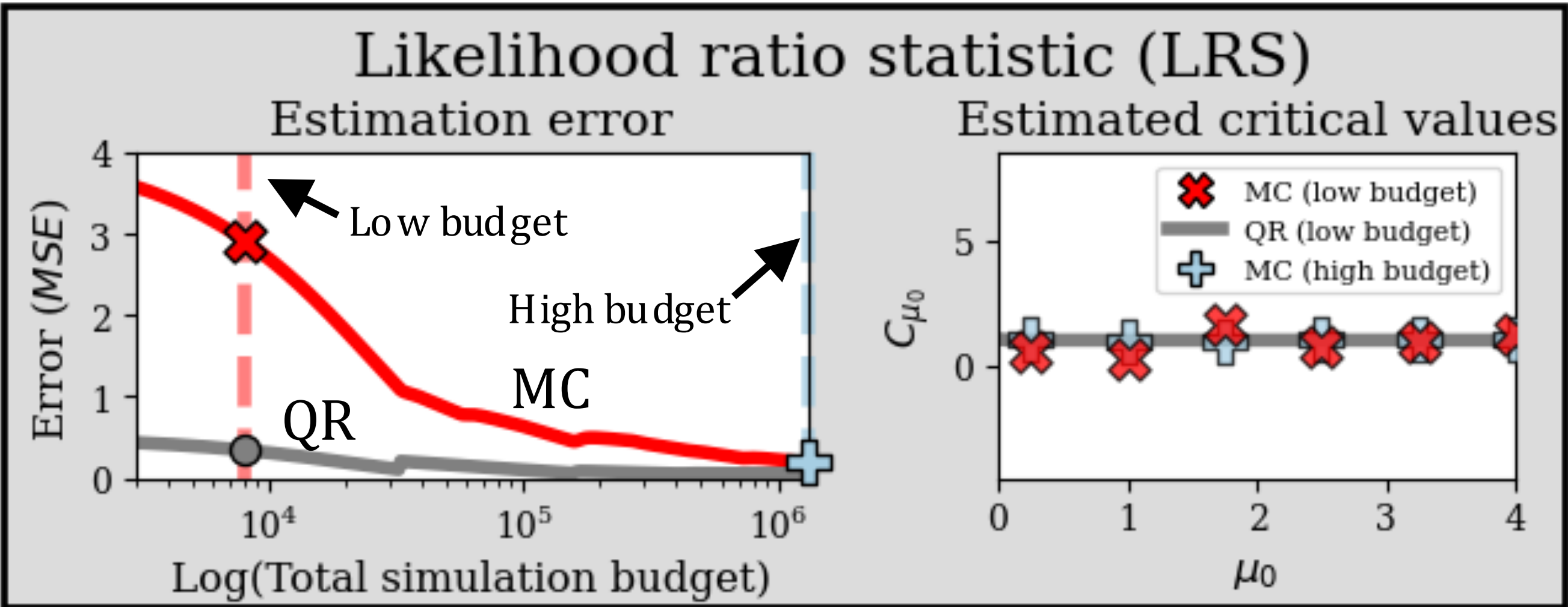
# Perform efficiently with quantile regression!

Can do it faster and more accurately with ML technique known as 'quantile regression' (QR)



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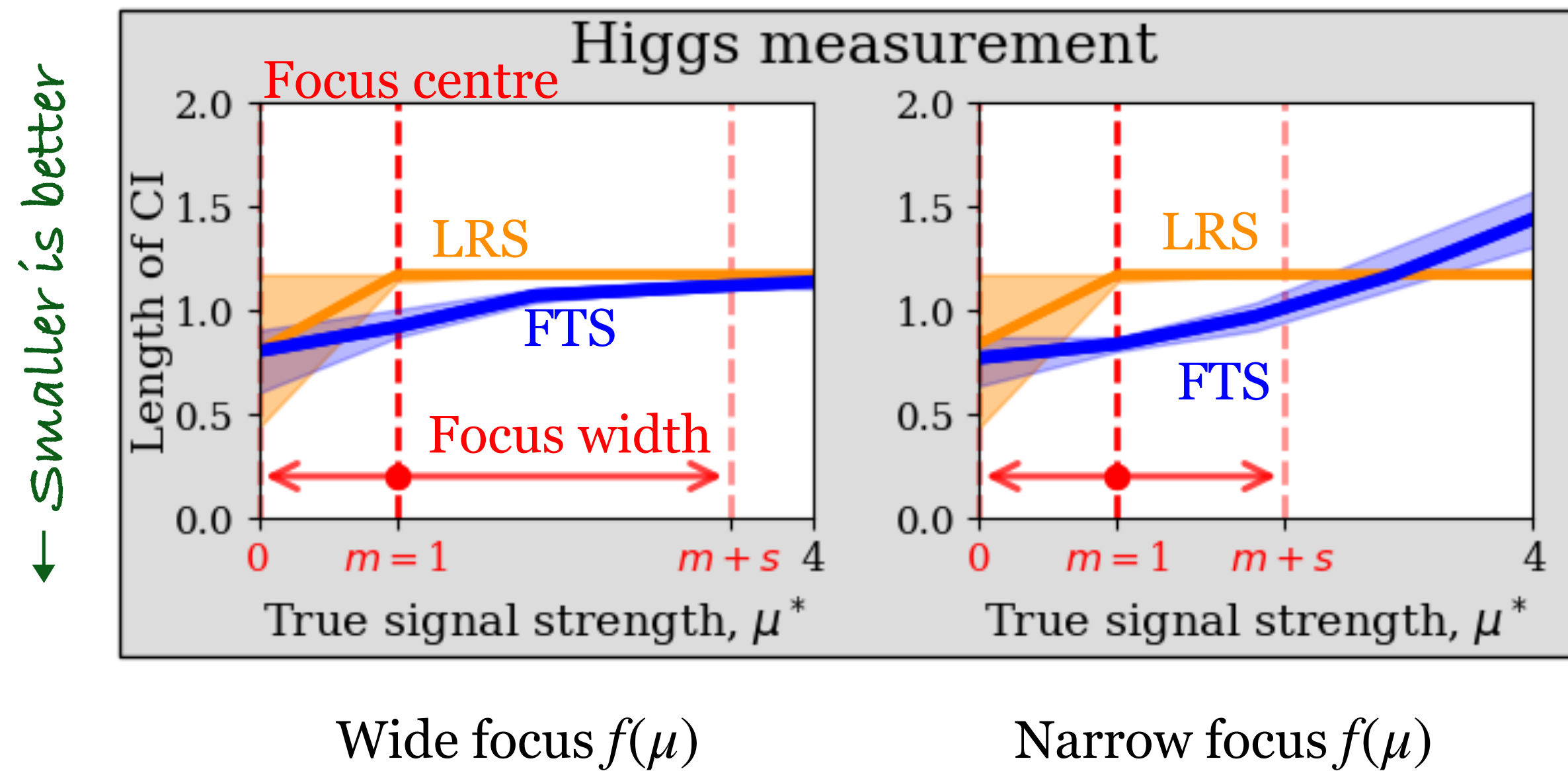


# Now ready to compare LRS to FTS on HiggsML benchmark dataset

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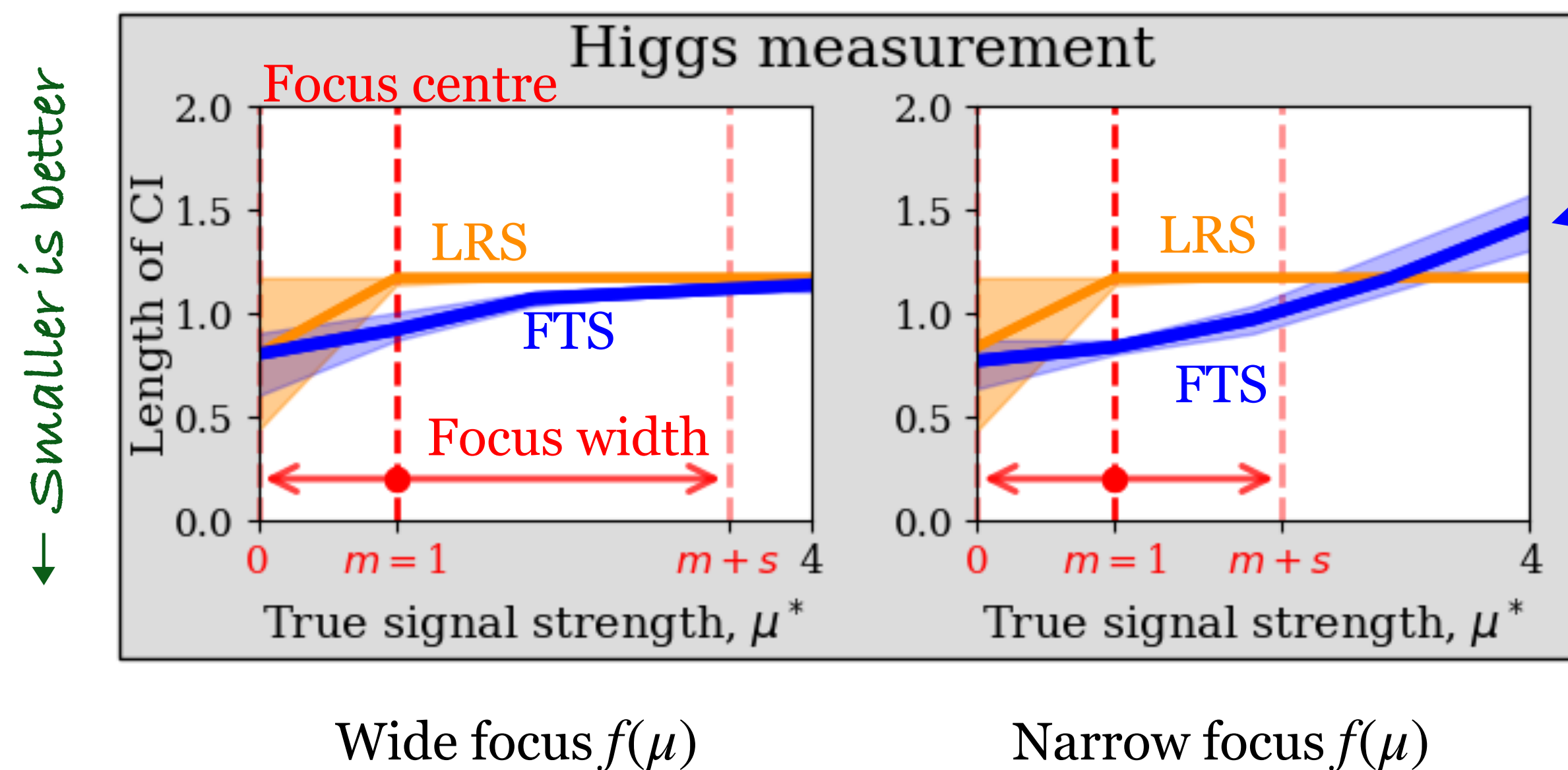
$H \rightarrow \tau\tau$  ATLAS simulated public [benchmark dataset](#)

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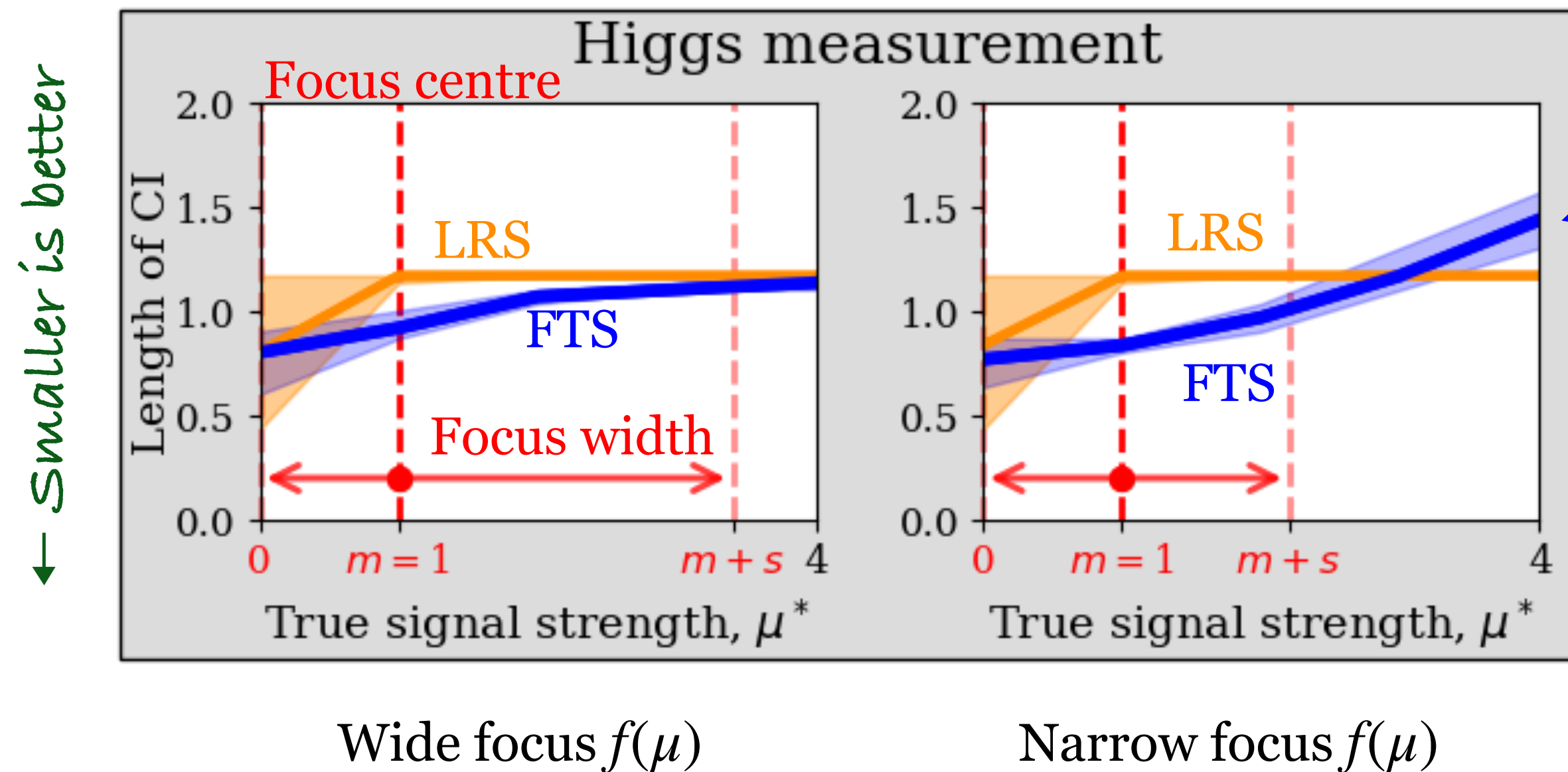
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FTS worse than LRS when true  $\mu$  is very far away from focus region, but coverage of confidence intervals still correct

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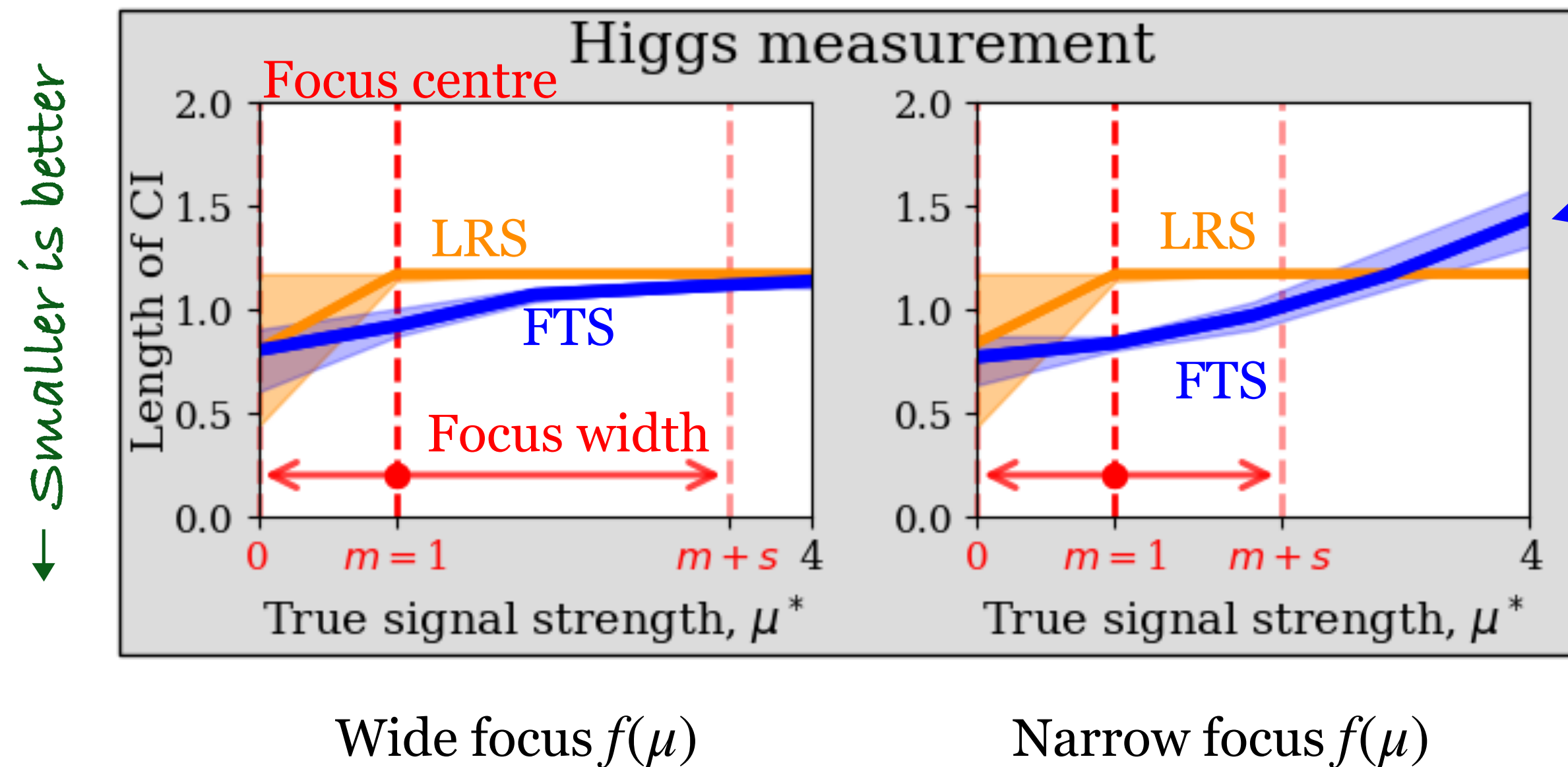


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Even when it is worse, the coverage of the confidence interval is still guaranteed by Neyman construction!

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FTS worse than LRS when true  $\mu$  is very far away from focus region, but coverage of confidence intervals still correct

$H \rightarrow \tau\tau$  ATLAS simulated public [benchmark dataset](#)

*which focus would you choose?*

Even when it is worse, the coverage of the confidence interval is still guaranteed by Neyman construction!

funding gets harder to secure, principal investigators are in their office writing grants while the trainees get to do the cool stuff.

**Bryan W. Jones** is a retinal neuroscientist at the University of Pittsburgh in Pennsylvania.

## **AISHIK GHOSH** **STUDENTS OVERTURN** **LONG-HELD ASSUMPTION**

I have worked on experimental particle physics since 2015, searching for Higgs bosons at CERN, Europe's particle-physics lab near Geneva, Switzerland, and now also working on the Deep Underground Neutrino Experiment (DUNE) in the United States. For this research, there's one statistical test we've used for decades to confirm the existence of a new particle – the generalized likelihood ratio test (GLRT). This compares two models – a simple null hypothesis, which includes no new particle or matter being discovered, and a more complicated alternative model, which includes a new particle with many possible values of strength.

In December 2024, a couple of PhD students working with my collaborator, Ann Lee, a data scientist at Carnegie Mellon University in Pittsburgh, Pennsylvania, were confident they could disprove the assumption that the GLRT was optimal. In the corner of my mind, I hoped they would prove us wrong. I gave them one of the most famous Higgs boson data sets to play around with. By early 2025, they showed that, although our previous physics results weren't wrong, our use of the GLRT wasn't ideal because it assumed large sample sizes

are always generated, which is often not the case. Instead, the test left valuable information on the table. That day was special. I was still sceptical and I went through a battery of checks because I had to go back to my community and defend the PhD students' work, but it was all correct. The paper is currently in review, receiving a great deal of scrutiny.

Together, we produced a statistical test that will drastically improve our ability to make discoveries in particle physics, for example in searches for a new particle such as dark matter, where we expect to see only a few signal events at best. As a scientist, I want deeply held beliefs to be questioned. It was a real shock to the particle-physics community. Young people find it exciting. Senior members are still highly sceptical, as they should be, but they are coming around. As the DUNE experiment comes online, with this new statistical model in place, we hope to make precise measurements about neutrinos much sooner than anticipated.

**Aishik Ghosh** is a fundamental physicist at the Georgia Institute of Technology in Atlanta.

## **RAFIK TAREK NEME GARRIDO** **SHOCKING** **CORAL FIND**

A couple of years ago, after a day of pouring rain, the water on the Caribbean coast of Colombia was crystal clear and my master's student, Jorge Mareno, managed to take pictures of corals that no one knew existed here. We could find no scientific reports of corals in the area. Typically, the water is pretty turbid because the

Magdalena River, which flows from the south of the country to the Caribbean Sea, brings chemicals and pollutants. It's an ongoing ecological and social challenge, but these corals must be adapting to these conditions. We did a sampling campaign across three days with a boat, using environmental DNA to find areas where corals, sponges and fish successfully survive the conditions. Most of the records are completely new for the region. It's super gratifying.

**Rafik Tarek Neme Garrido** is an evolutionary biologist at the University of the North in Barranquilla, Colombia.

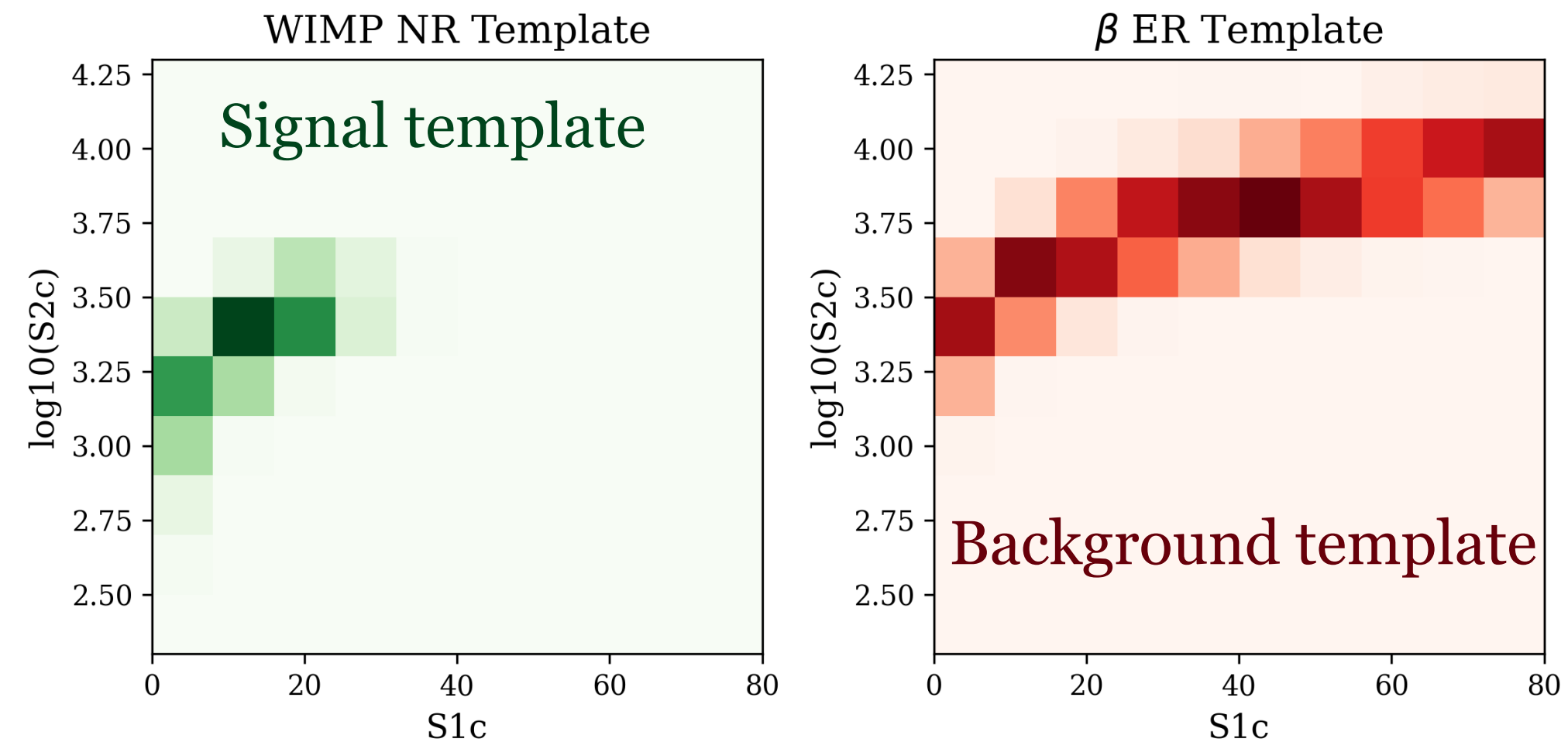
## **TIM CURRAN** **BURN** **PREDICTIONS**

In my group, we test the flammability of plant species using a barbecue. The results can help with fire-mitigation policies and with understanding the evolution of flammability. As part of an outreach activity, we host school-children at the university who haven't had much exposure to academia before. We ask the kids to predict how a particular plant species will behave – for example, what characteristics will make it burn less or more – and then we see who is right. The kids get really into it. They ask amazing questions, the same kind that peer reviewers have asked us, including questioning our methodological assumptions, such as "why do you only blowtorch them for ten seconds?"

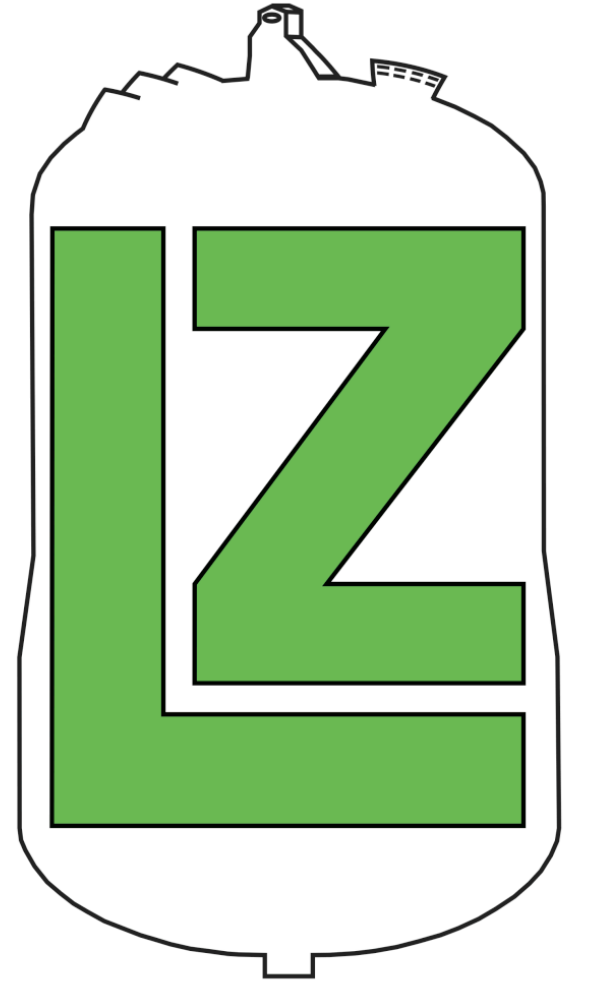
Most of the really good days doing science have been associated with young students having a light-bulb moment. In the rather



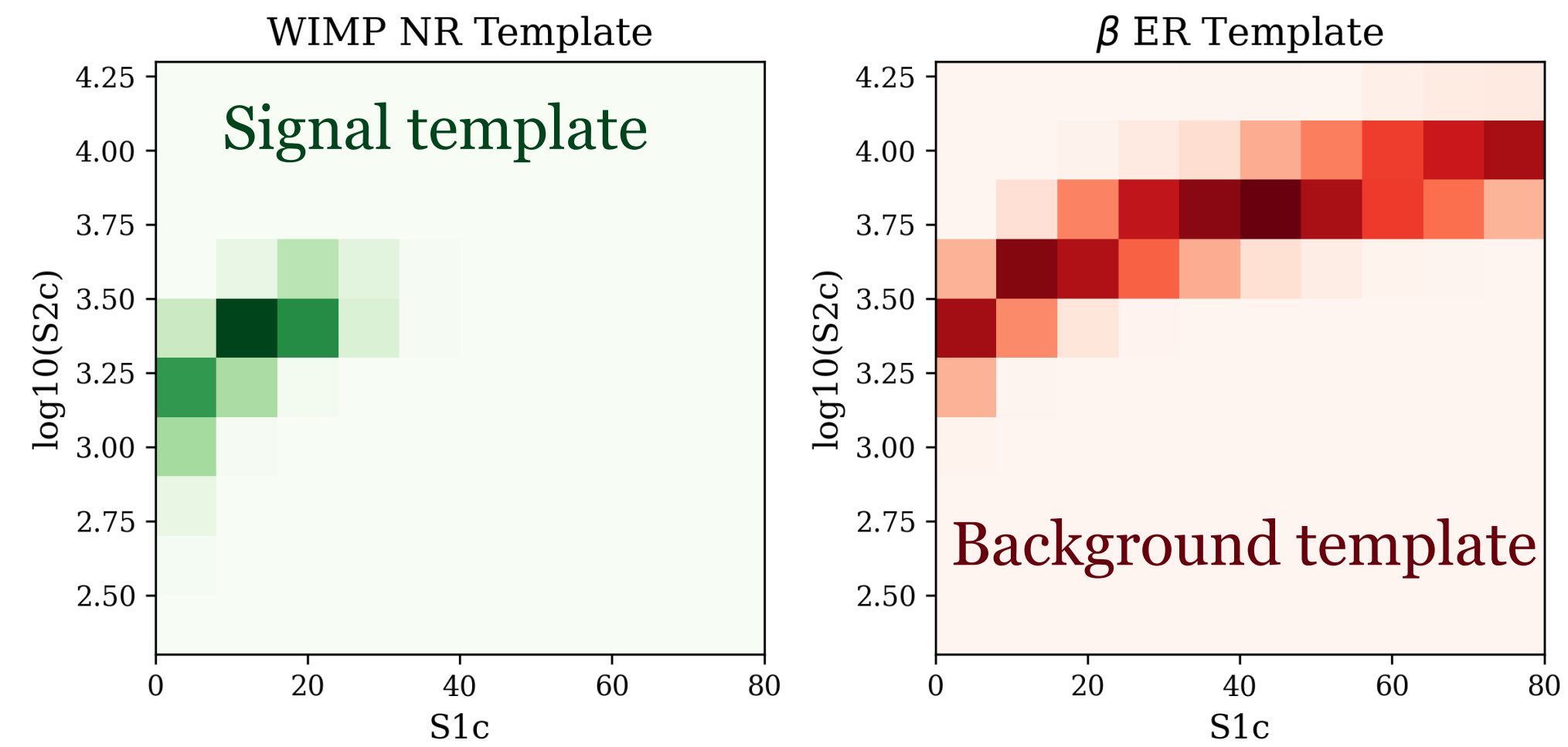
# Second case study: Dark matter search in LZ experiment



Simulated to mimic [LZ data](#)

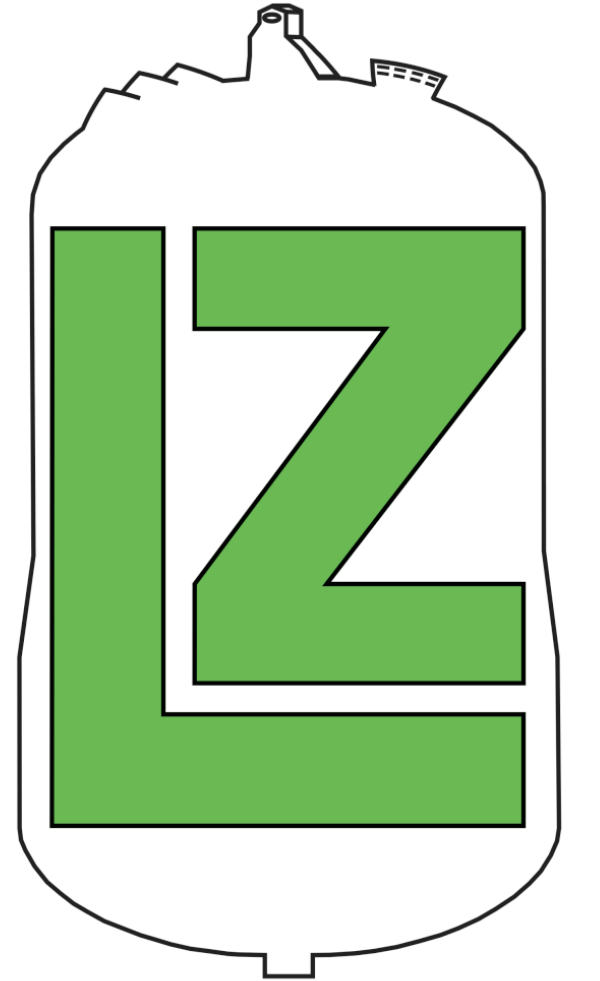


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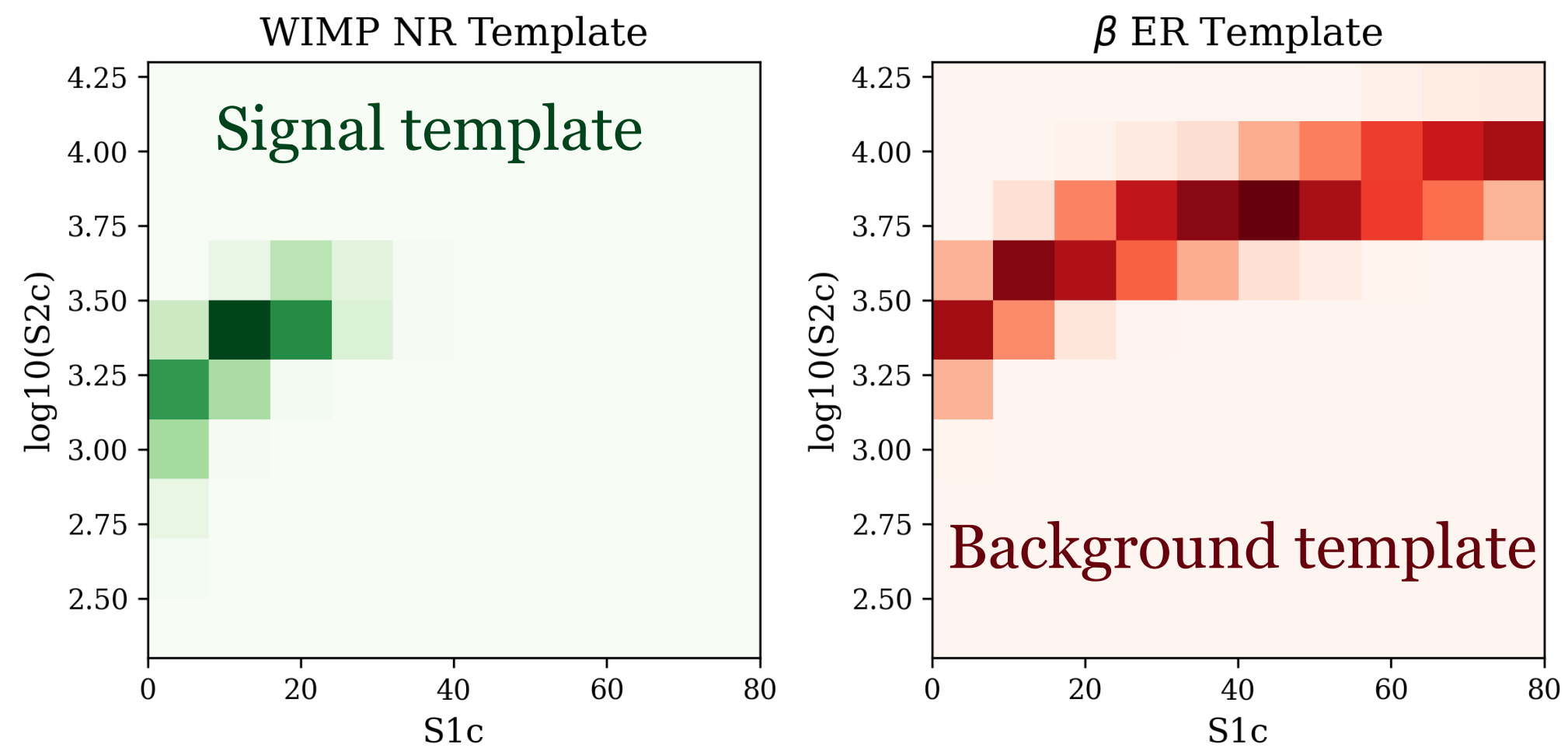


Simulated to mimic [LZ data](#)

**Focus now placed at  $\mu = 0$ , we want to set upper limits on DM signal**

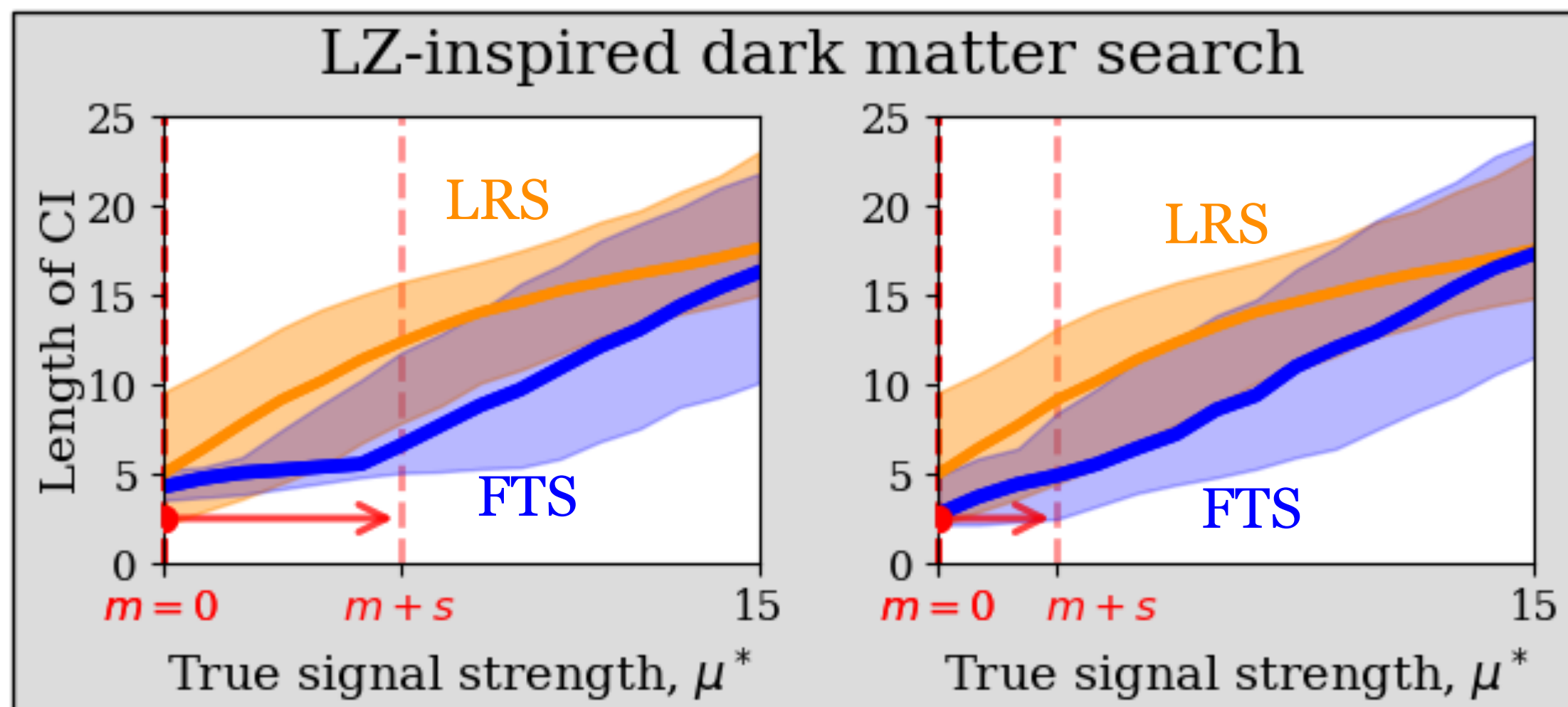
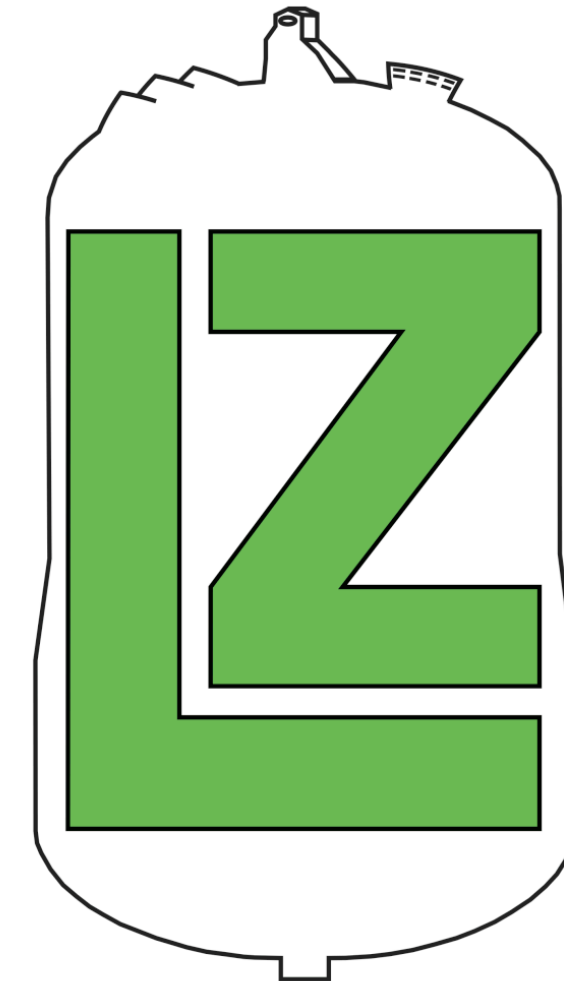


# Second case study: Dark matter search in LZ experiment



Simulated to mimic [LZ data](#)

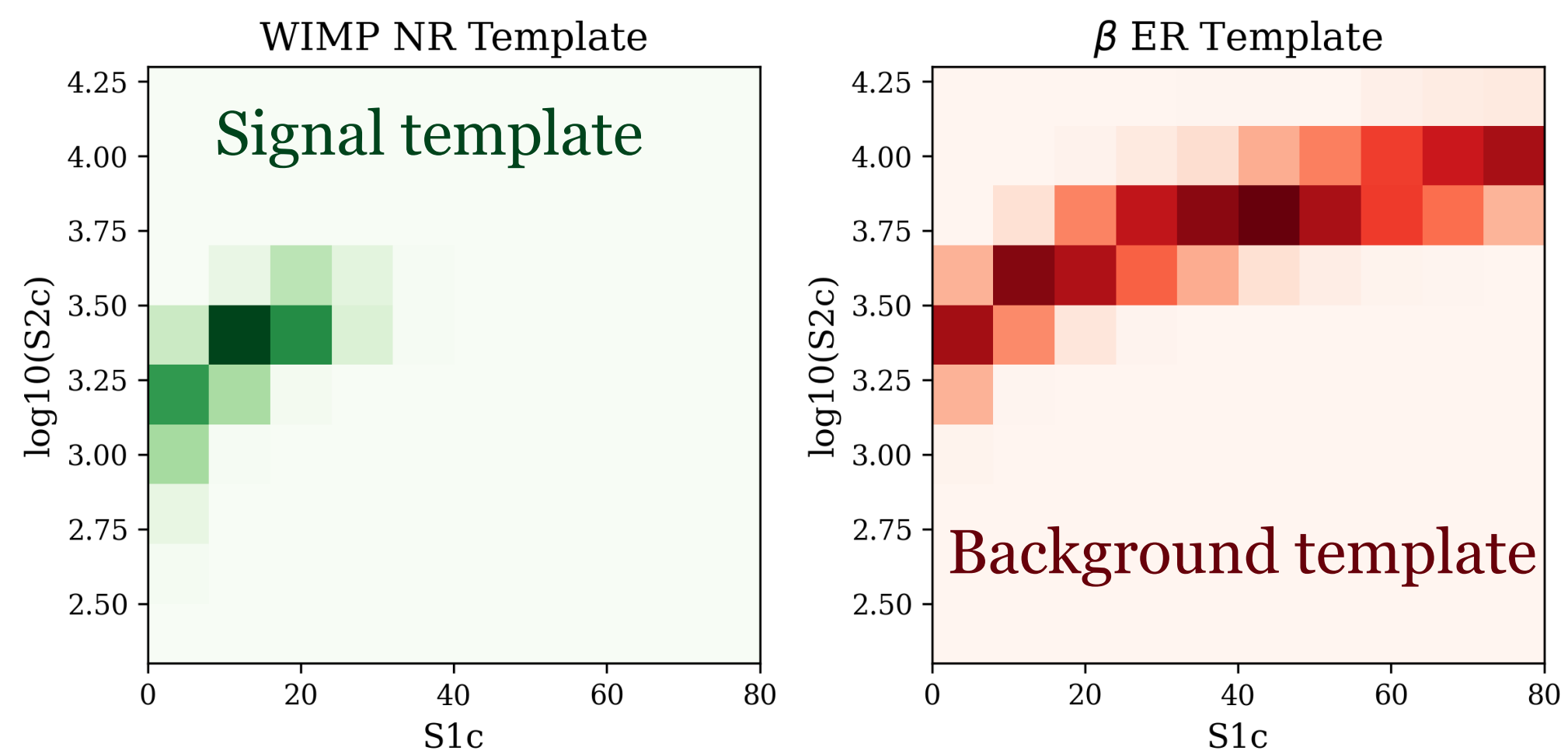
**Focus now placed at  $\mu = 0$ , we want to set upper limits on DM signal**



Wide focus

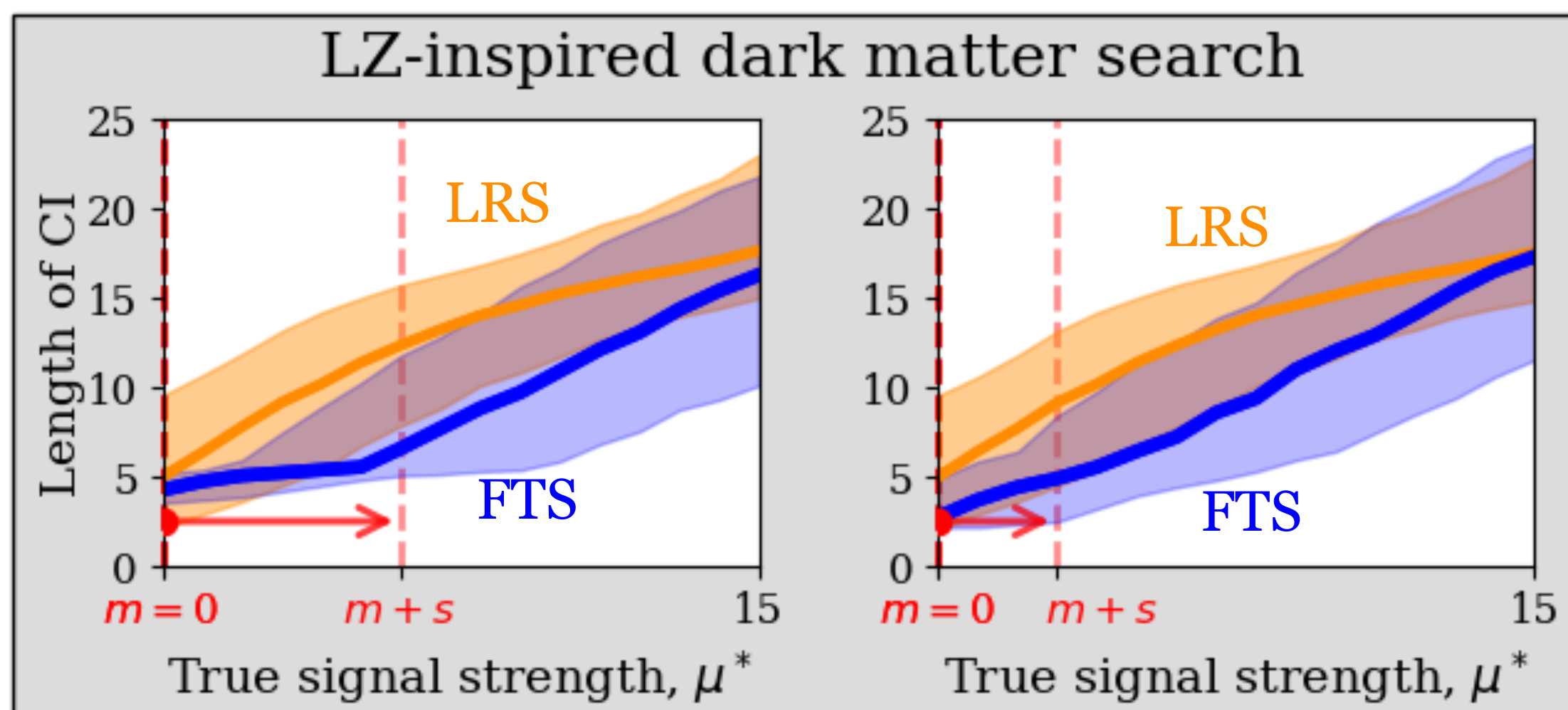
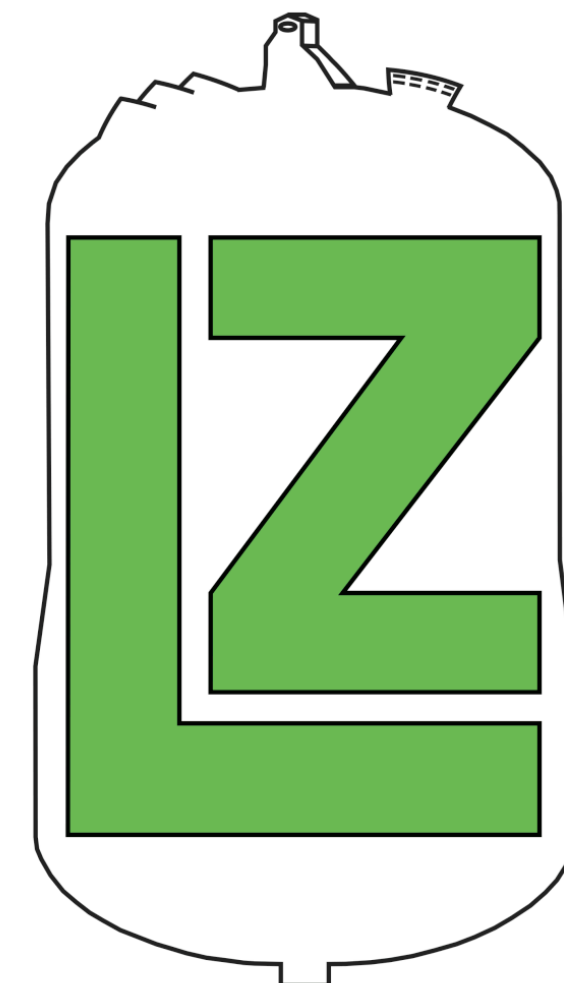
Narrow focus

# Second case study: Dark matter search in LZ experiment



Simulated to mimic [LZ data](#)

**Focus now placed at  $\mu = 0$** , we want to set upper limits on DM signal



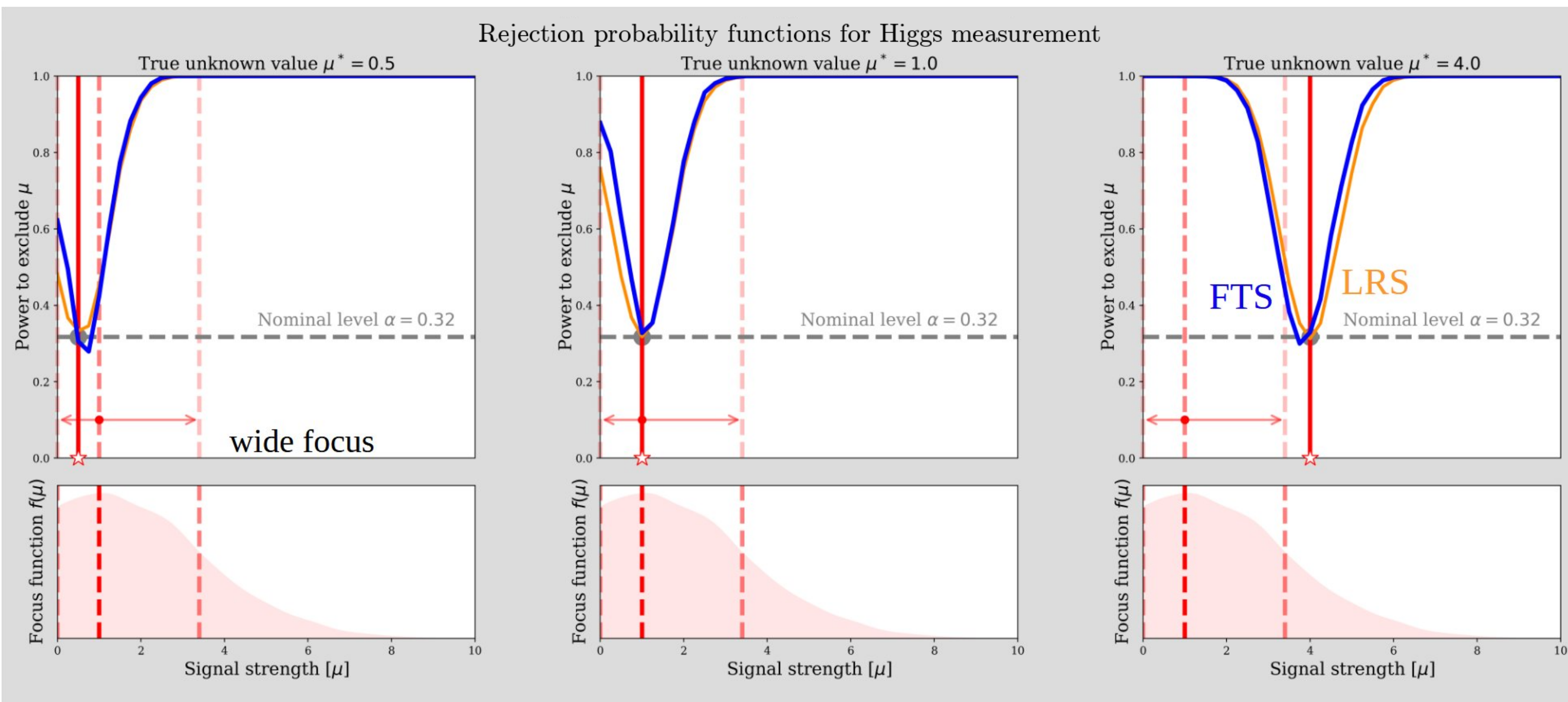
FTS gains stable over large parameter space for both focus functions

More precision even when there is actually a signal in the observed data ( $\mu > 0$ )

Wide focus

Narrow focus

# Introducing a technical bias does not compromise coverage



Bias: FTS could accept some other value of  $\mu$  more frequently than the true value

Coverage: CIs have correct local coverage because we enforced it by construction

Any history of biased tests in particle physics?:  
The CLs method

Even for point estimation in high dimensions, one may prefer a low variance estimator over an unbiased estimator

FTS sets better lower bounds even with a focus centred at  $\mu = 0$

---

# FTS sets better lower bounds even with a focus centred at $\mu = 0$

Consistently larger lower bounds if signal exists, smaller upper bounds when it doesn't

Lower bounds when there is a signal:

$\mu^*$	LRS	FTS, wide	FTS, narrow
10.0	0.0 (0.0)	<b>2.08</b> (0.0)	1.60 (0.0)

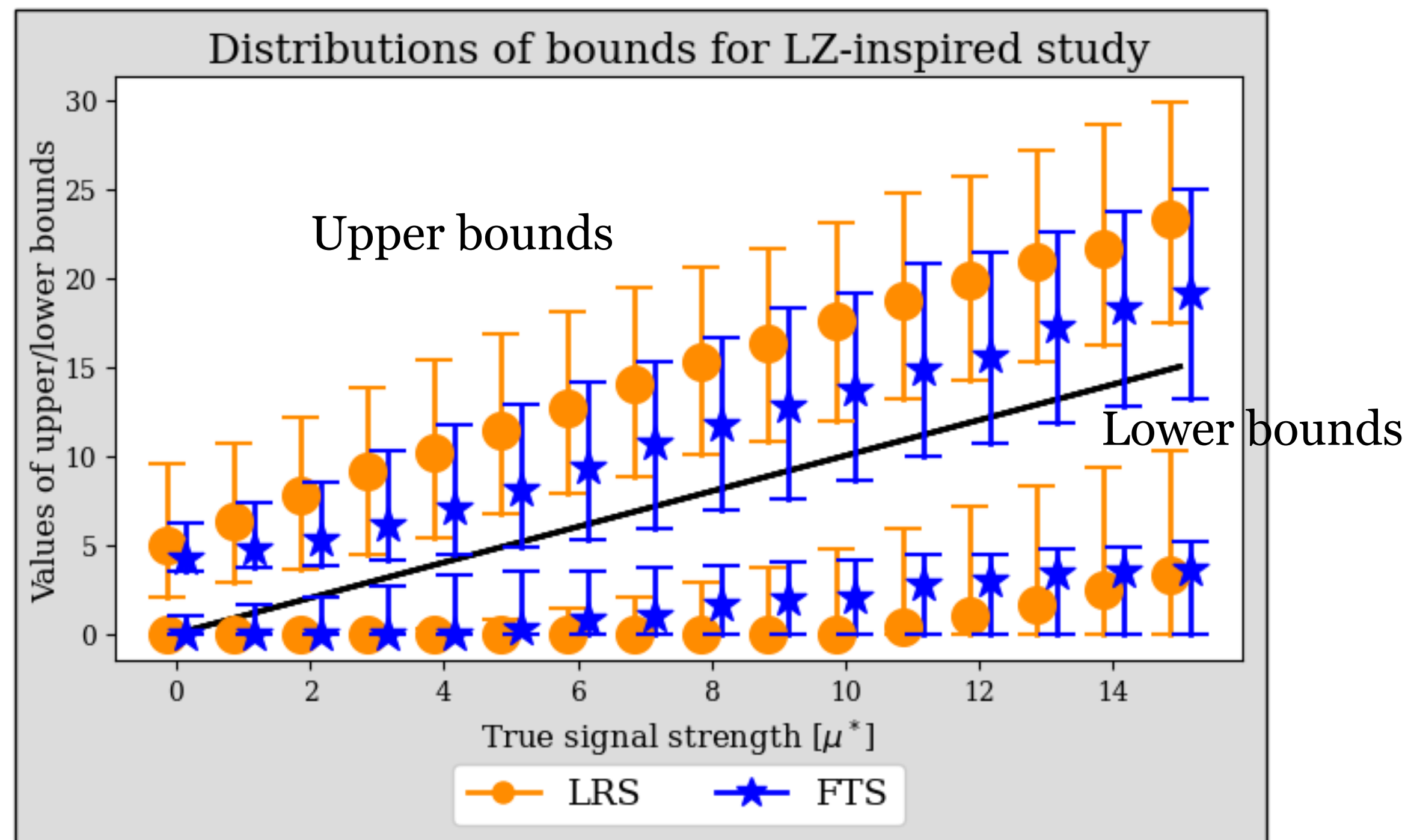
Upper bounds when there is no signal:

$\mu^*$	LRS	FTS, wide	FTS, narrow
0.0	4.97 (12.59)	4.25 ( <b>10.98</b> )	<b>2.73</b> (11.14)

Numbers represent length of confidence intervals for  $1\sigma$  ( $2\sigma$ )

# FTS sets better lower bounds even with a focus centred at $\mu = 0$

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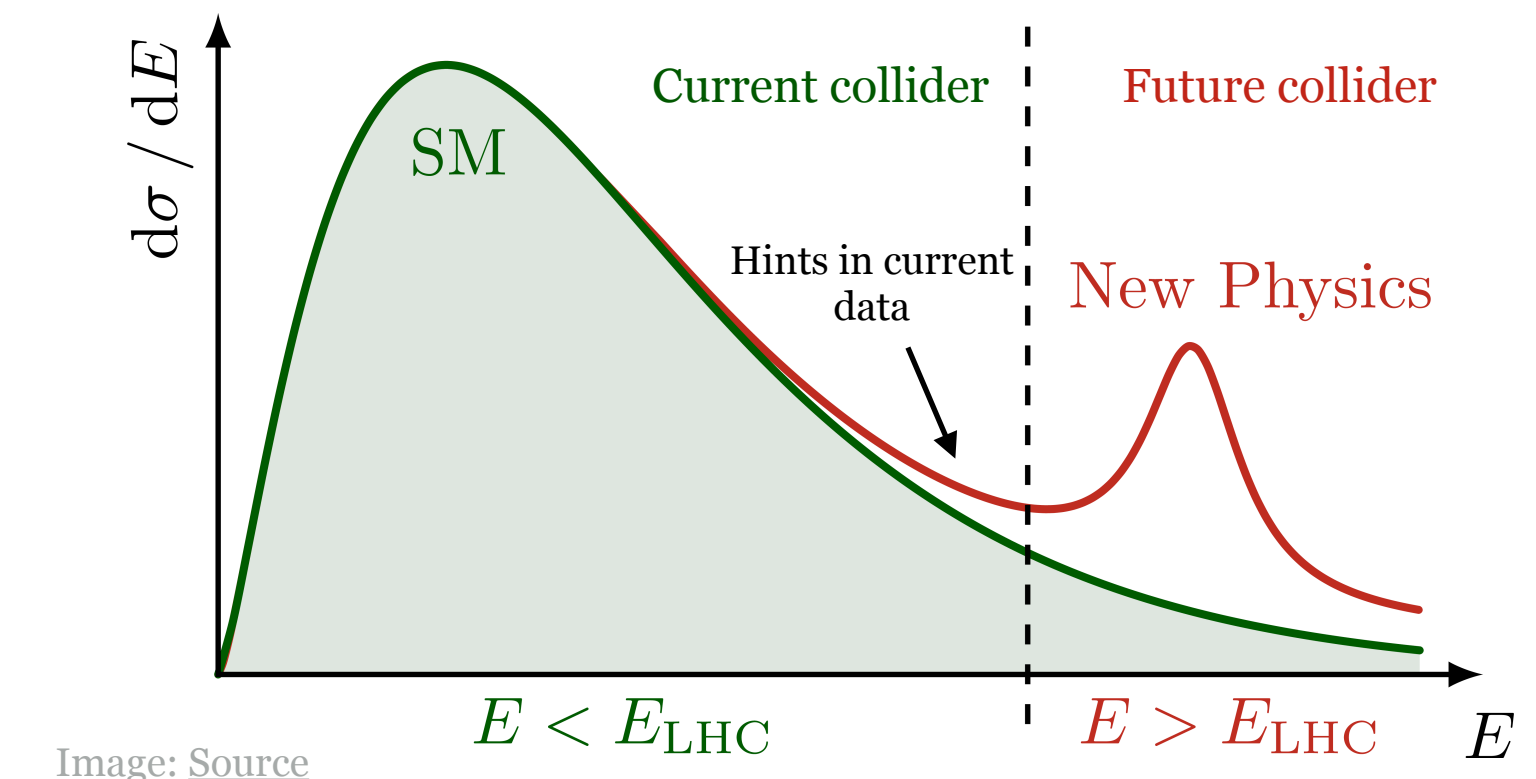
Numbers represent length of confidence intervals for  $1\sigma$  ( $2\sigma$ )

# Conclusion

- Semi-parametric NSBI let's us maximally extract information in high-dimensional unbinned data
- FTS + QR let's us robustly **outperform the likelihood ratio test**, by focusing power in physics-motivated regions
- NSBI+FTS gives us control over high-dimensional problems in particle physics

$$FTS(\mathcal{D}; \mu_0) = -2 \log \left( \frac{p(\mathcal{D} | \mu_0)}{\int_{\Theta} p(\mathcal{D} | \mu) f(\mu) d\mu} \right)$$

- 'Asymptopia' a fundamental misconception about our field, new physics is always just out of reach!



Thank you!

Backup

# Quantitative improvements with FTS

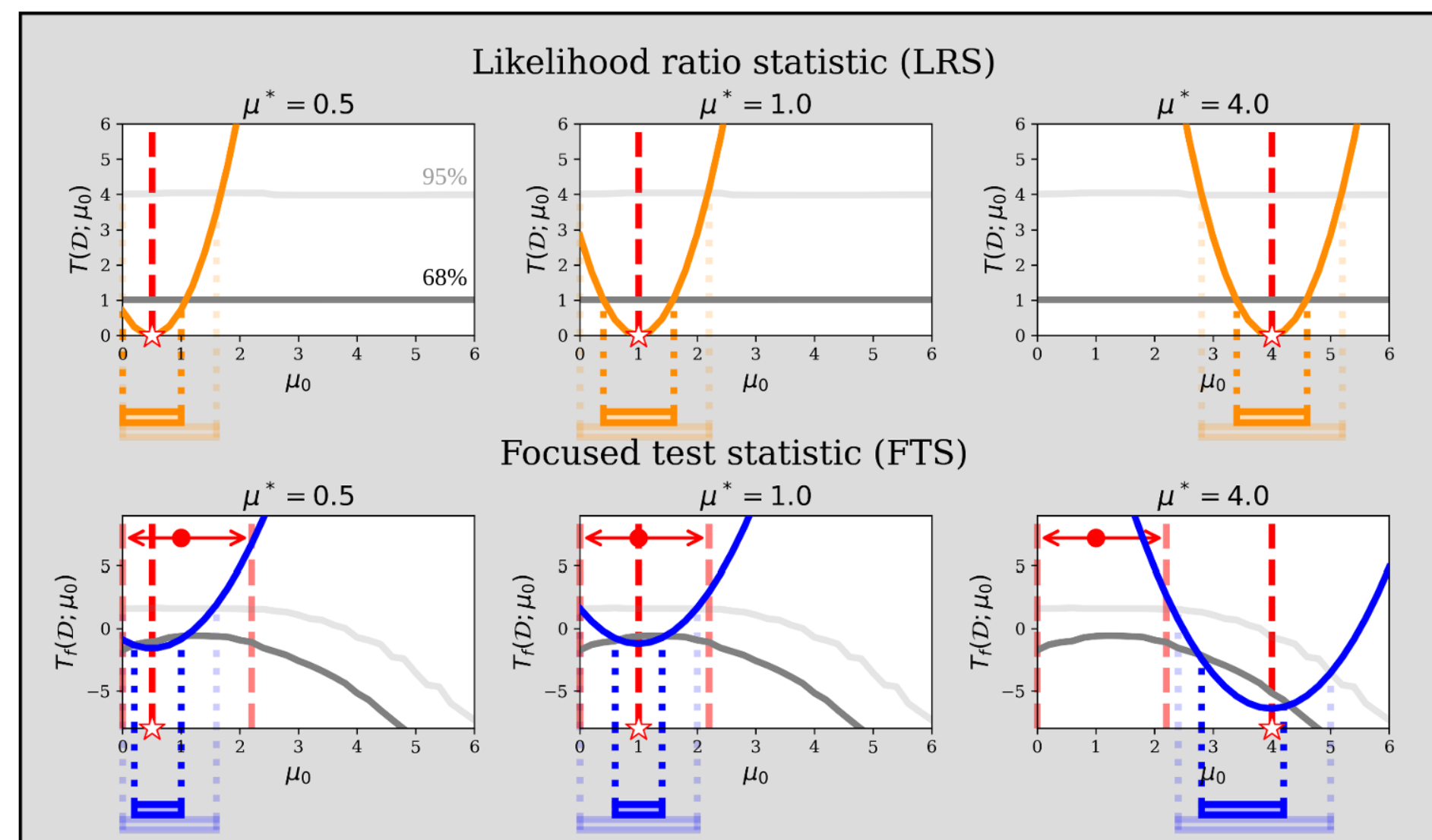
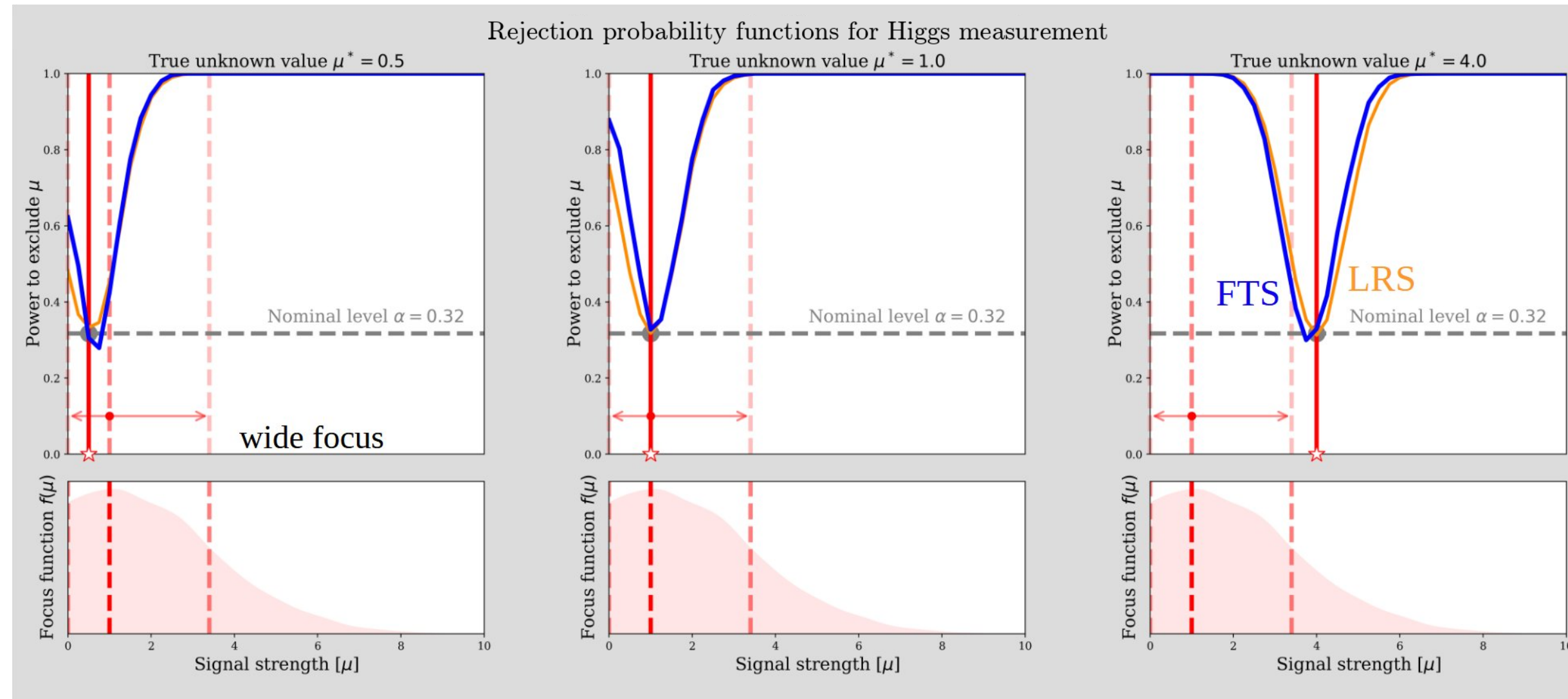
Using reconstructed  
Higgs mass as  
observable to  
construct histogram

Setting	$\mu^*$	Test statistic		
		LRS	FTS-wide	FTS-narrow
Higgs (mass)	1.0	1.08 (2.01)	0.94 (1.89)	<b>0.85 (1.79)</b>
Higgs (vis. mass)	1.0	1.26 (2.31)	1.10 (2.17)	<b>0.98 (2.08)</b>
LZ-inspired	0.0	5.99 (13.87)	4.68 (11.47)	<b>3.89 (11.29)</b>
LZ-inspired	1.0	7.08 (15.18)	5.29 (12.82)	<b>4.68 (12.75)</b>

Using visible energy  
as observable

Numbers represent length of confidence intervals for  $1\sigma$  ( $2\sigma$ )

# Introducing a technical bias does not compromise coverage



# Search-Oriented Mixture Model

---

$x_i$  vector representing one individual event      General Formula

$$p(x_i|\mu) = \frac{1}{\nu(\mu)} \sum_j^{\mathcal{C}} f_j(\mu) \cdot \nu_j p_j(x_i)$$

$j$  runs over different physics process  
(Eg.  $gg \rightarrow H^* \rightarrow 4l$ ,  $gg \rightarrow ZZ \rightarrow 4l$ )

---

Example use case

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$$p_{\text{ggF}}(x|\mu) = \frac{1}{\nu_{\text{ggF}}(\mu)} \left[ (\mu - \sqrt{\mu}) \nu_S p_S(x) + \sqrt{\mu} \nu_{\text{SBI}_1} p_{\text{SBI}_1}(x) + (1 - \sqrt{\mu}) \nu_B p_B(x) \right]$$

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Comes from theory model chosen to interpret data

Example use case

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$f_i(\mu)$  will depend on morphing bases points (which values of  $\mu$  were used to simulate samples)

# Search-Oriented Mixture Model

$x_i$  vector representing one individual event

General Formula

$$p(x_i|\mu) = \frac{1}{v(\mu)} \sum_j^C f_j(\mu) \cdot v_j p_j(x_i)$$

Event rates estimated from simulations

Comes from theory model chosen to interpret data

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$$\frac{p(x_i|\mu)}{p_{\text{ref}}(x_i)} = \frac{1}{v(\mu)} \sum_j^C f_j(\mu) \cdot v_j \frac{p_j(x_i)}{p_{\text{ref}}(x_i)}$$

Event rates estimated from simulations
Comes from theory model chosen to interpret data

Reference hypothesis  $j$  runs over different physics process  
 (Eg.  $gg \rightarrow H^* \rightarrow 4l, gg \rightarrow ZZ \rightarrow 4l$ )

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Estimated using an ensemble of networks

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Reference hypothesis

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Example use case

$$p_{\text{ggF}}(x|\mu) = \frac{1}{\nu_{\text{ggF}}(\mu)} \left[ (\mu - \sqrt{\mu}) \nu_S p_S(x) + \sqrt{\mu} \nu_{\text{SBI}_1} p_{\text{SBI}_1}(x) + (1 - \sqrt{\mu}) \nu_B p_B(x) \right]$$

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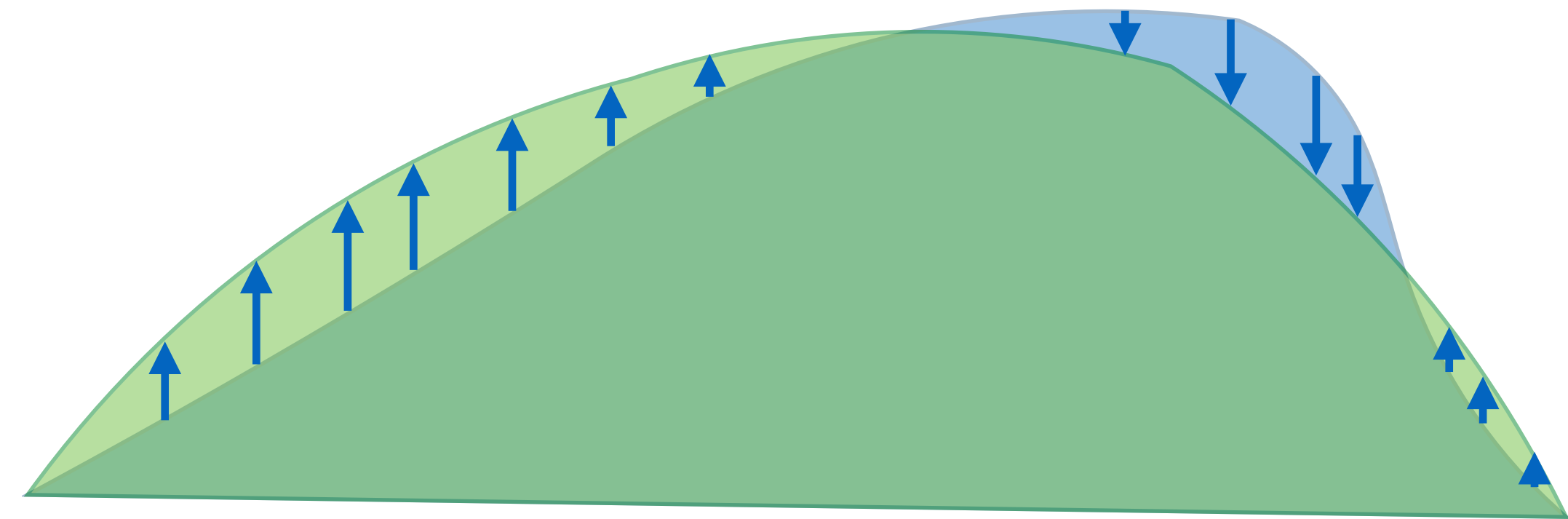
$f_i(\mu)$  will depend on morphing bases points (which values of  $\mu$  were used to simulate samples)

More NSBI diagnostics

# Validate quality of LR estimation with re-weighting task

---


Reweighting: Calculate weights  $w_i$  for events  $x_i$  in **blue sample** to match **green sample**

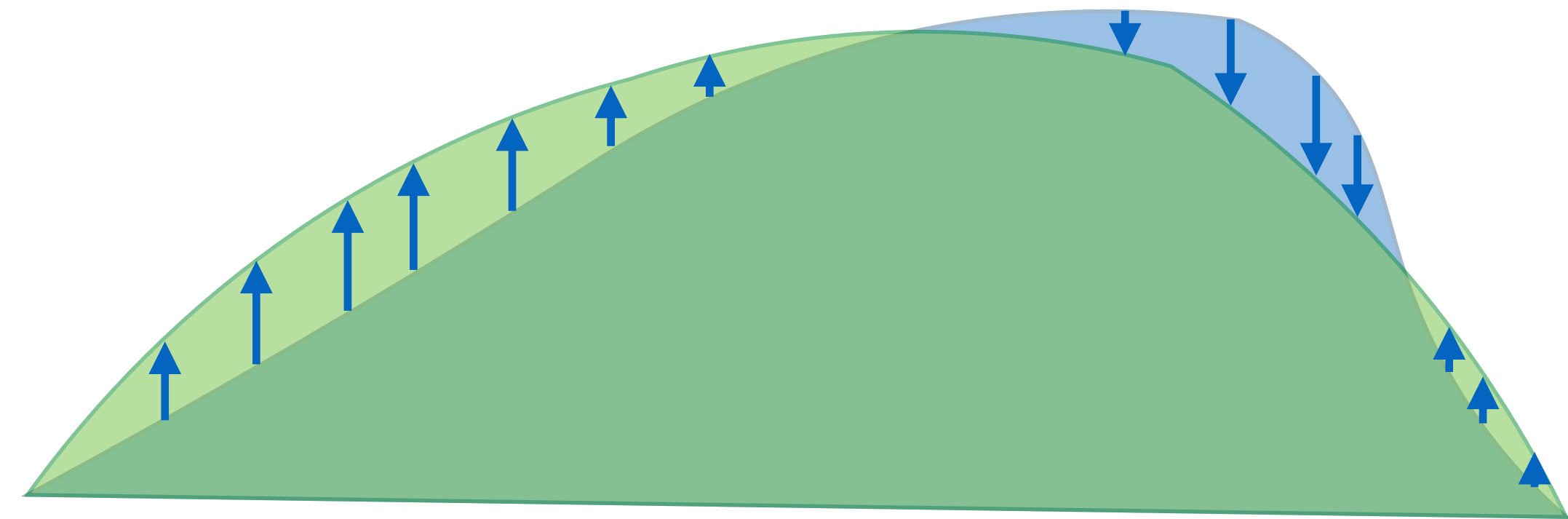


# Validate quality of LR estimation with re-weighting task

---

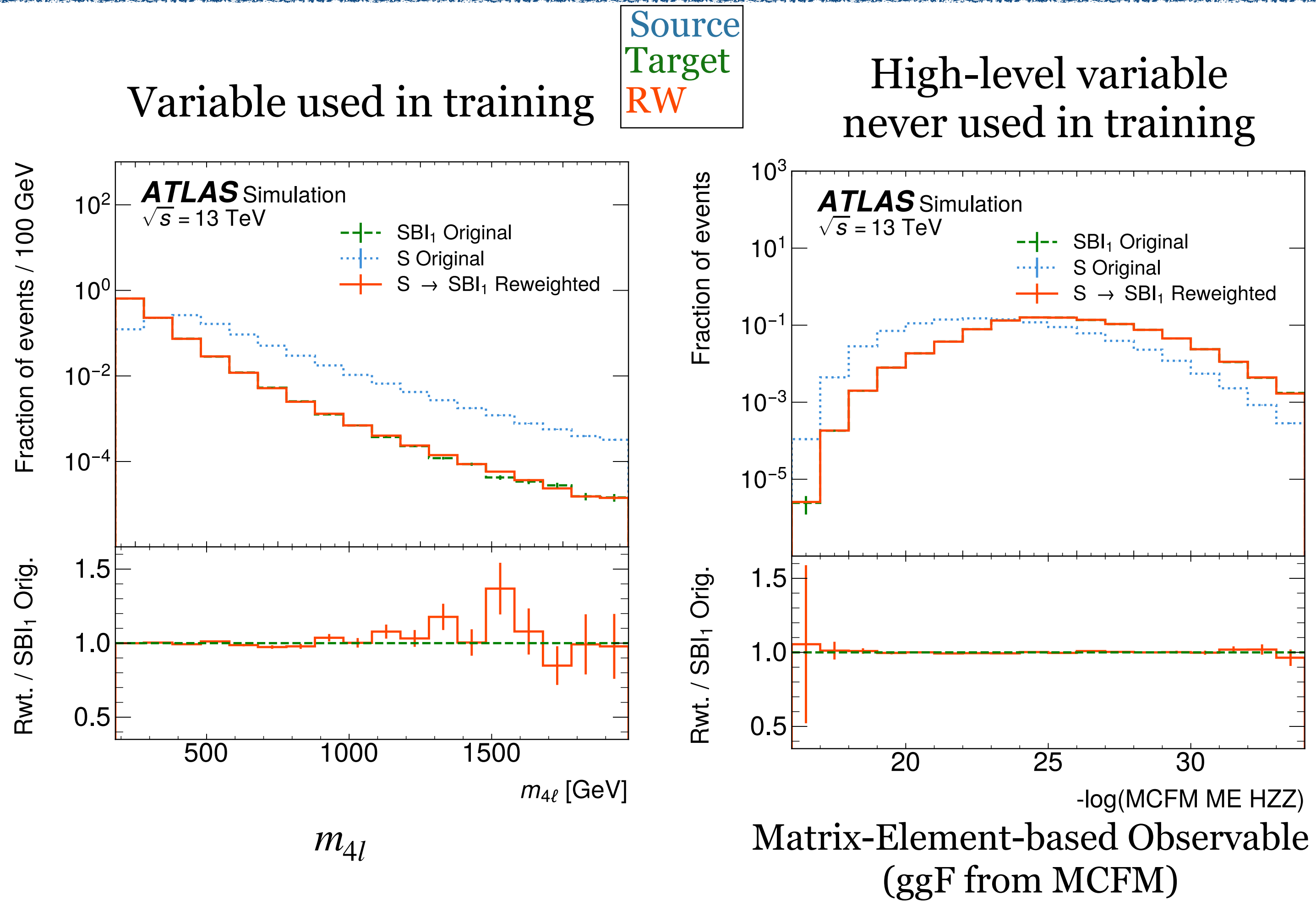
Reweighting: Calculate weights  $w_i$  for events  $x_i$  in **blue sample** to match **green sample**

$$w_i = r(x_i, \mu_0, \mu_1) = \frac{p(x_i | \mu_0)}{p(x_i | \mu_1)}$$




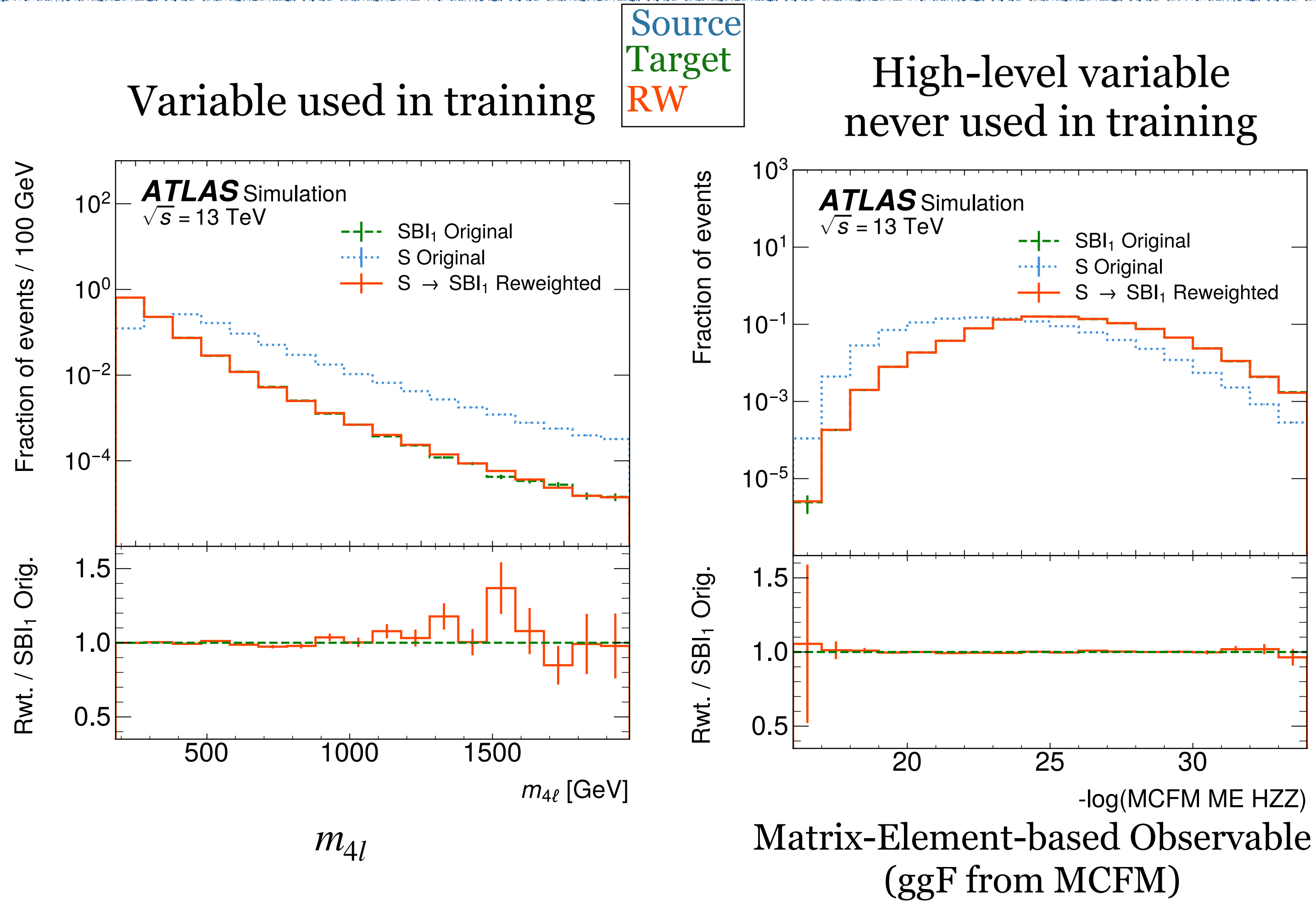
Already estimated using an ensemble of networks

# Re-weight closures



Hoping to see Target and RW match

# Re-weight closures



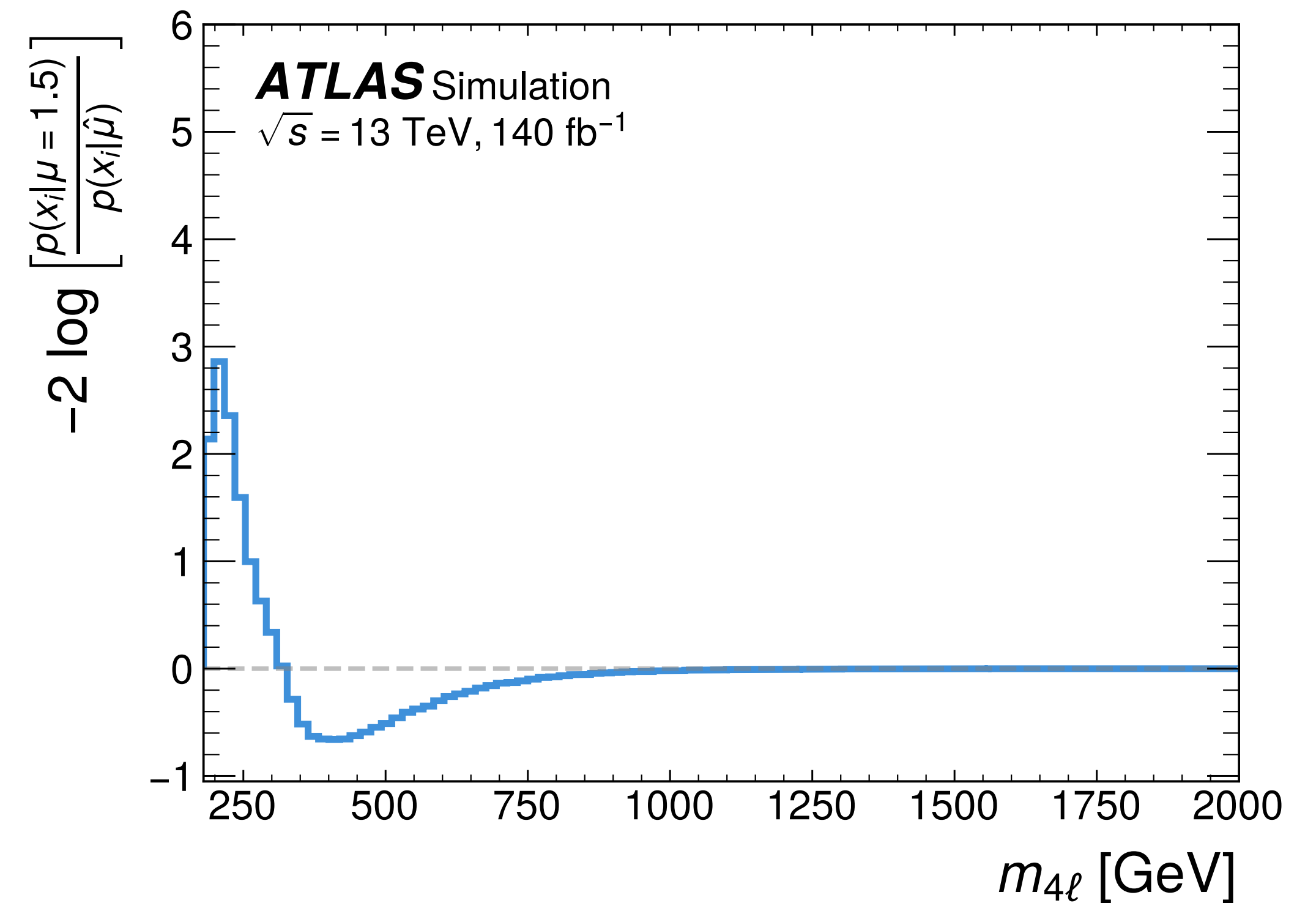
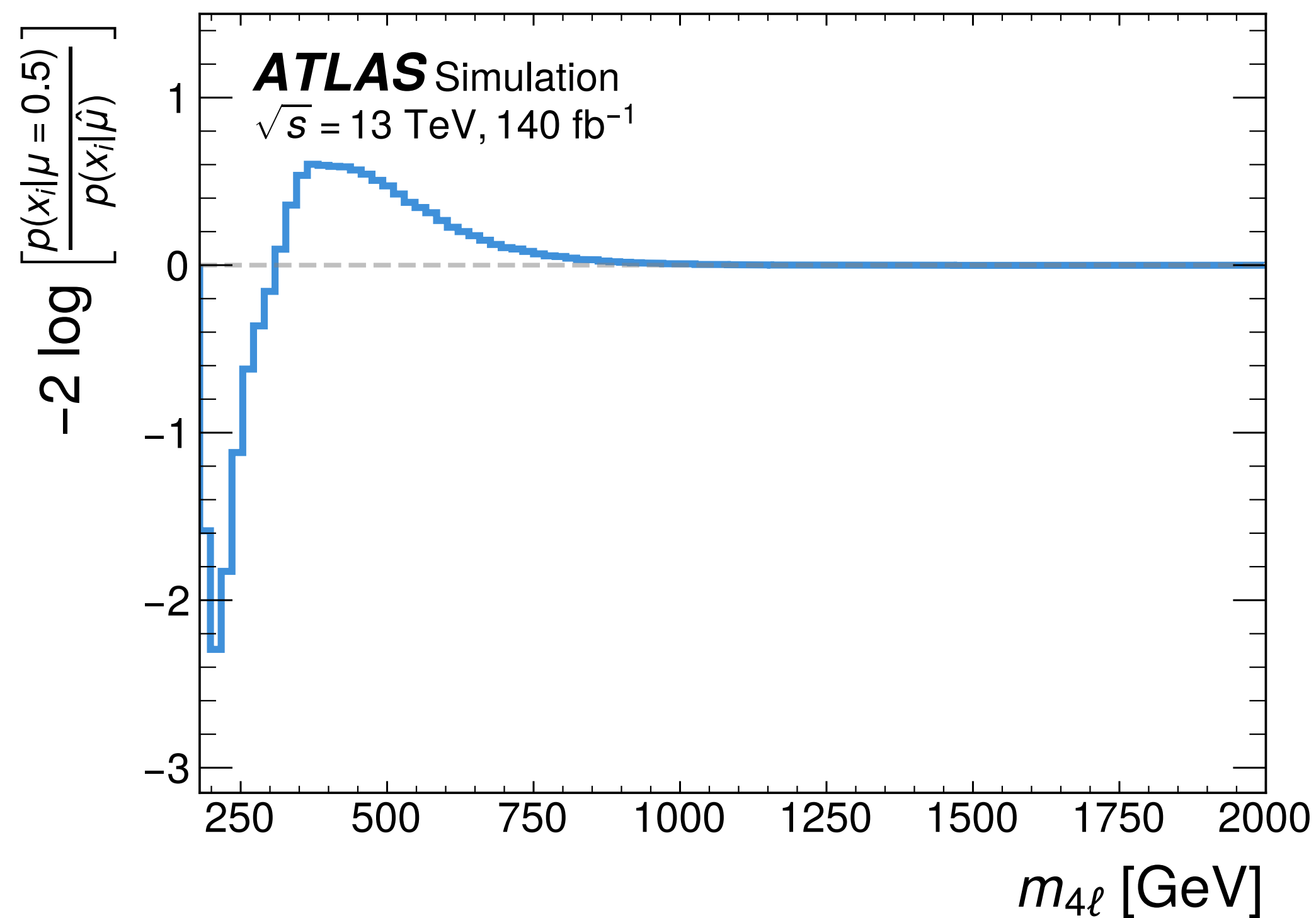
High-dim goodness-of-fit test:  
Train independent classifier on **RW** vs **Target**  
AUC=0.5  $\Rightarrow$  LRs well estimated

Hoping to see **Target** and **RW** match

# Interpretability: Which phase space favours one hypothesis over another?

$$-2 \cdot \log \frac{P(x_i | \mu = 0.5)}{P(x_i | \mu = 1)}$$

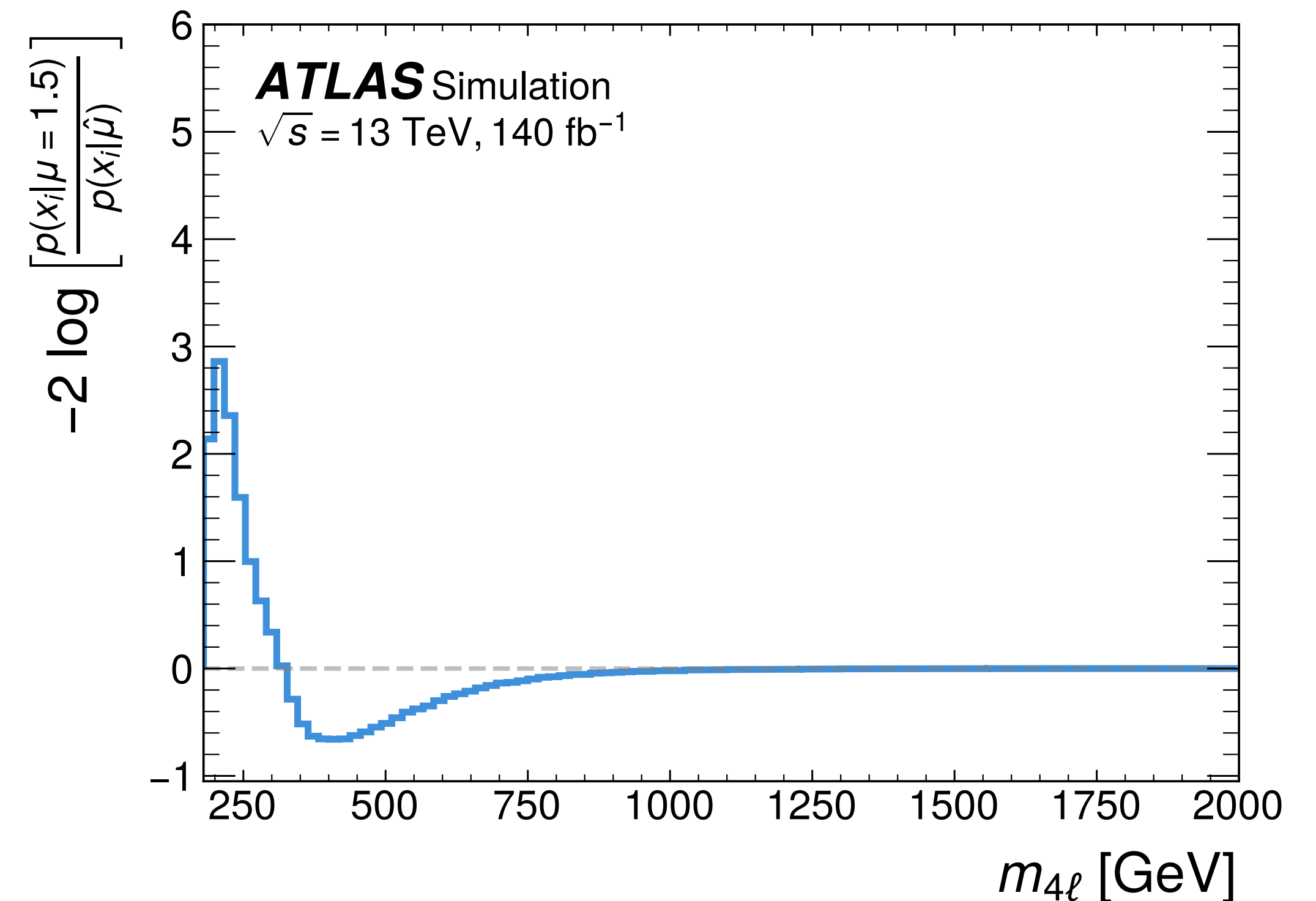
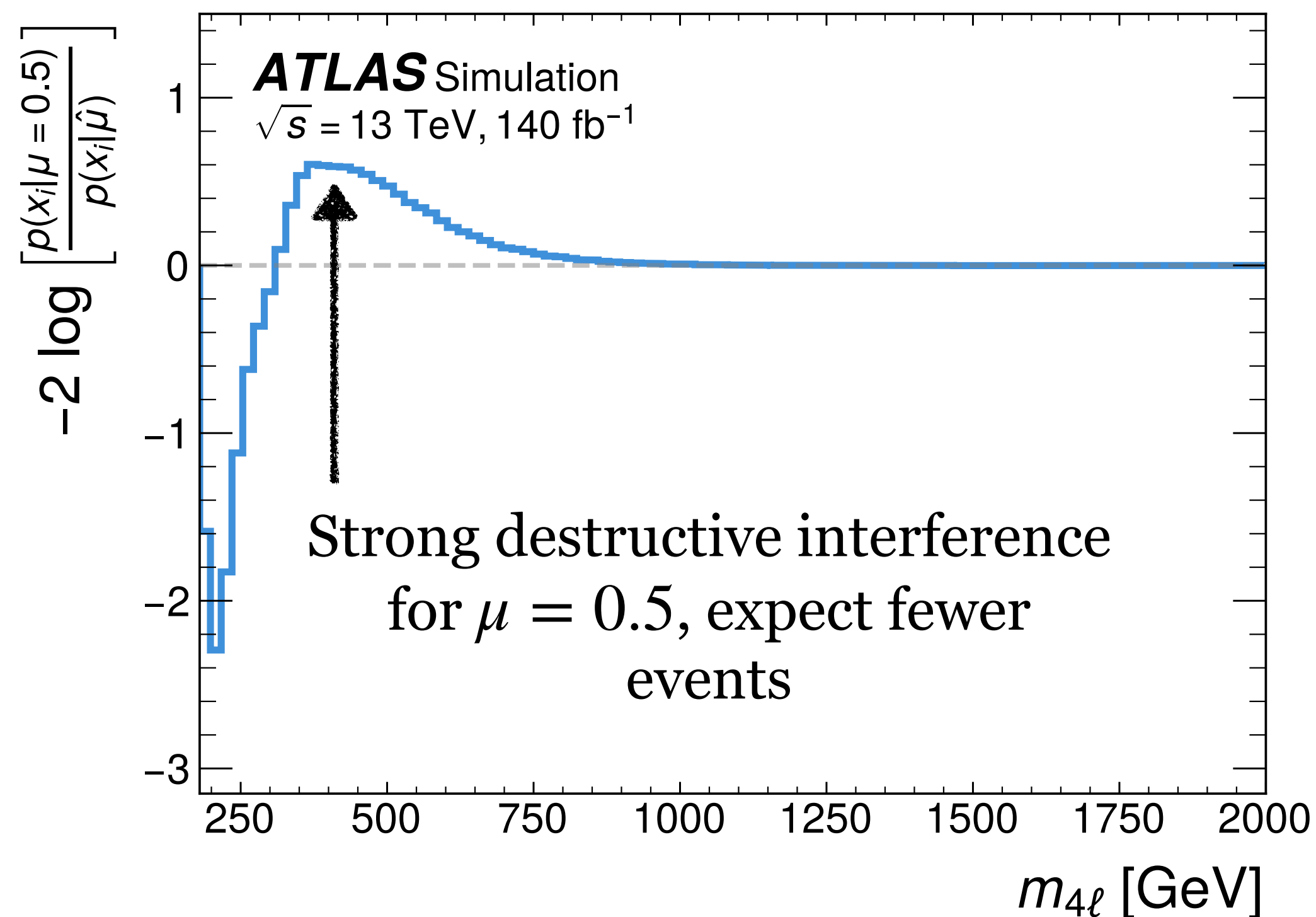
$$-2 \cdot \log \frac{P(x_i | \mu = 1.5)}{P(x_i | \mu = 1)}$$



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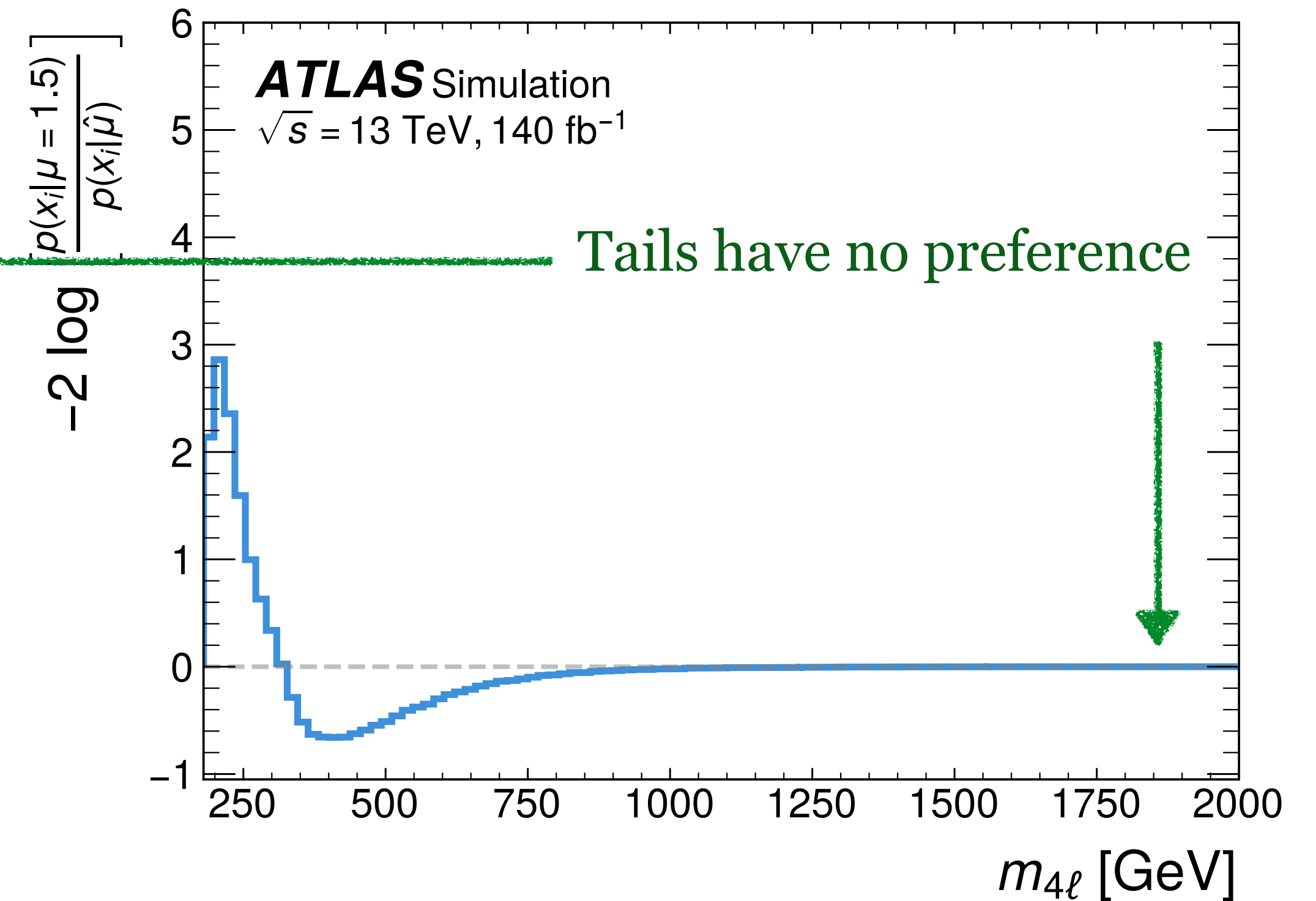
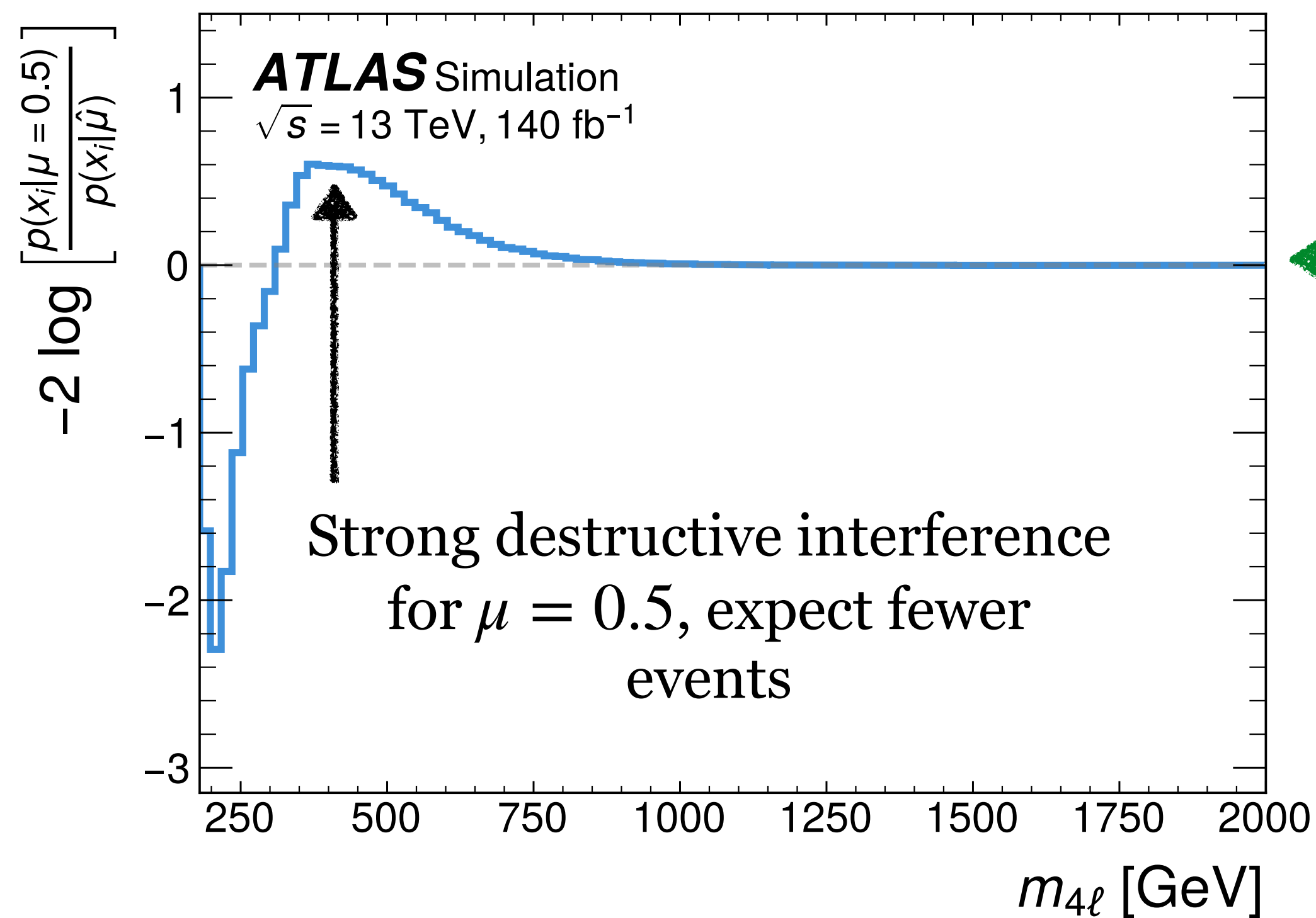
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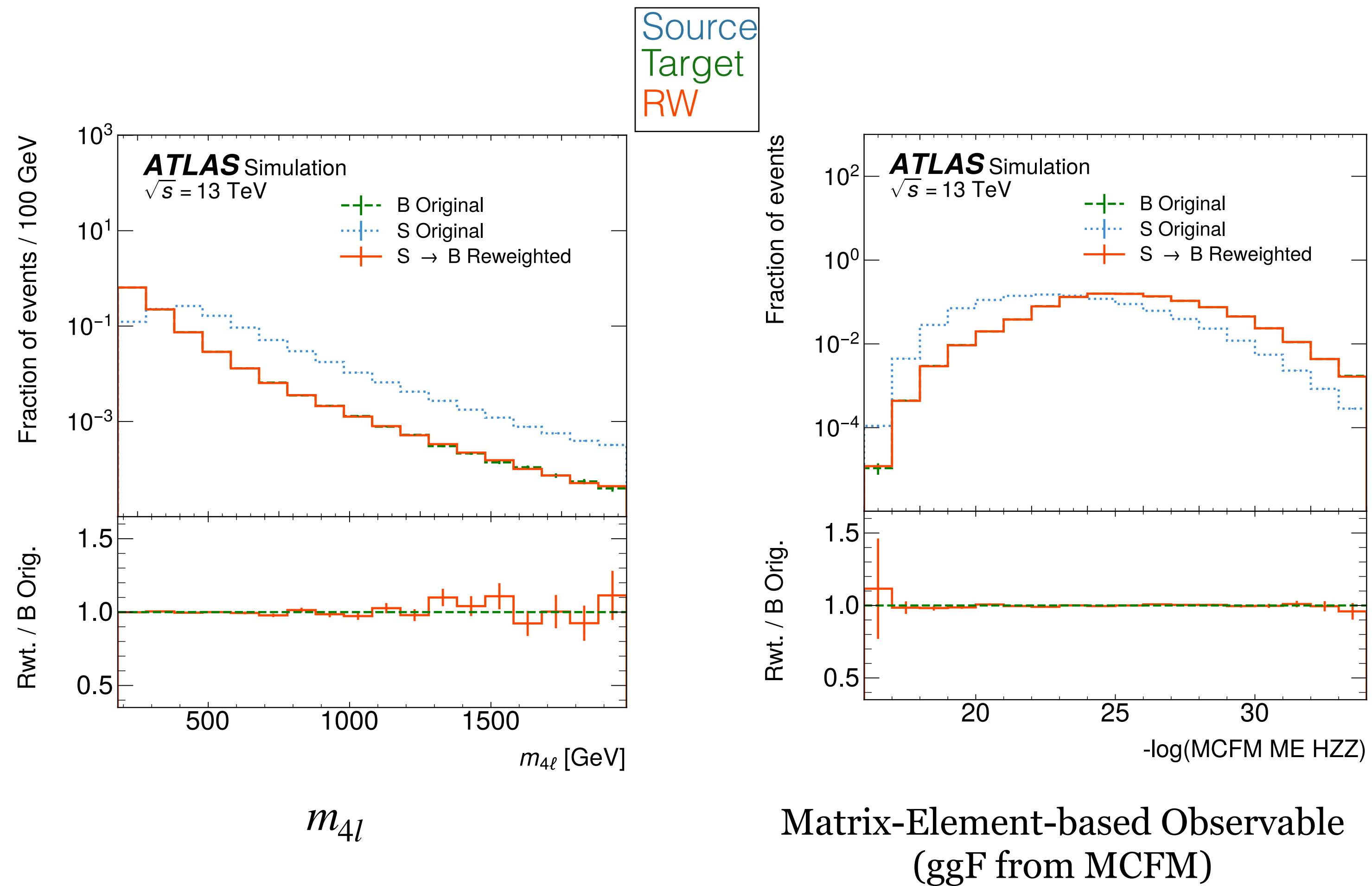
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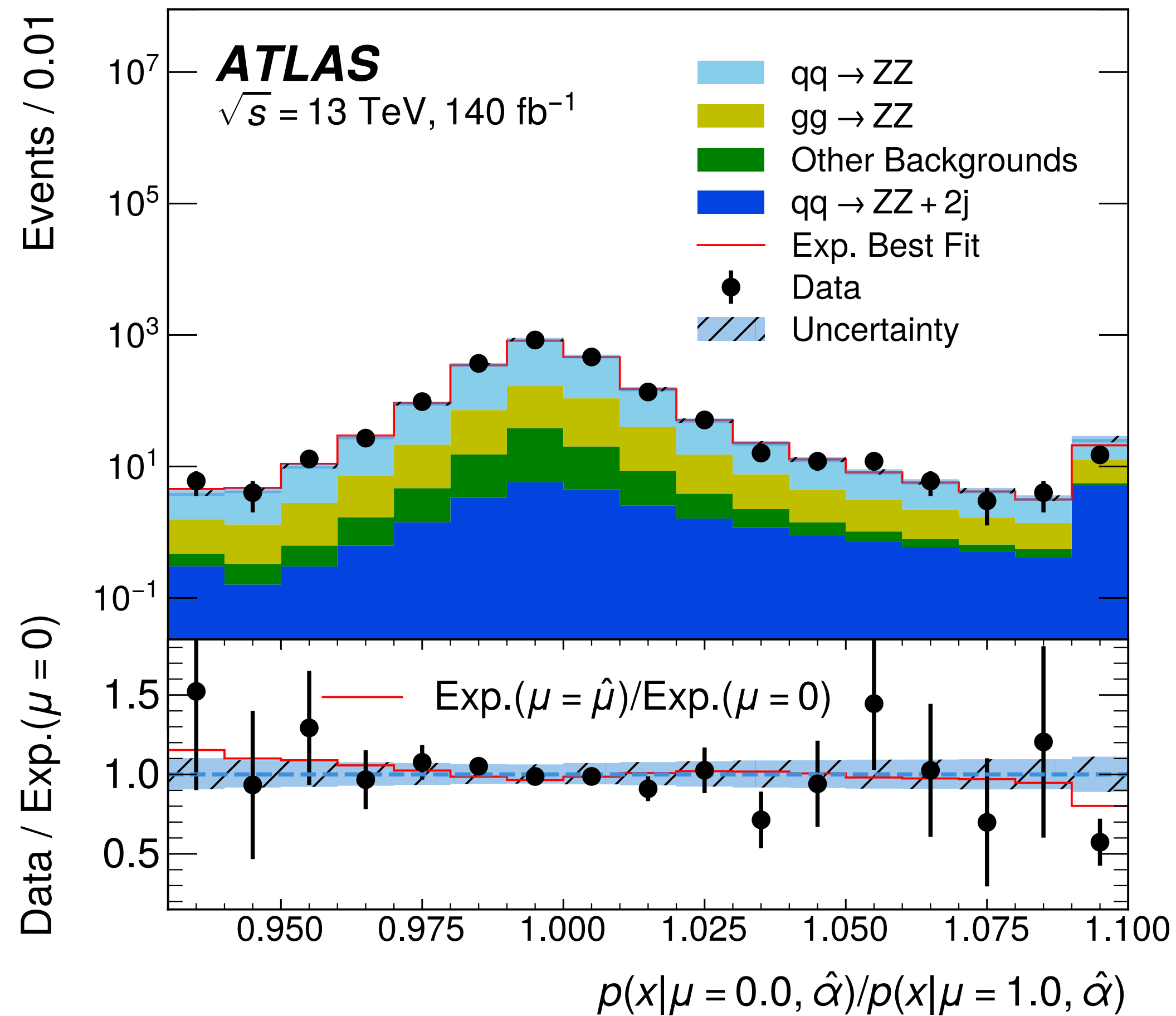
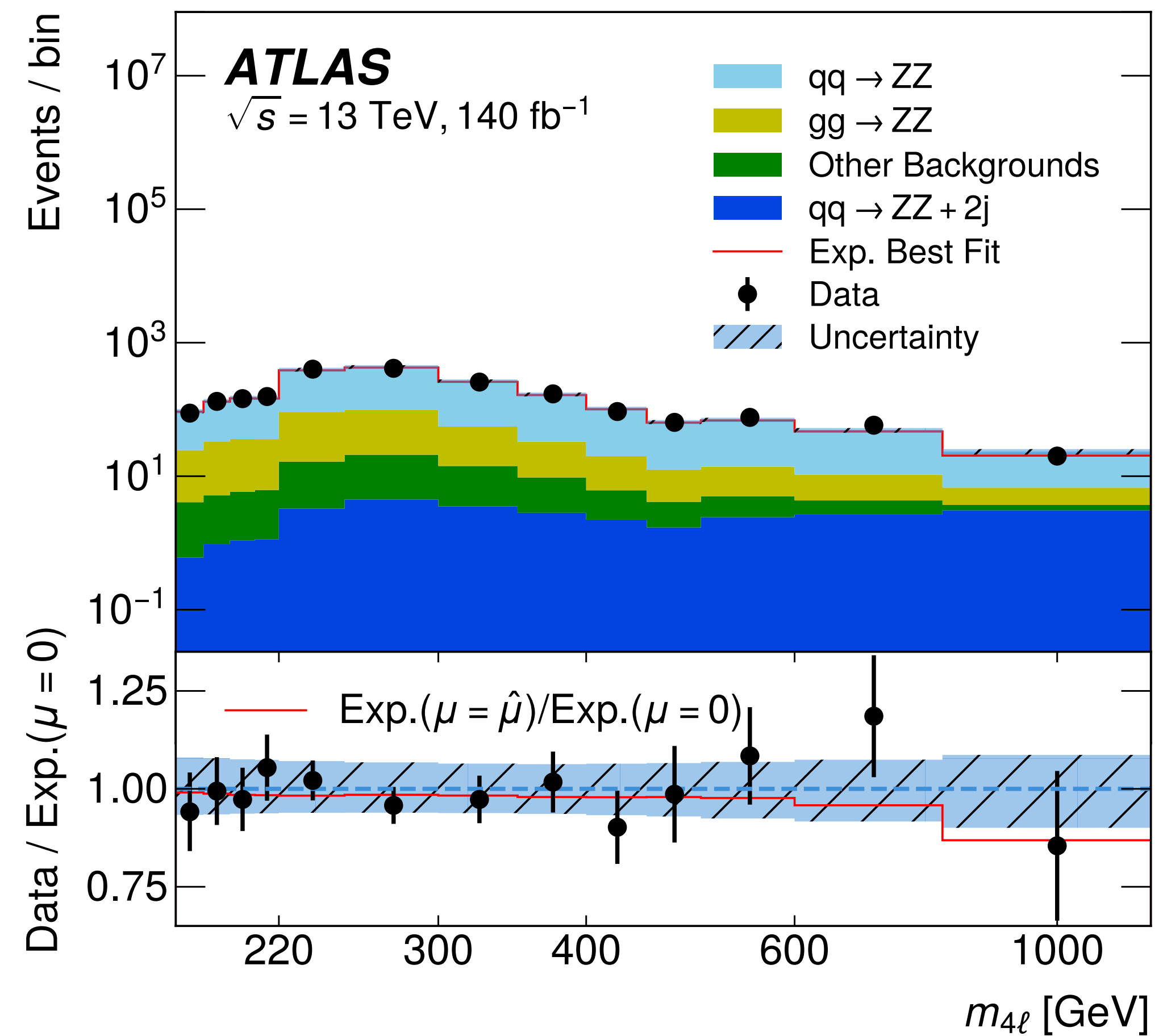


# Re-weight closures for B



## Data-MC validation

NN observable

 $m_{4\ell}$ 

## Data-MC validation

## Different NN observables

